



HPCI project field 5
“The origin of matter and the universe”

2014/06/05 ARIS2014
at Ito Hall, Tokyo

Frontier of Nuclear Shell- Model Calculations and High- Performance Computing



Noritaka SHIMIZU
CNS, University of Tokyo

T. Abe (Tokyo), M. Honma (Aizu), T. Mizusaki (Senshu),
T. Otsuka (Tokyo), T. Togashi (CNS),
Y. Tsunoda (Tokyo), and T. Yoshida (CNS)

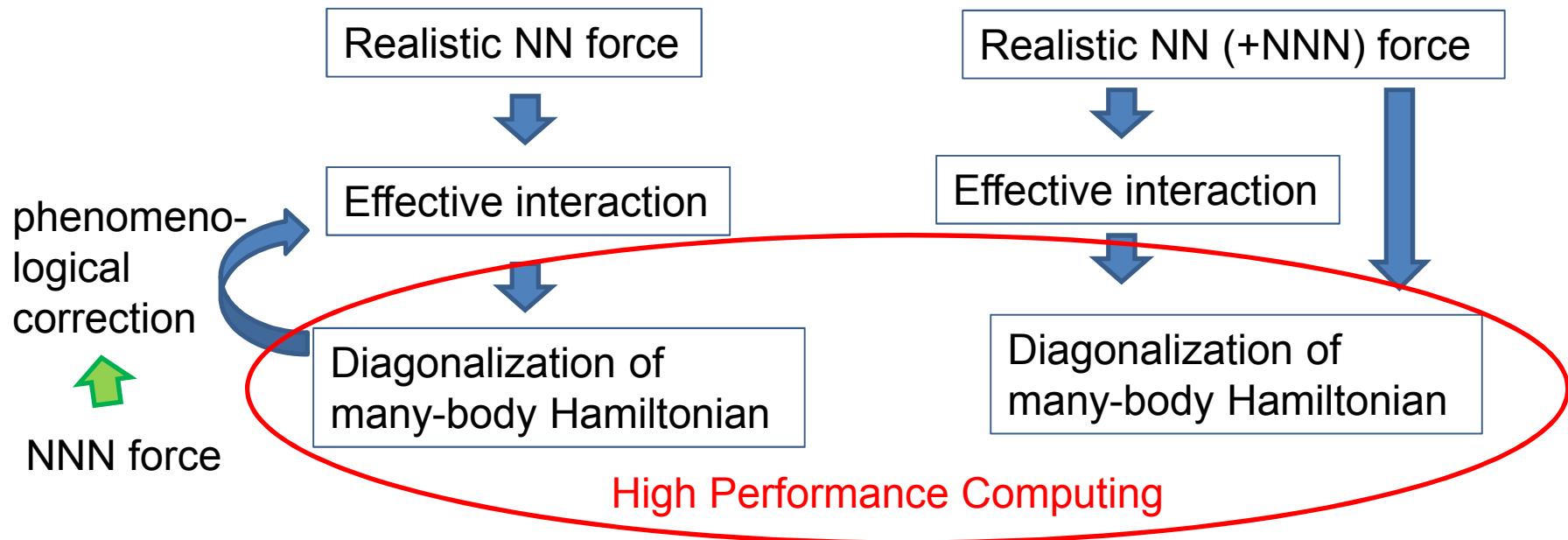
Recipe for shell-model calculations

large-scale shell-model calc.

model space: valence shell (+ extention)

no-core shell model

excitation energy (Nmax hw)



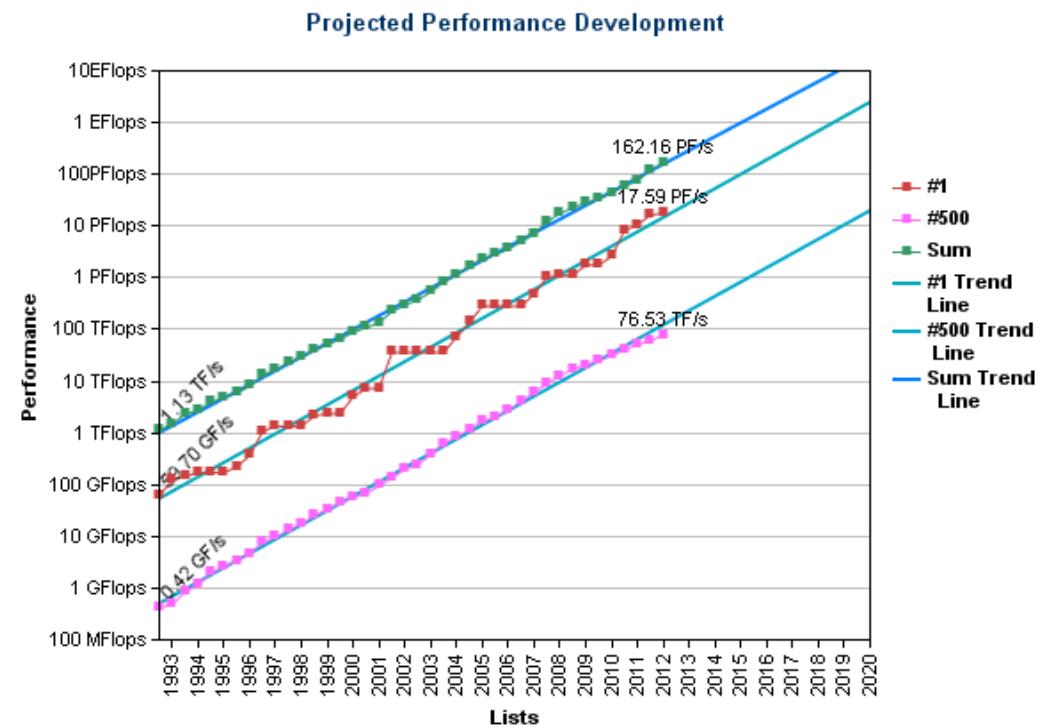
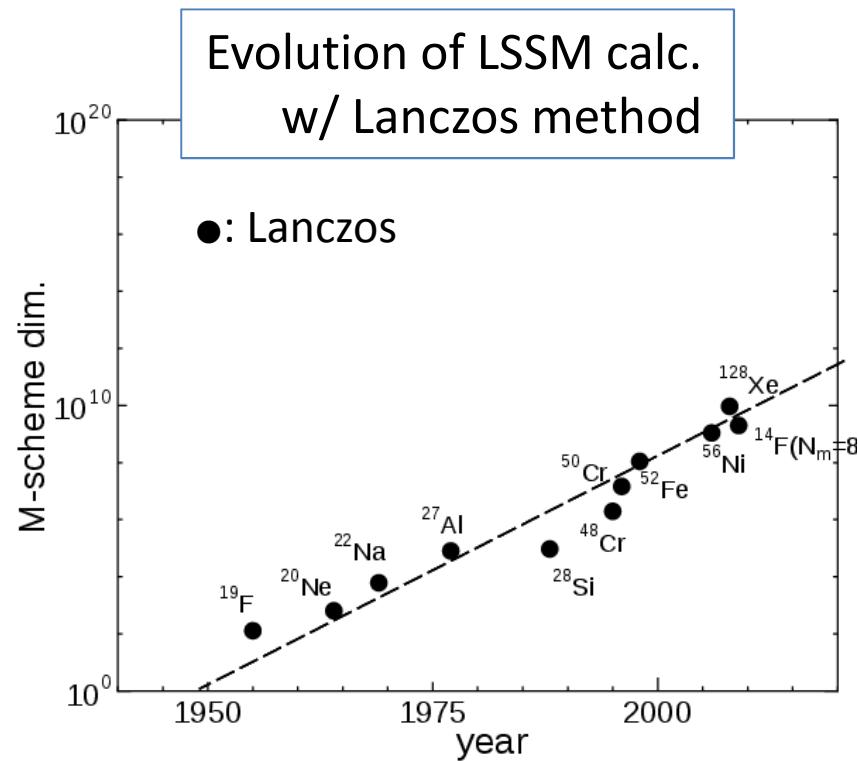
$$|\Psi\rangle = \sum_m v_m |m\rangle \quad \sum_{m'} \langle m | H | m' \rangle v_{m'} = E |m\rangle$$

$|m\rangle$: configuration, “spherical” Slater determinant of harmonic oscillator

$$\text{The number of } |m\rangle \approx \frac{1}{f} \binom{N_p}{Z} \binom{N_p}{Z}$$

Hamiltonian matrix is sparse
⇒ Lanczos algorithm

Frontier of large-scale shell-model calculations



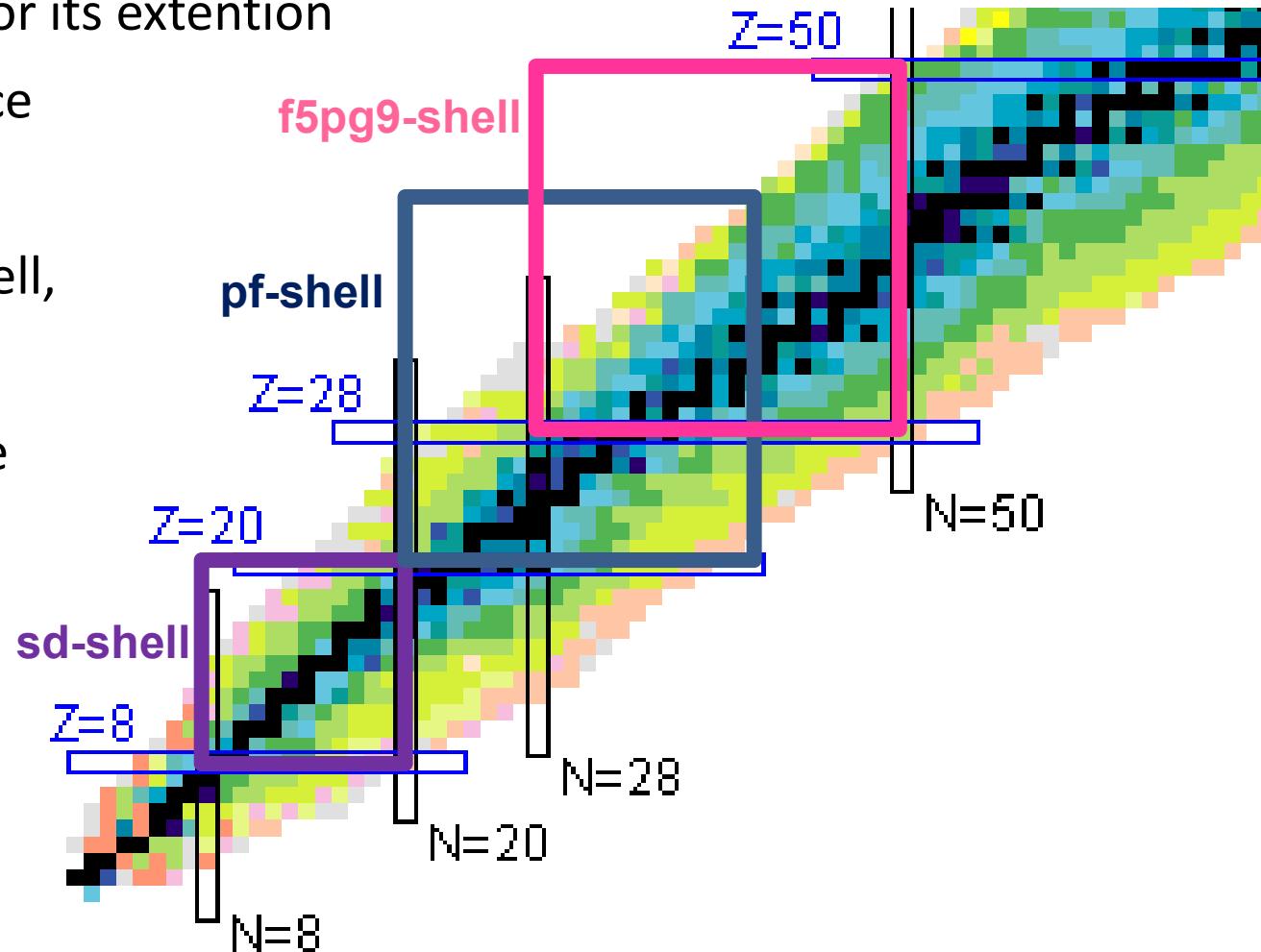
- The maximum dimension of the M -scheme Hamiltonian matrix of the LSSM evolve exponentially as the performance development of supercomputers
- Present limitation $\approx 10^{10-11}$ (as expected by the extrapolation in 2000)
e.g. ^{68}Ni in pf+g9/2+d5/2 5×10^{15} → 2040 for 10^{15} dim.?

Large-scale shell model calculations

valence shell or its extention

as model space

sd shell, *pf* shell,
f5pg9 shell
diagonalizable



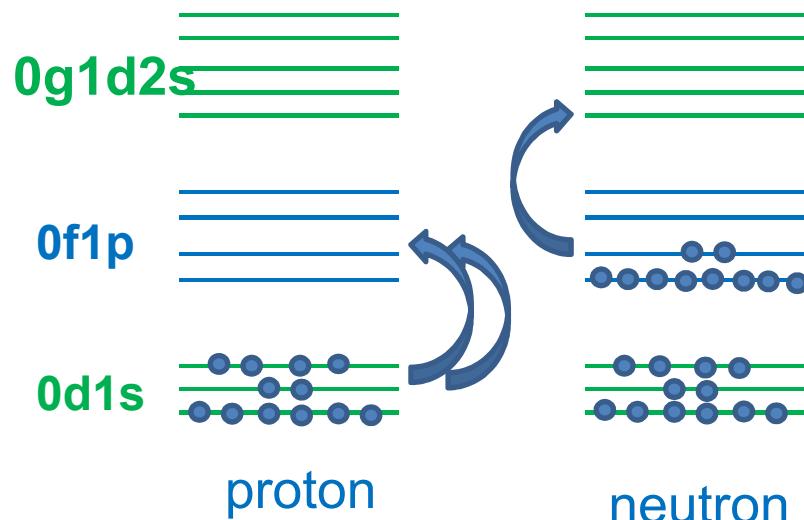
^{60}Zn in pf-shell : 2×10^9

^{78}Y in f5pg9-shell : 1.3×10^{10}

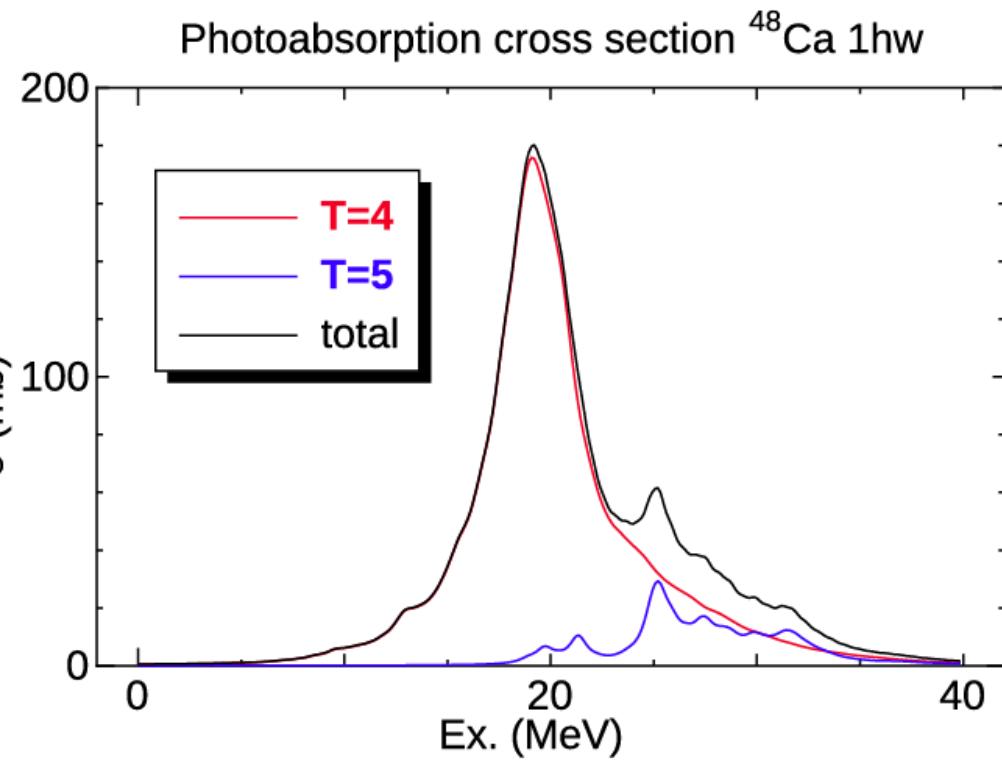
Example of the large-scale shell-model calculations

NS, Y. Utsuno, S. Ebata, T. Otsuka, M. Honma and T. Mizusaki, in preparation

1hw/3hw sd-pf-sdg shell calculations
for negative-parity states of Ca isotopes
to describe E1 excitations



We develop M -scheme shell-model code with Lanczos method for massive parallel computation "KSHELL".



1hw : upto 1hw excitation in sd-pt-sdg shell

4.1×10^6 M-scheme dim. at PC

(1+3)hw: up to 3hw excitation in sd-pf-sdg shell

1.2×10^{10} M-scheme dim.
at supercomputer



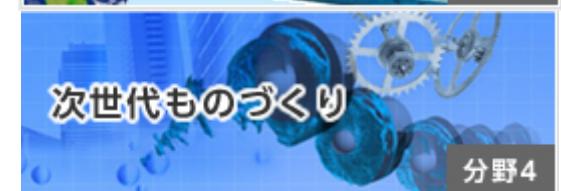
HPCI Strategic Program for Innovative Research

by MEXT, Japan FY2011 – FY2015

to facilitate efficient use of K computer and HPC

- Strategic 5 Fields:

- Field1: “Supercomputational Life Science”
- Field2: “Computational Materials Science Initiative”
- Field 3: “Projection of Planet Earth Variations for Mitigating Natural Disasters”
- Field 4: “Industrial Innovation”
- Field 5: “The origin of matter and the universe”
 - lattice QCD
 - Nucleus ← Our group
 - Supernova Explosion Head: Takaharu Otsuka (Tokyo)
 - Early Star Formation



nuclear shell-model calculations using Monte Carlo shell model

Monte Carlo Shell Model

Otsuka, Honma, Mizusaki, Shimizu,
Utsuno, PPNP47, 319 (2001)

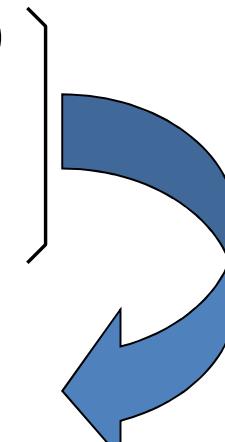
a tool to go beyond the conventional Lanczos diagonalization method
for Large-Scale Shell Model calculations

$$\mathbf{H} = \begin{pmatrix} * & * & * & * & * & \dots \\ * & * & * & * & \cdot & \dots \\ * & * & * & \cdot & \cdot & \cdot \\ * & * & \cdot & \cdot & \cdot & \cdot \\ \cdot & \cdot & \cdot & \cdot & \cdot & \cdot \\ \cdot & \cdot & \cdot & \cdot & \cdot & \cdot \end{pmatrix}$$

Conventional Shell Model
all Slater determinants

diagonalization

$$\begin{pmatrix} \varepsilon_1 & & & & & \\ \varepsilon_2 & & & & & \\ \varepsilon_3 & & & & & \\ & \ddots & & & & \\ & & 0 & & & \\ & & & \ddots & & \\ & & & & \ddots & \end{pmatrix}$$



Stochastically selected basis

$$\mathbf{H} \approx \begin{pmatrix} * & * & * & \cdot \\ * & * & * & \cdot \\ * & * & \cdot & \cdot \\ \cdot & \cdot & \cdot & \cdot \end{pmatrix}$$

Monte Carlo Shell Model
bases important for a specific eigenstate

diagonalization

$$\begin{pmatrix} \varepsilon'_1 & \varepsilon'_2 & 0 \\ & \cdot & \cdot \\ 0 & & \cdot \end{pmatrix}$$

$\hat{h}(\sigma)$ one-body Hamiltonian

σ ... auxiliary field

random numbers

$$|\phi_k\rangle = \prod_{\alpha=1}^N \left(\sum_{i=1}^{N_{sp}} c_i^\dagger D_{i\alpha}^{(k)} \right) |-\rangle$$

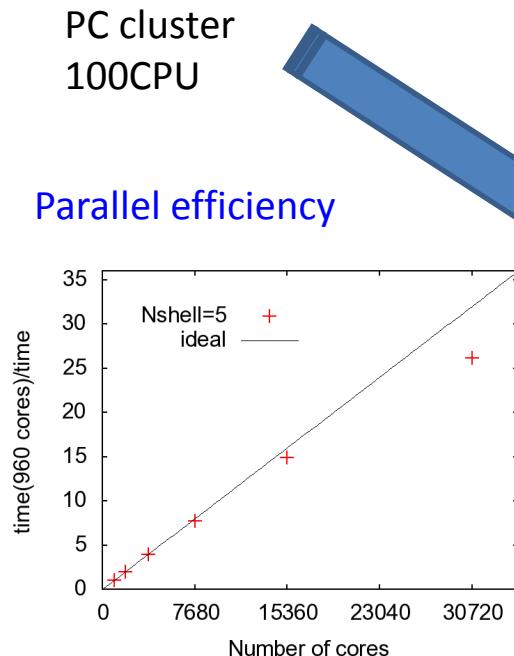
$$|\Psi\rangle = \sum_{k=1}^{N_{MCSM}} f_k P^{J,\pi} |\phi_k\rangle$$

↑ MCSM basis, deformed Slater det.

Wave function is described as a linear combination of J -projected deformed Slater determinants.
Increase MCSM dimension, or number of deformed Slater det. till energy converges.

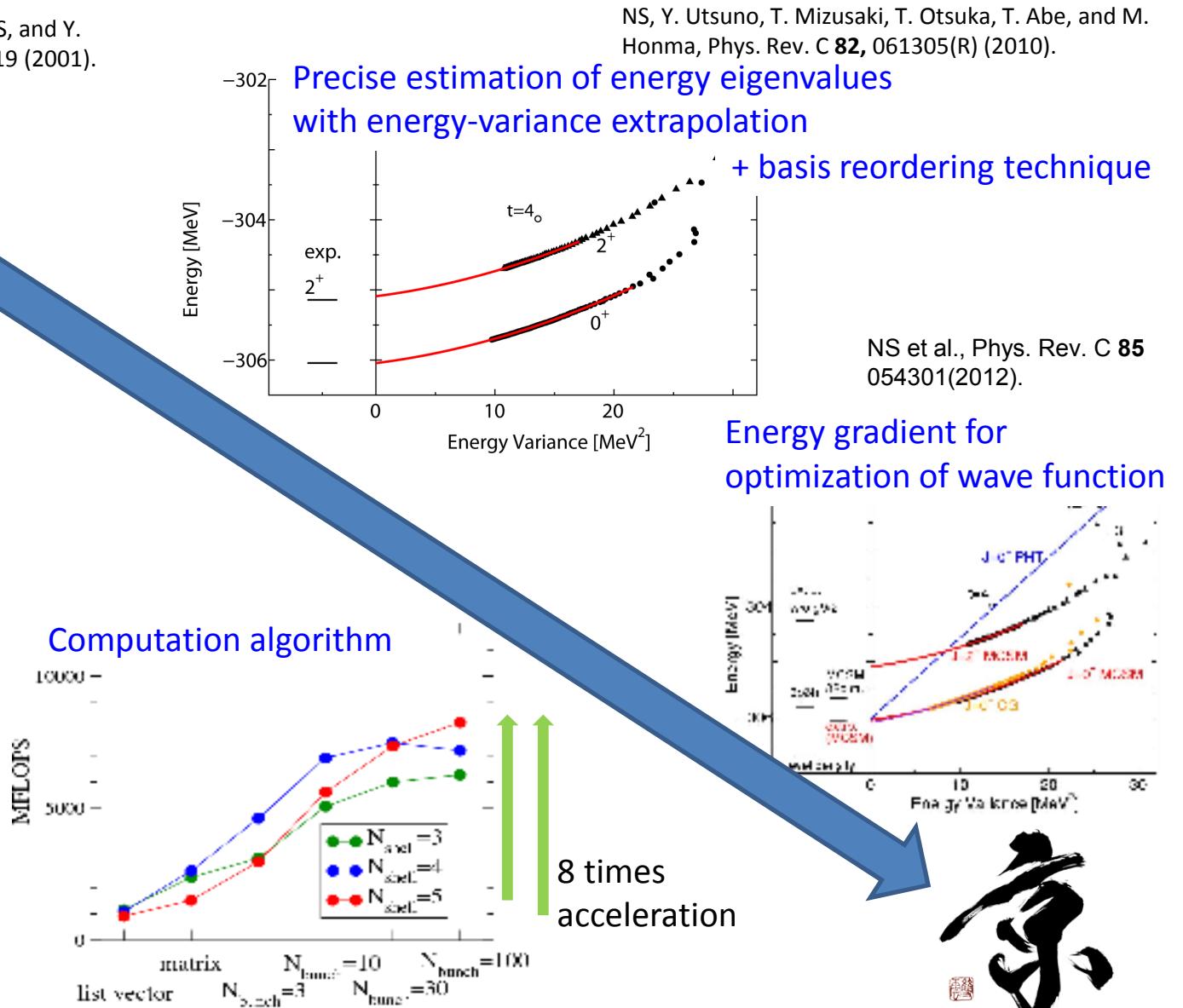
“Advanced” Monte Carlo Shell Model towards K computer

T. Otsuka, M. Honma, T. Mizusaki, NS, and Y. Utsuno, Prog. Part. Nucl. Phys. **47**, 319 (2001).



OpenMP+MPI hybrid parallel
100,000 cores

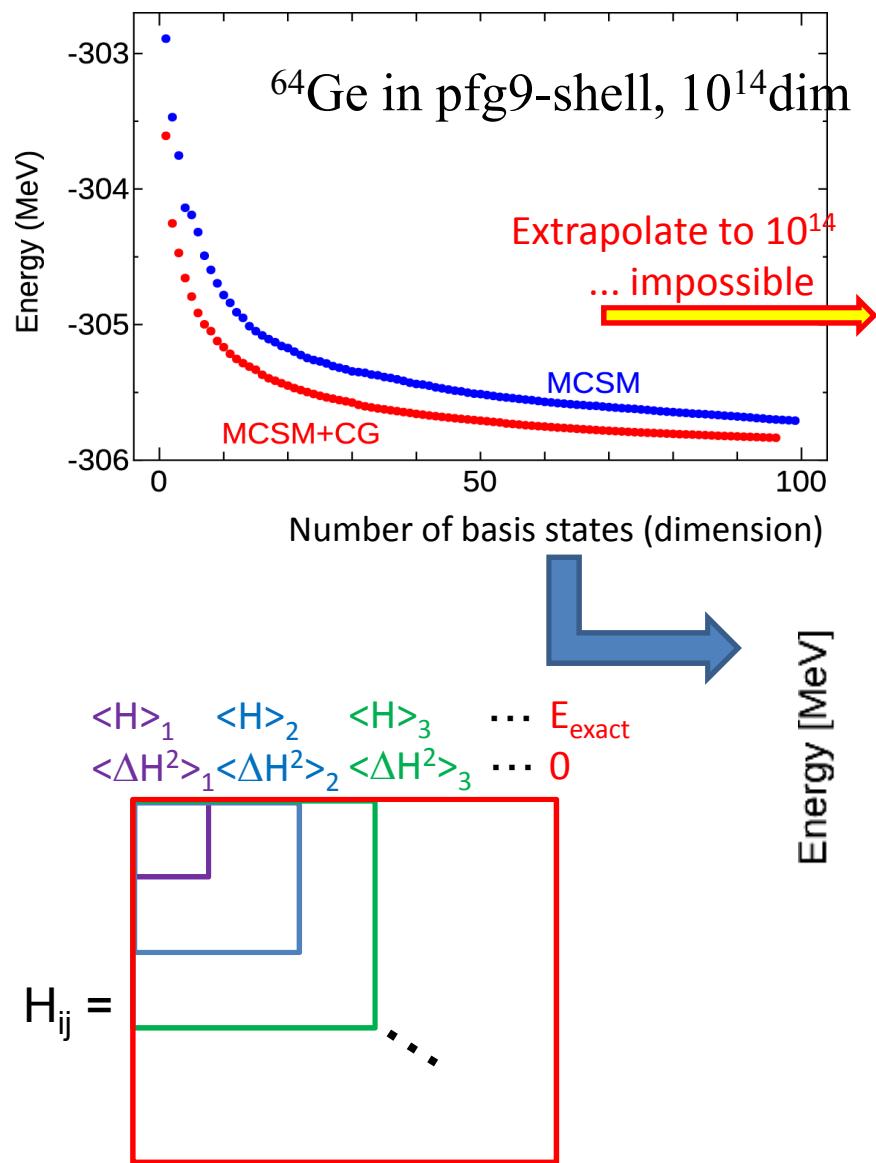
NS, T. Abe, Y. Tsunoda, Y. Utsuno, T. Yoshida, T. Mizusaki, M. Honma and T. Otsuka, Prog. Theor. Exp. Phys. **2012** 01A205 (2012).



Y. Utsuno, NS, T. Otsuka, and T. Abe, Comp. Phys. Comm. **184** 102 (2013).



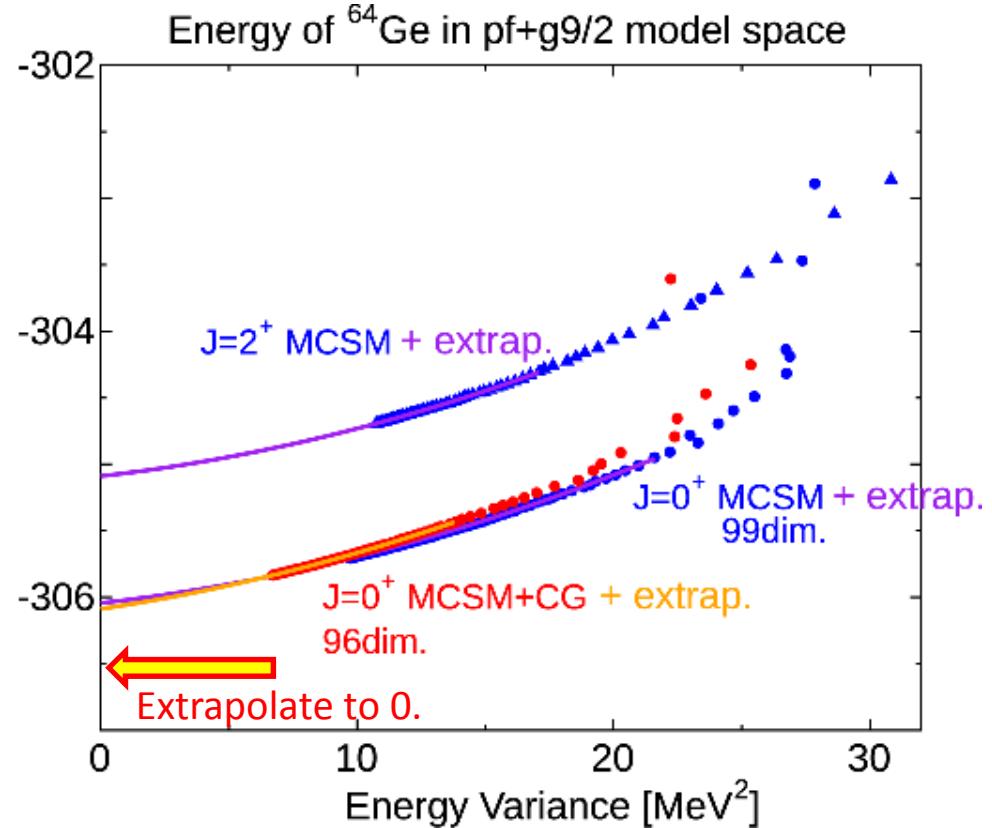
Energy variance extrapolation in the MCSM: ^{64}Ge in $pfg9$ -shell



$$\text{Energy variance: } \langle \Delta H^2 \rangle = \langle H^2 \rangle - \langle H \rangle^2$$

As the number of basis states increases, the approximated w.f. approaches the exact one and the energy variance approaches zero.

Extrapolate towards $\langle \Delta H^2 \rangle \rightarrow 0$



Advantages of the MCSM

- computation intensive, rather small memory access, small I/O ... good at massive parallel computation

M-scheme Lanczos method

$$|\Psi\rangle = \sum_{m=1}^{\sim 10^{10}} f_k |\phi_m\rangle$$

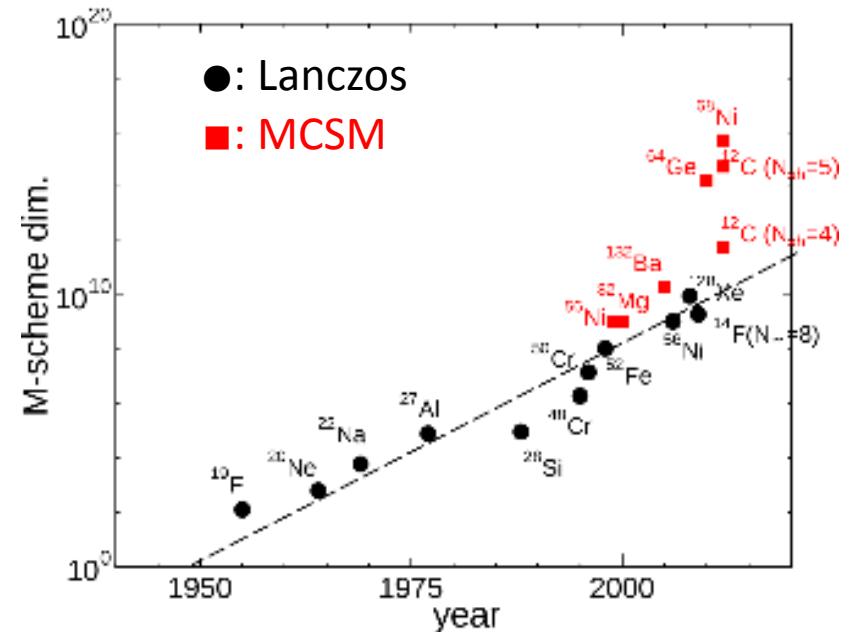
huge memory and disk I/O

vs MCSM wave function

$$|\Psi\rangle = \sum_{k=1}^{N_{MCSM} \sim 100} f_k P^{J,\pi} |\phi_k\rangle$$

discretized integral
of Euler's angle

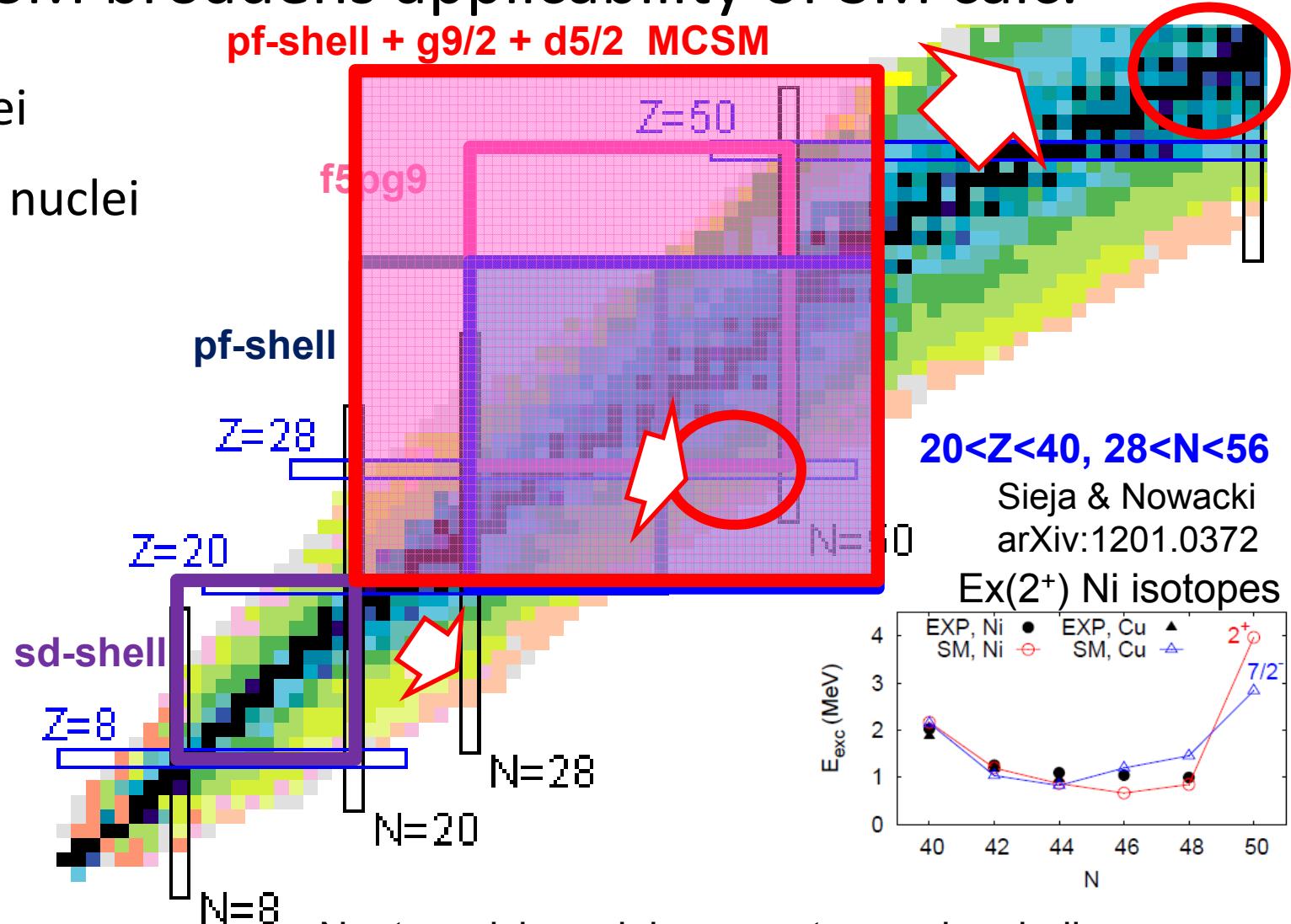
- overcome the limitation of Lanczos method
- analyze “intrinsic” wave function and shape (discussed later)



MCSM broadens applicability of SM calc.

heavier nuclei

neutron-rich nuclei



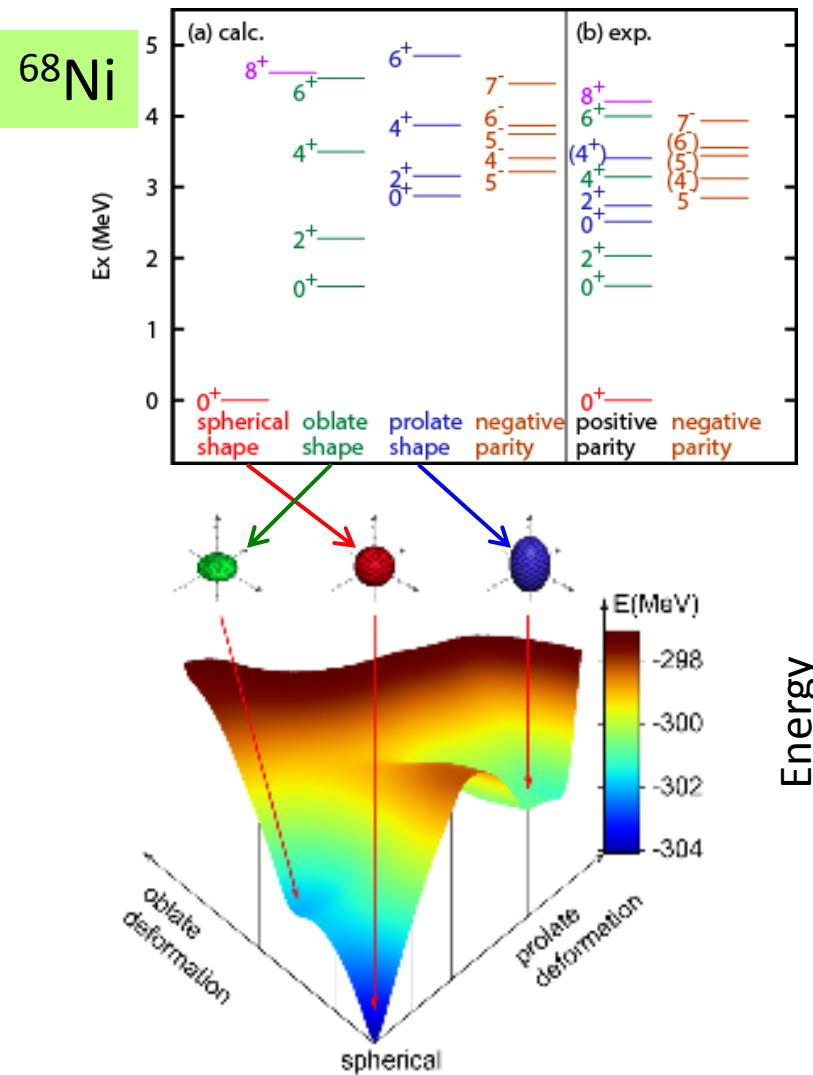
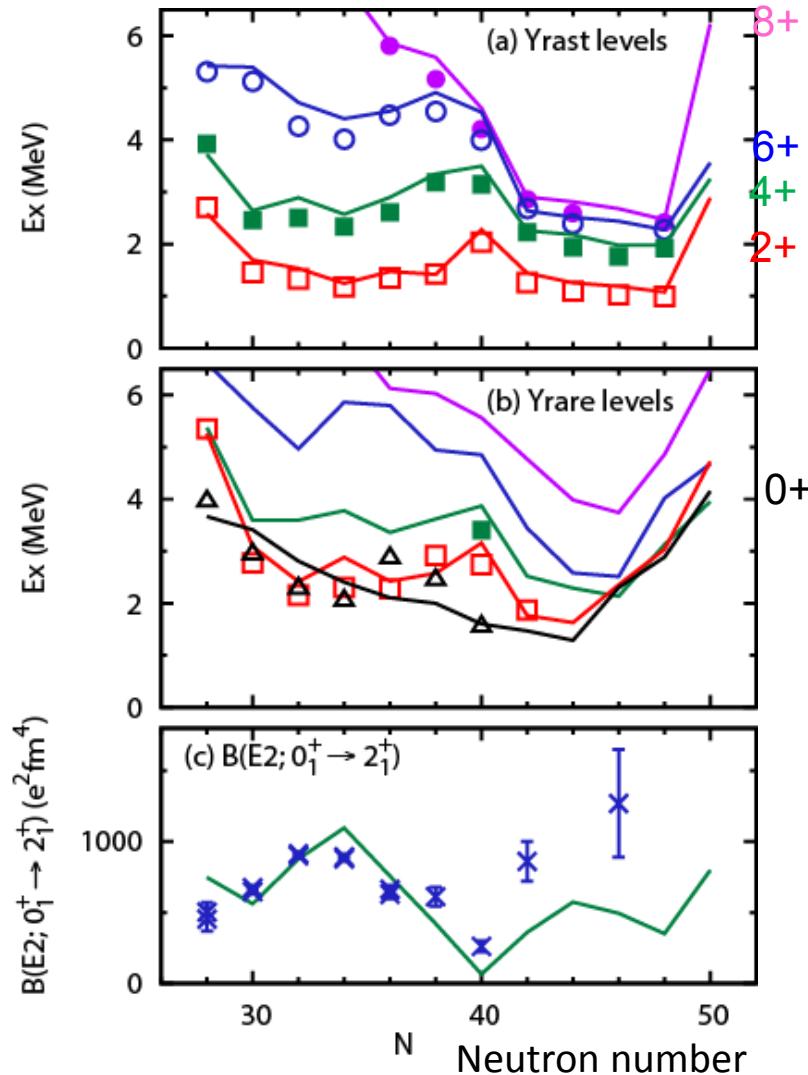
Neutron-rich nuclei: across two major shells
 ⇒ proton neutron different model space
 + truncation
 ⇒ full isospin-invariant space + MCSM

Neutron-rich Ni isotopes and shell evolution

Exp.: S. Suchyta PS2-A039

pf-shell + g9/2, d5/2 orbits w/o truncation
... 5.1×10^{15} M-scheme dim.

Y. Tsunoda, T. Otsuka, NS, M. Honma and Y. Utsuno,
PRC 89, 031301(R) (2014)



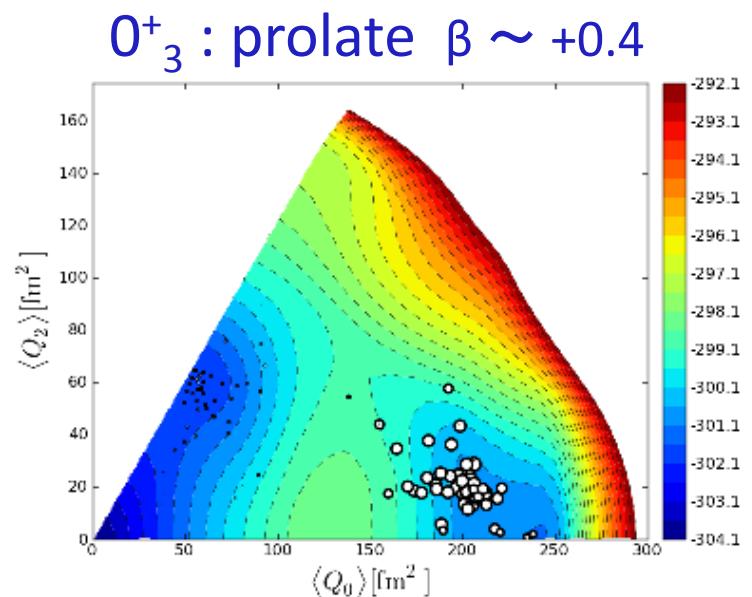
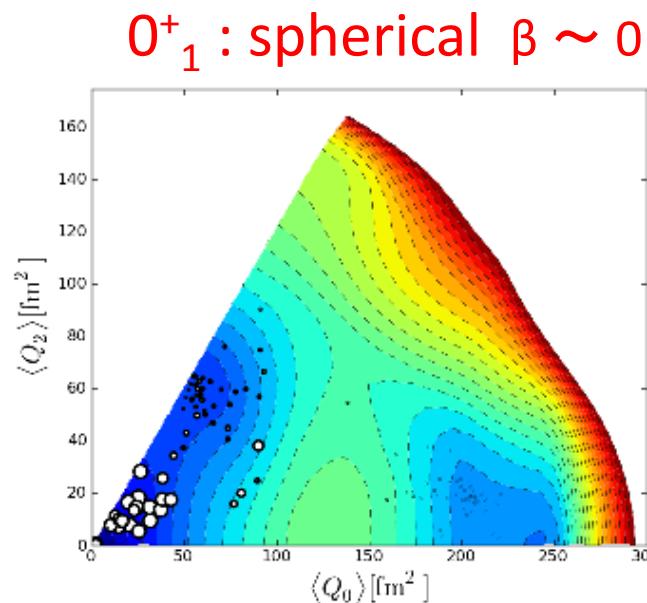
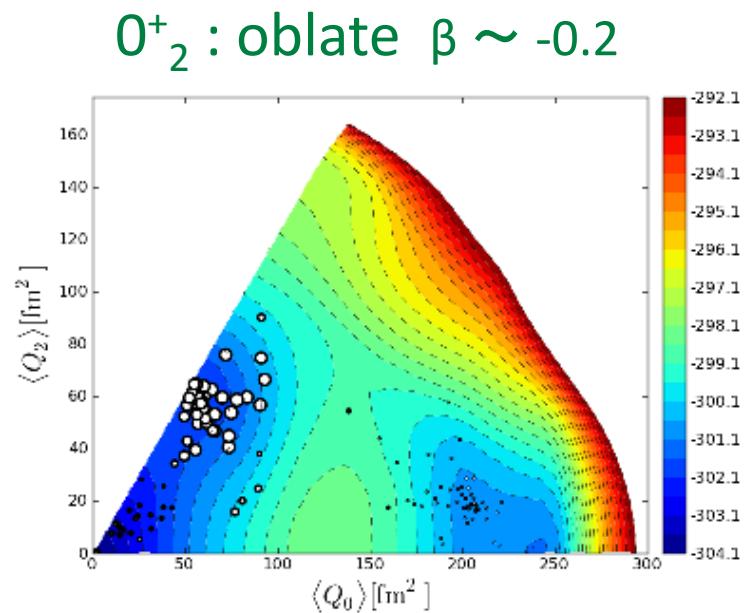
^{68}Ni 0^+ states \leftrightarrow different shapes

MCSM w.f. $|\Psi\rangle = \sum_{k=1}^{N_{MCSM} \sim 100} f_k P^{J,\pi} |\phi_k\rangle$

PES is calculated by Q-constraint Hartree-Fock calc.

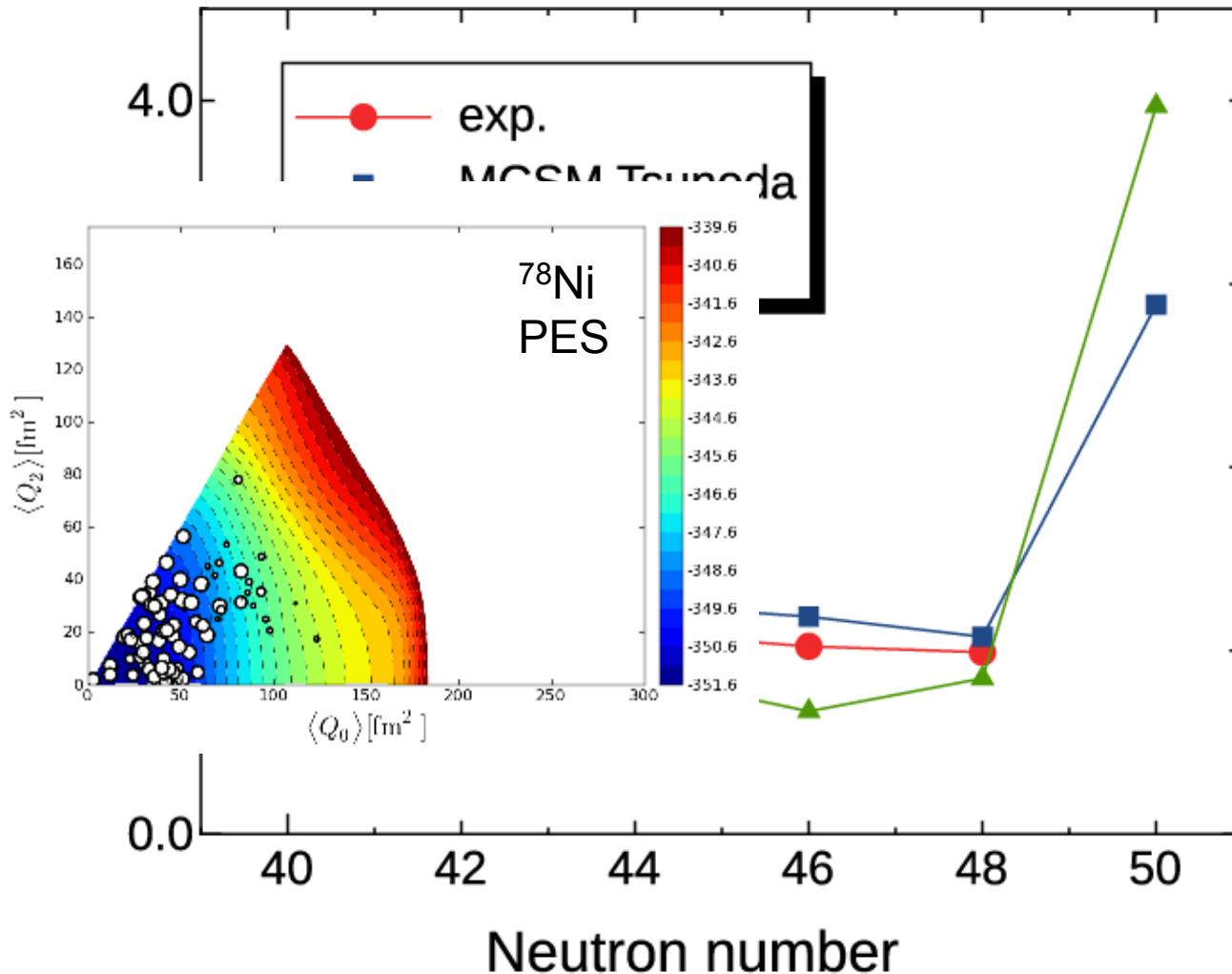
Location of circle : quadrupole deformation of $|\phi_k\rangle$

Area of circle : overlap probability $\frac{1}{N_k} |\langle \Psi | P^{J,\pi} | \phi_k \rangle|^2$



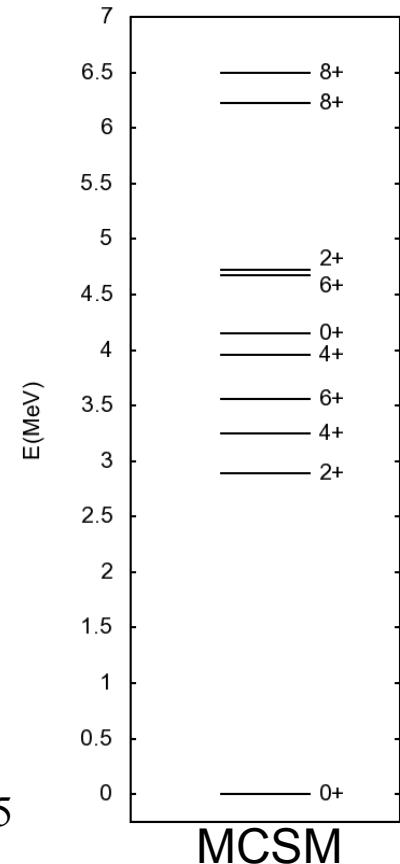
Towards ^{78}Ni : prediction

Excitation energies of Ni isotopes



$Z=28$, $N=50$
How magic?

Both SM results give
large 2^+ energy

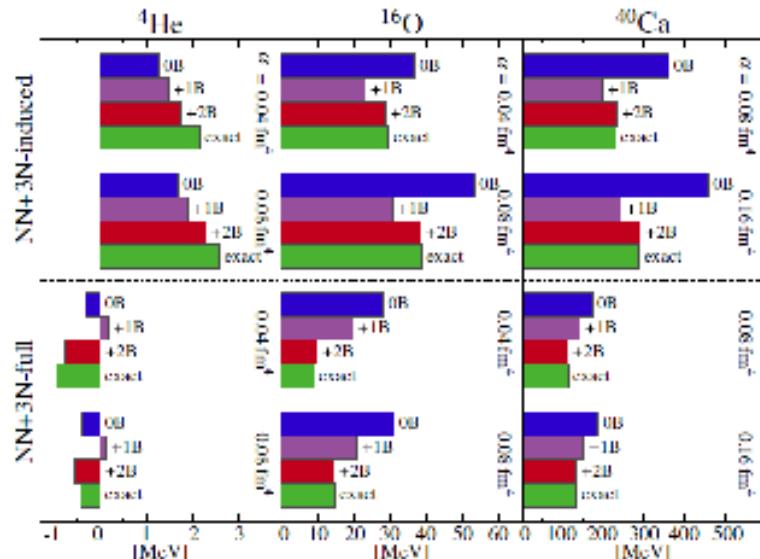
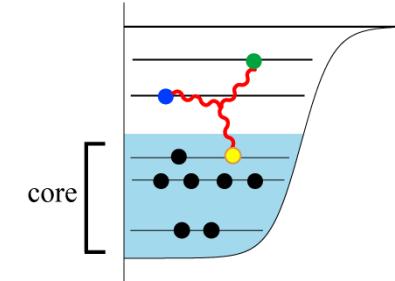


SM: Sieja and Nowacki, arXiv:1201.0373
MCSM: Y. Tsunoda *et al.*, PRC 89 031301(R) (2014)

π pf, ν f5pg9d5
full pfpg9d5

Towards SM calc. w/o phenomenological correction

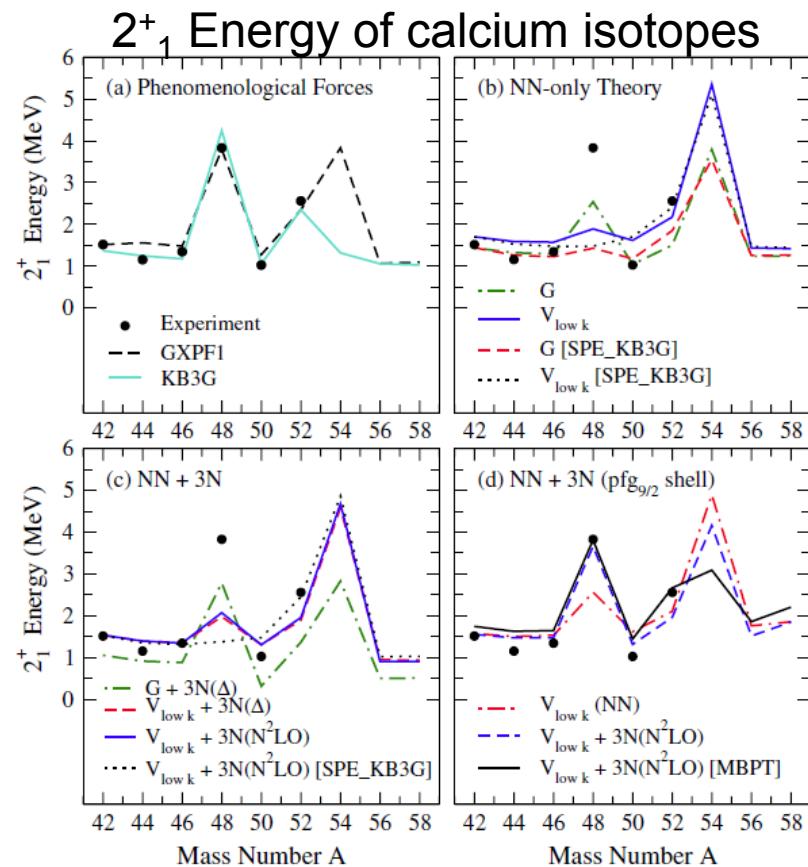
3-body force into normal ordered 2-body force



Normal ordered 2-body force is good approximation for no-core SM

Ref. R. Roth et al., PRL 109, 052501

Phenomenological fit of SM interaction
← 3-body force origin



SM calc. w/o phenom. correction

Normal ordered 2-body force

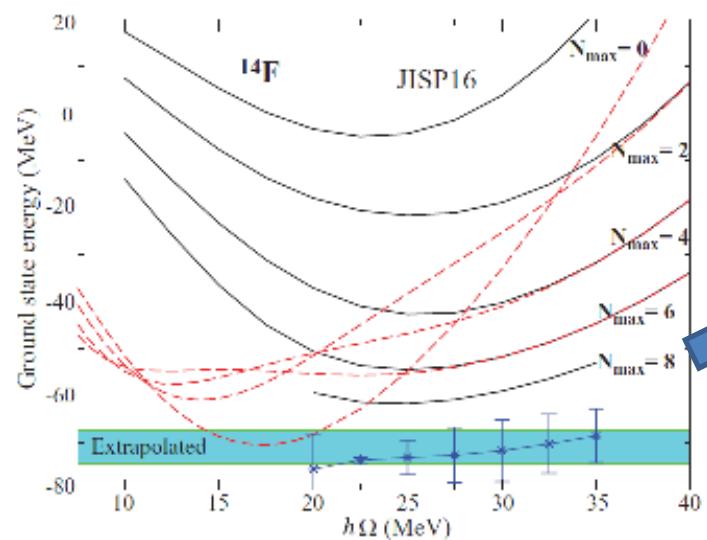
Ref. J. Holt, T. Otsuka, A. Schwenk and T. Suzuki, JPG 39 085111 (2012)

No-core shell model

Importance Truncated
NCSM

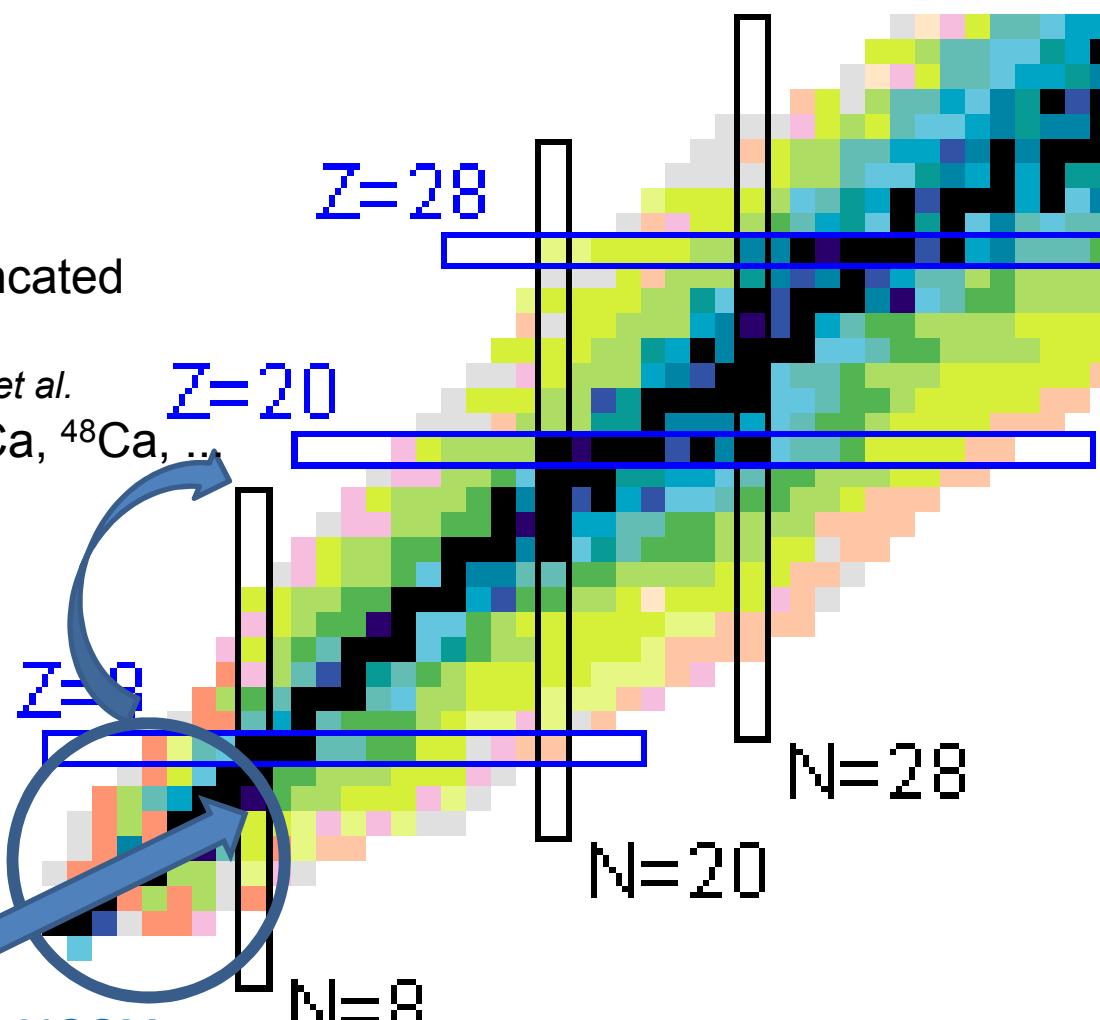
R. Roth, P. Navratil, *et al.*

Excitation energy ($N_{max}hw$) truncation
Extrapolation to infinite model space



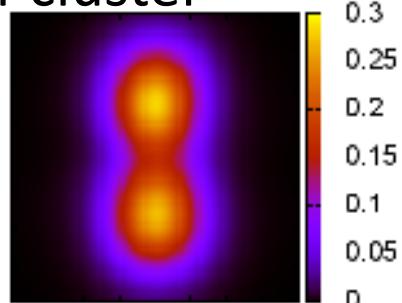
P. Maris, J. P. Vary and A. Shirokov
PRC 79, 014308 (2009)

NCSM and reaction ... S. Quaglioni *et al.*

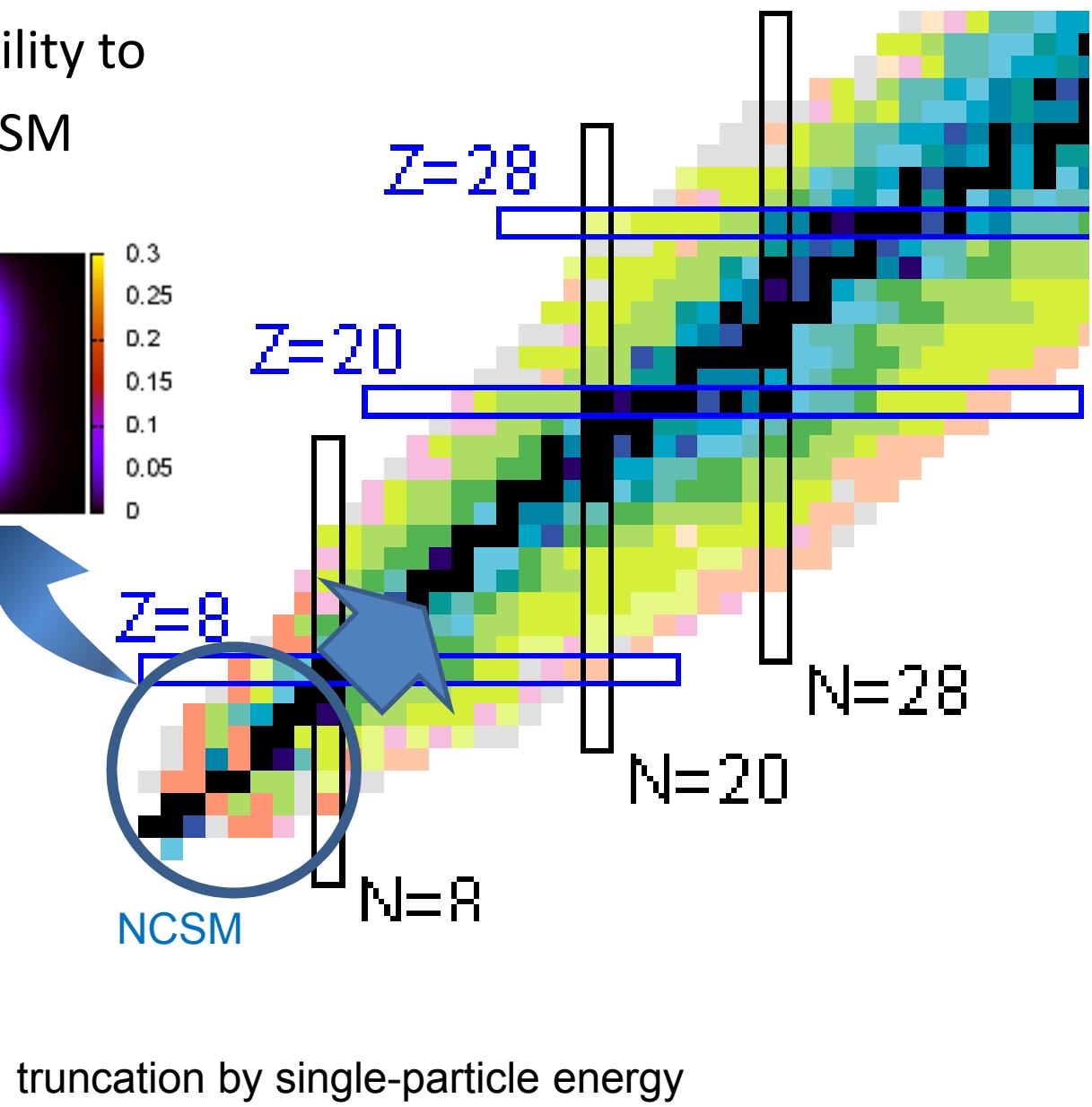
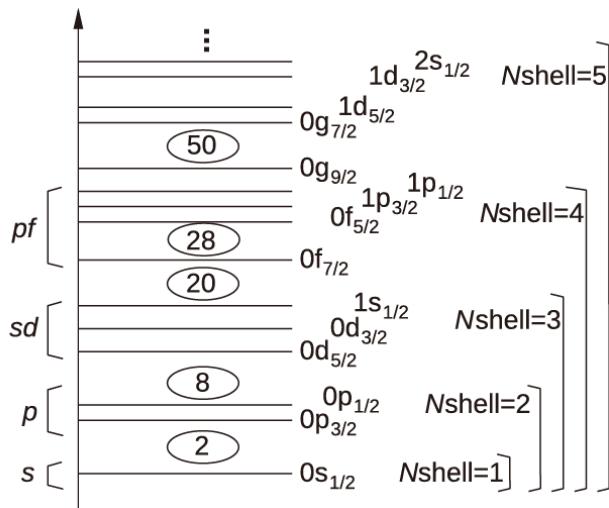


MCSM for no-core shell model calc.

- Enhance the applicability to heavier region by MCSM
- Description of cluster structure by MCSM



- N_{shell} truncation



truncation by single-particle energy

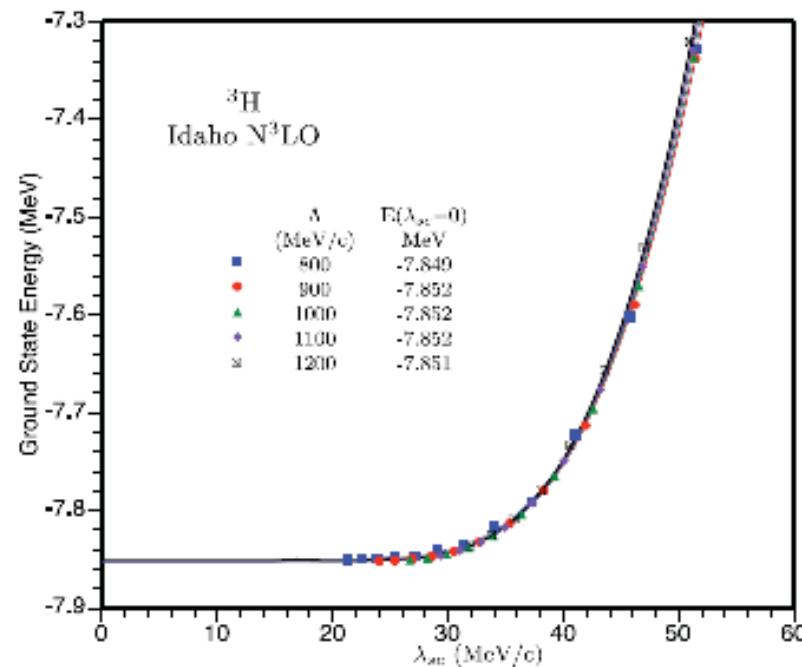
Extrapolation to infinite model space based on EFT

$$\lambda_{sc} = \sqrt{m\hbar\omega/(N + 3/2)} - \lambda^2/\Lambda$$

$$\Lambda = \sqrt{m(N + 3/2)\hbar\omega}$$

NCSM

Excitation energy ($N_{max}hw$) truncation

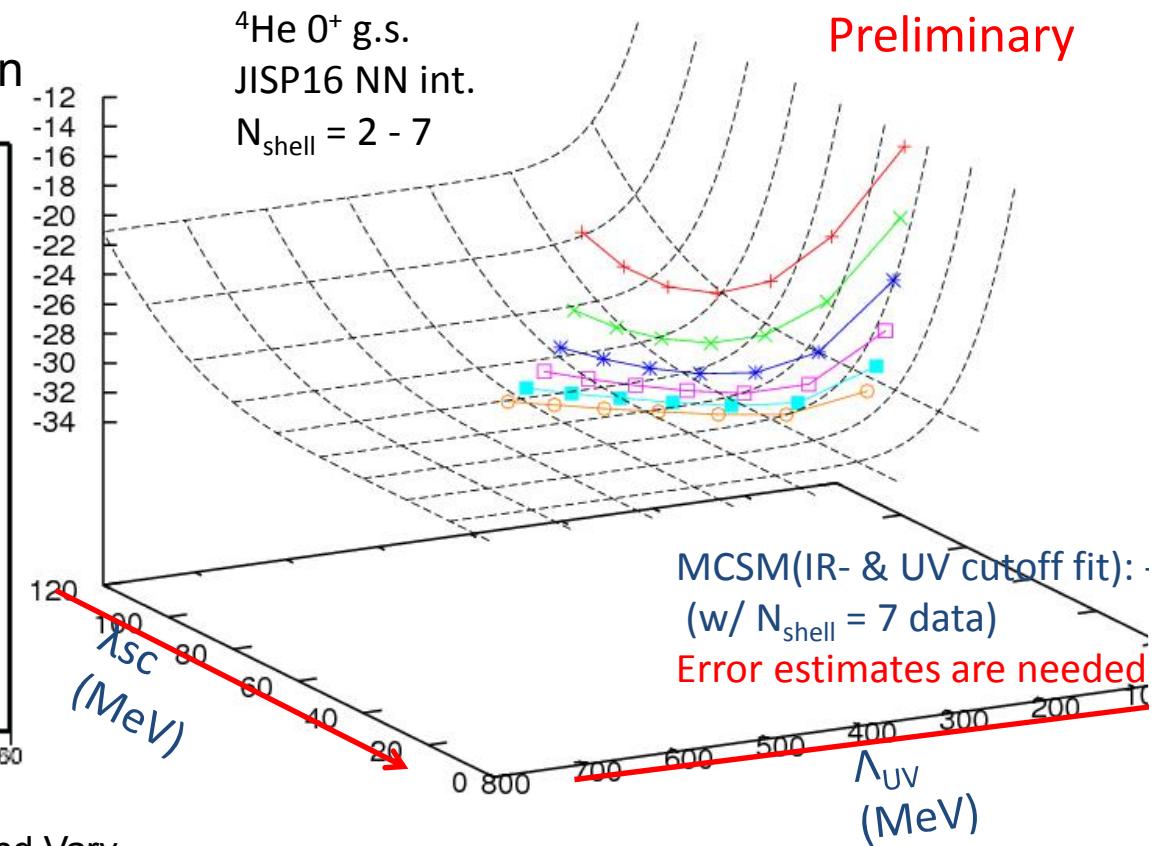


Ref. Coon, Avetian, Kruse, van Klock, Maris and Vary
PRC 86 054002 (2012)

MCSM

Excitation energy (N_{nshell}) truncation

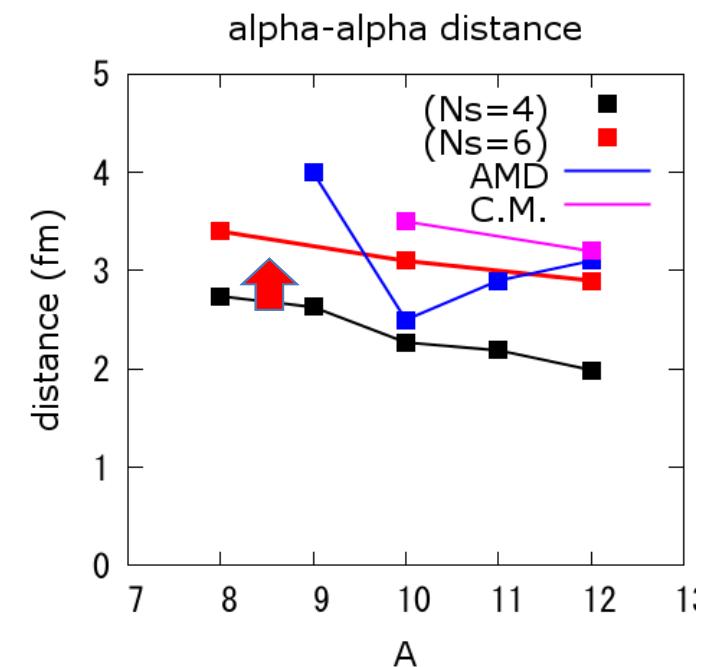
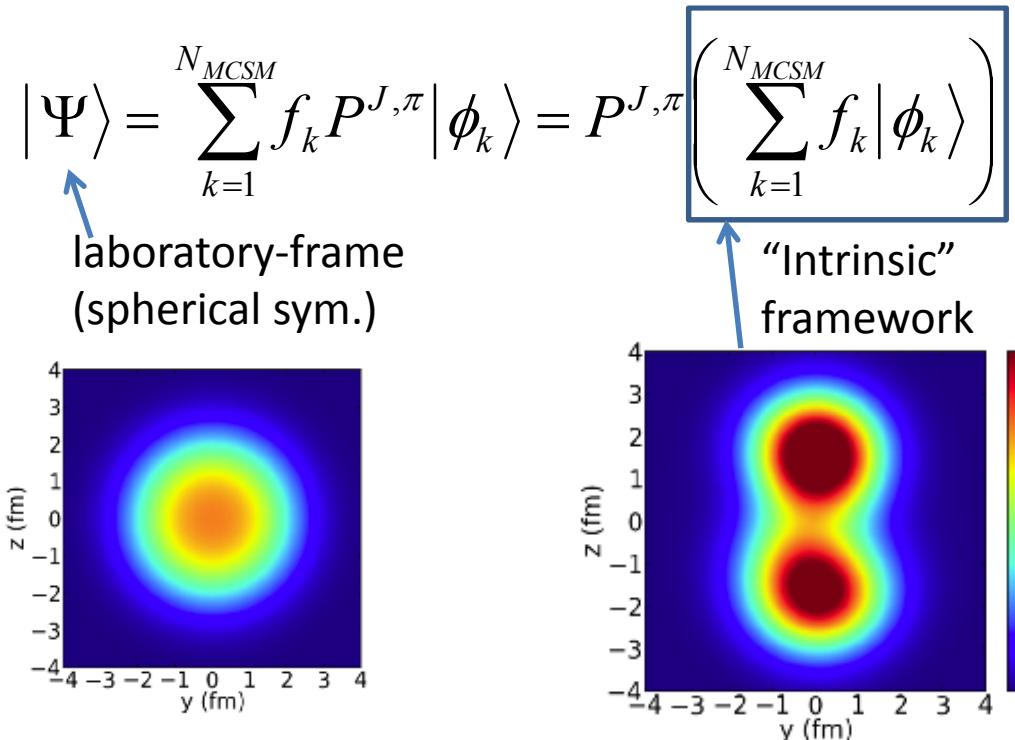
$$E(\lambda, \Lambda) = E(\lambda=0, \Lambda=\infty) + a \exp(-b/\lambda) + c \exp(-\Lambda^2/d^2)$$



On going: ${}^8\text{Be}(0^+)$, ${}^{12}\text{C}(0^+)$, ${}^{16}\text{O}(0^+)$, ...

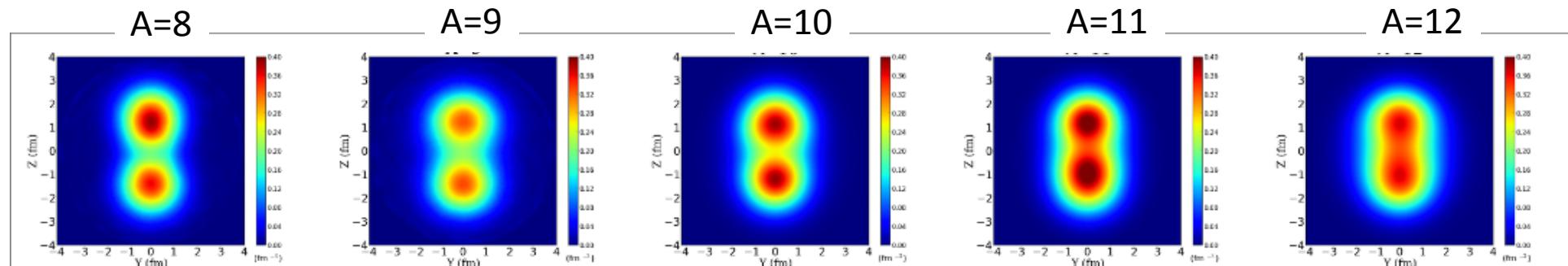
Density profile of the MCSM wave function

Density distribution of the ground state $J^\pi = 0^+$ of ${}^8\text{Be}$



[AMD: Y.Enyo et al. PRC 2003]
[C.M. : M.Ito, N.Itagaki et al, PRL 2008]

Matter density of Be isotopes JISP16 interaction



Summary

- HPCI project promotes nuclear physics. In HPCI Field 5 Subject 2, we developed the “advanced” MCSM to run on K computer
 - energy-variance extrapolation, code R&D
 - unified description of neutron-rich Ni isotopes, shell evolution and shape coexistence
 - test the feasibility of no-core MCSM and propose a new method to draw the density profile
- MCSM purvey analysis of “intrinsic” shapes and their mixing, which enlighten cluster structure and shape coexistence.

Collaborators

HPCI project, Field 5

- Takaharu Otsuka (Tokyo) Co-chair
- Takashi Abe (Tokyo) Poster AS2-A018
- Shuichiro Ebata (Hokkaido) Poster PS2-B022
- Michio Honma (Aizu)
- Yoritaka Iwata (CNS, Tokyo) Poster PS2-A061
- Takahiro Mizusaki (Senshu Univ.)
- Tomoaki Togashi (CNS, Tokyo) Poster PS2-A035
- Naofumi Tsunoda (CNS, Tokyo)
- Yusuke Tsunoda (Tokyo) Poster PS2-A037
- Yutaka Utsuno (JAEA) Oral Jun/3
- Tooru Yoshida (CNS, Tokyo) Poster PS2-A005
- James Vary (Iowa)
- Pieter Maris (Iowa) ICNT workshop, next week

Thank you !