

第47回理研原子核コロキウム
仁科ホール
June. 28, 2010

不安定核と核力

Exotic nuclei and nuclear forces

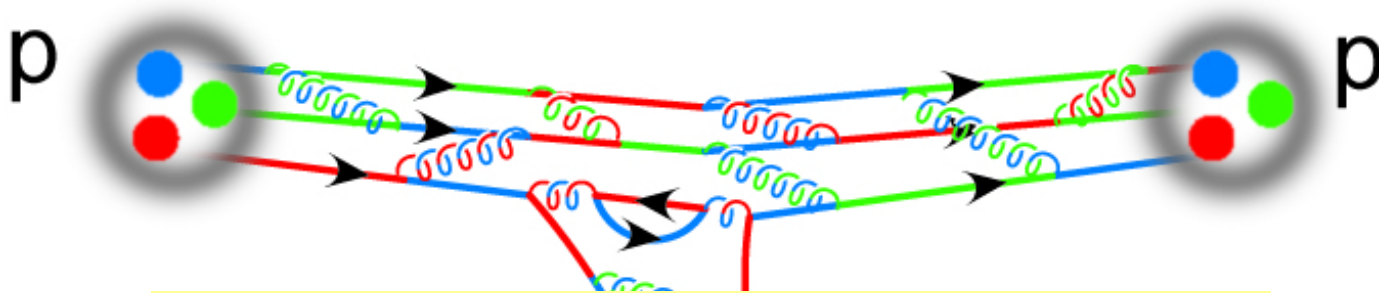


Takaharu Otsuka
University of Tokyo / MSU



3時40分

Graphical Image of NN force



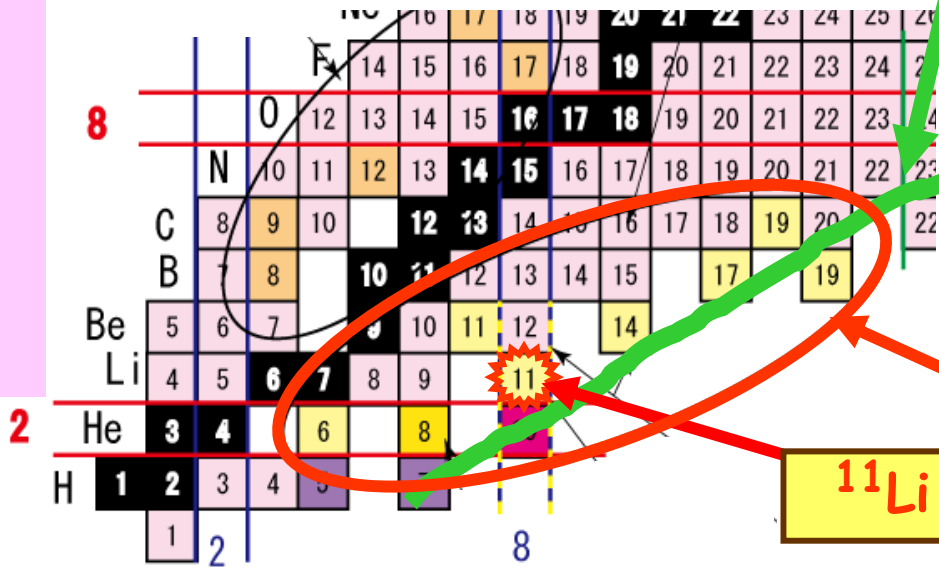
What simple regularity can we imagine
from such a complexity ?



Studies on exotic nuclei in the 80~90's

Left-lower part of the Nuclear Chart

Proton number →



Stability line and drip lines are not so far from each other

→ Physics of loosely bound neutrons, e.g., halo while other issues like ^{32}Mg

neutron halo

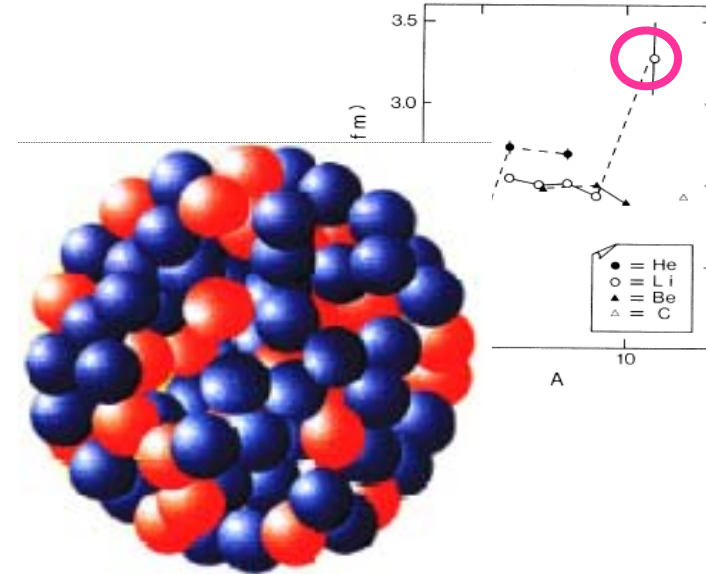
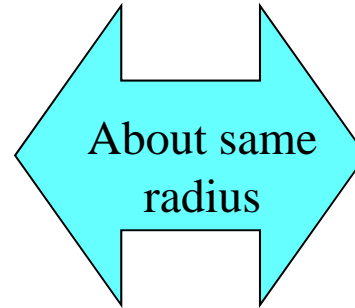
^{11}Li

- A nuclei (mass number)
- stable
- exotic
- with halo

Neutron number →

Neutron halo

Strong tunneling of loosely bound excess neutrons



VOLUME 55, NUMBER 24

PHYSICAL REVIEW LETTERS

9 DECEMBER 1985

Measurements of Interaction Cross Sections and Nuclear Radii in the Light p -Shell Region

I. Tanihata,^(a) H. Hamagaki, O. Hashimoto, Y. Shida, and N. Yoshikawa
Institute for Nuclear Study, University of Tokyo, Tanashi, Tokyo 188, Japan

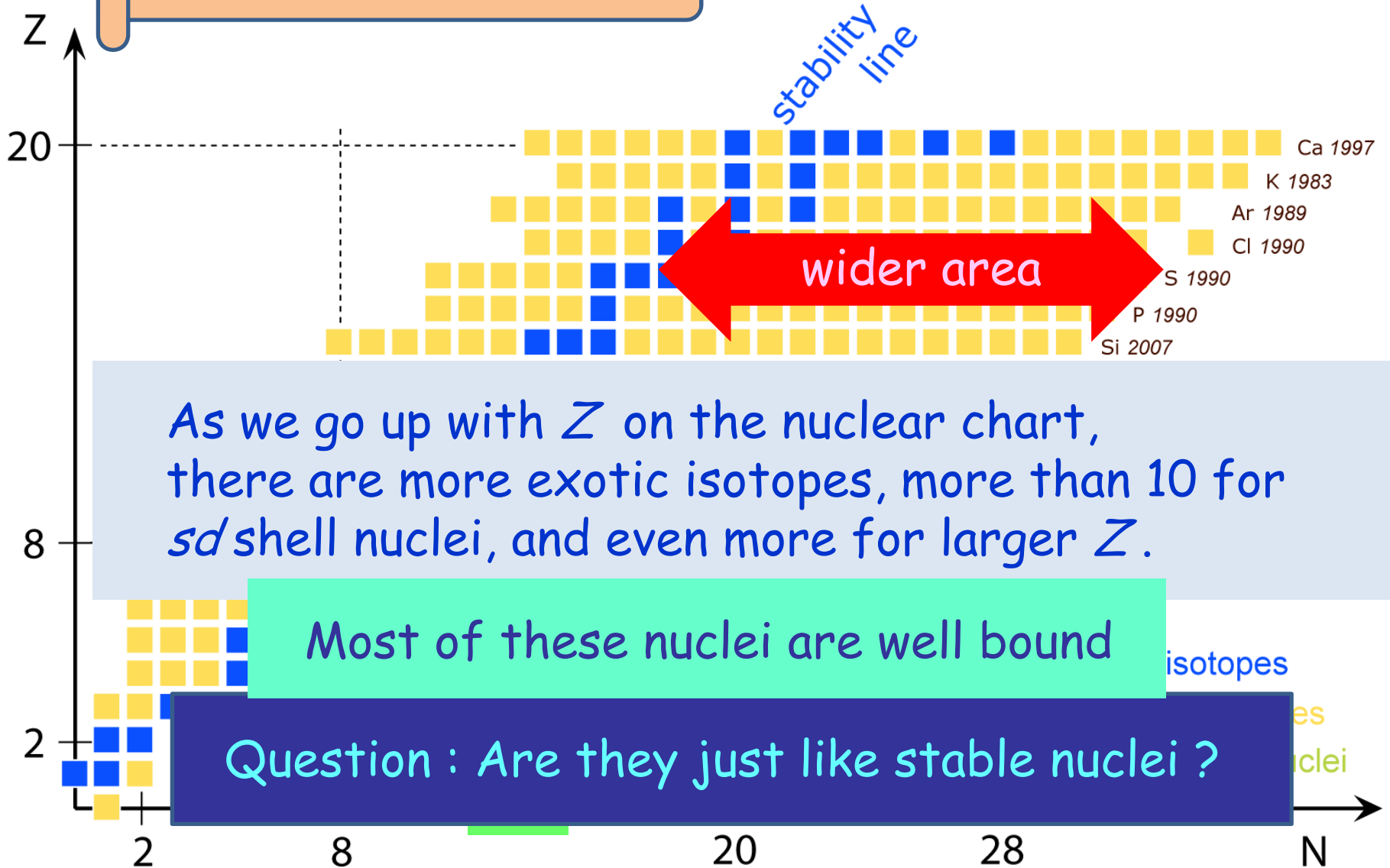
K. Sugimoto,^(b) O. Yamakawa, and T. Kobayashi
Nuclear Science Division, Lawrence Berkeley Laboratory, University of California, Berkeley, California 94720

and

N. Takahashi

Nuclear Chart - Left Lower Part -

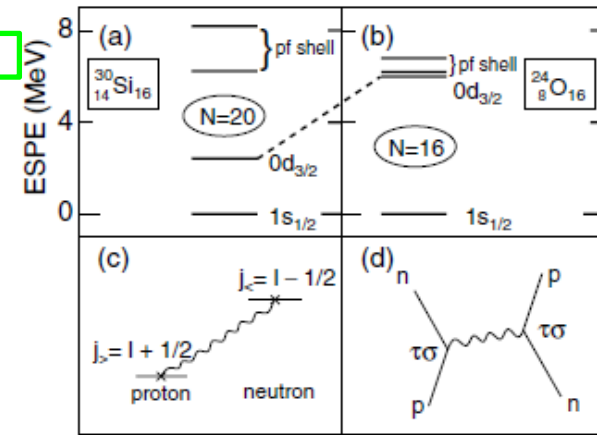
For Z larger, more exotic isotopes.



Magic Numbers in Exotic Nuclei and Spin-Isospin Properties of the NN Interaction

Takaharu Otsuka,^{1,2} Rintaro Fujimoto,¹ Yutaka Utsuno,³ B. Alex Brown,⁴ Michio Honma,⁵ and Takahiro Mizusaki⁶
¹Department of Physics, University of Tokyo, Hongo, Bunkyo-ku, Tokyo, 113-0033, Japan
²RIKEN, Hirosawa, Wako-shi, Saitama 351-0198, Japan

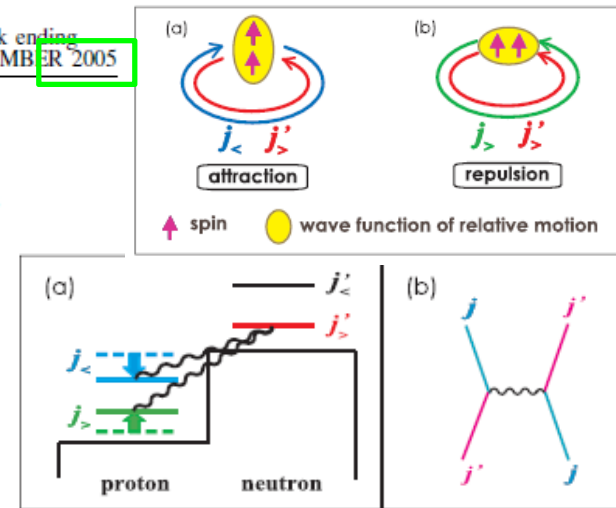
魔法数までもがスピンスピン核力で変わる



Evolution of Nuclear Shells due to the Tensor Force

Takaharu Otsuka,^{1,2,3,*} Toshio Suzuki,⁴ Rintaro Fujimoto,¹ Hubert Grawe,⁵ and Yoshinori Akaishi⁶
¹Department of Physics, University of Tokyo, Hongo, Bunkyo-ku, Tokyo 113-0033, Japan
²Center for Nuclear Study, University of Tokyo, Hongo, Bunkyo-ku, Tokyo 113-0033, Japan
³RIKEN, Hirosawa, Wako-shi, Saitama 351-0198, Japan

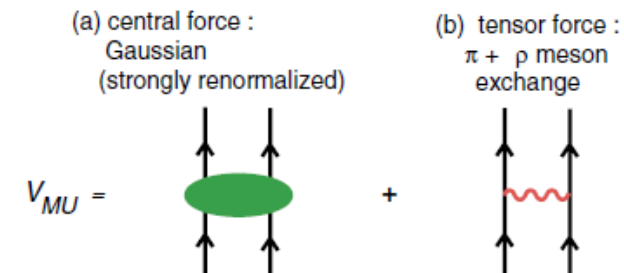
テンソル力の特異な効果



Novel Features of Nuclear Forces and Shell Evolution in Exotic Nuclei

Takaharu Otsuka,^{1,2} Toshio Suzuki,³ Michio Honma,⁴ Yutaka Utsuno,⁵ Naofumi Tsunoda,¹ Koshiroh Tsukiyama,¹ and Morten Hjorth-Jensen⁶

テンソル力と中心力によるWeinberg 型モデル



Focus of this talk

Changes in the structure of exotic nuclei
as compared to stable ones

shell structure (incl. magic numbers)
binding energies, drip lines

experimentally, a lot of recent progress by
combining direct reactions and γ -ray spectroscopy

Nuclear and hadronic forces are responsible
for such changes

2-body and 3-body forces

Monopole-based universal force V_{MU}

→ similarity to Weinberg's chiral perturbation

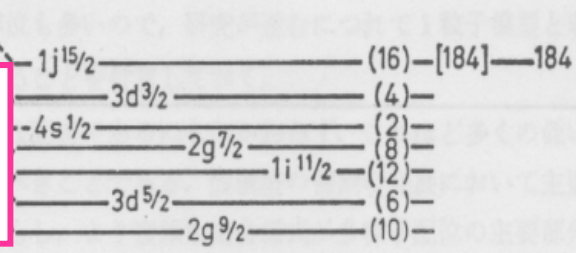
Importance of Fujita-Miyazawa force

Eigenvalues of HO potential

Magic numbers Mayer and Jensen (1949)

5ħω
4ħω
3ħω
2ħω
1ħω
0

Will these magic numbers remain the same in exotic nuclei?



126

28
20
8
2



R SHELL MODEL

図 2-23 1 粒子軌道の順序。図は G. Mayer and J. H. D. Jensen, *Elementary Theory of Nuclear Shell Structure*, p. 58, Wiley, New York, 1955 からとった。

From undergraduate nuclear-physics course,

density saturation
+ short-range NN interaction
+ spin-orbit splitting

→ Mayer-Jensen's magic number
with rather **constant gaps**
(except for gradual A dependence)

robust feature -> no way out

As N or Z is changed to a large extent in exotic nuclei, the shell structure is changed (evolved) by

- Monopole component of the NN interaction

$$v_{m;j,j'} = \frac{\sum_{k,k'} \langle jk j' k' | V | jk j' k' \rangle}{\sum_{k,k'} 1},$$

➔ Averaged over possible orientations

Linearity: Shift

$$\Delta \epsilon_j = v_{m;j,j'} n_{j'}$$

$n_{j'}$: # of particles in j'

$\langle n_{j'} \rangle$ can be ~ 10 in exotic nuclei

-> effect quite relevant to neutron-rich exotic nuclei

Strasbourg group made a major contribution in initiating systematic use of the monopole interaction. (Poves and Zuker, Phys. Rep. 70, 235 (1981))

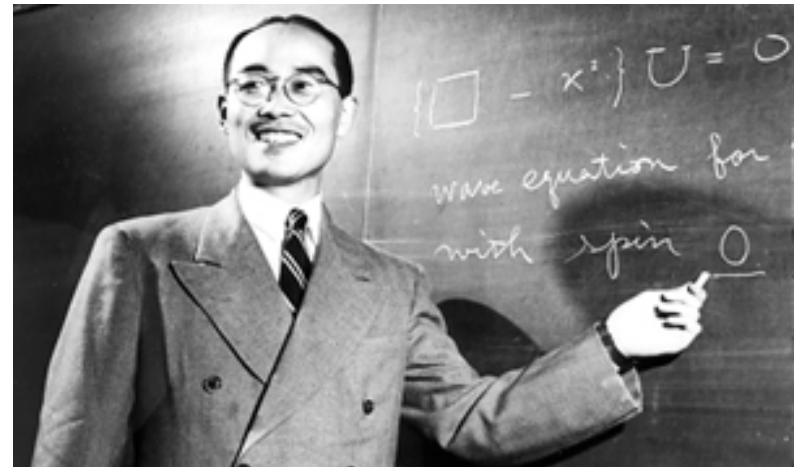
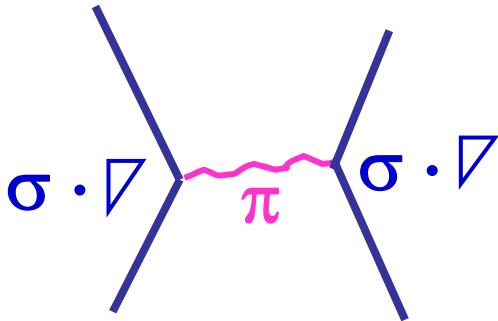
Tensor Interaction by pion exchange

$$V_T = (\tau_1 \tau_2) ([\sigma_1 \sigma_2]^{(2)} Y^{(2)}(\Omega)) Z(r)$$

contributes
only to **S=1** states

relative motion

π meson : primary source



Yukawa

ρ meson ($\sim \pi + \pi$) : minor ($\sim 1/4$) cancellation

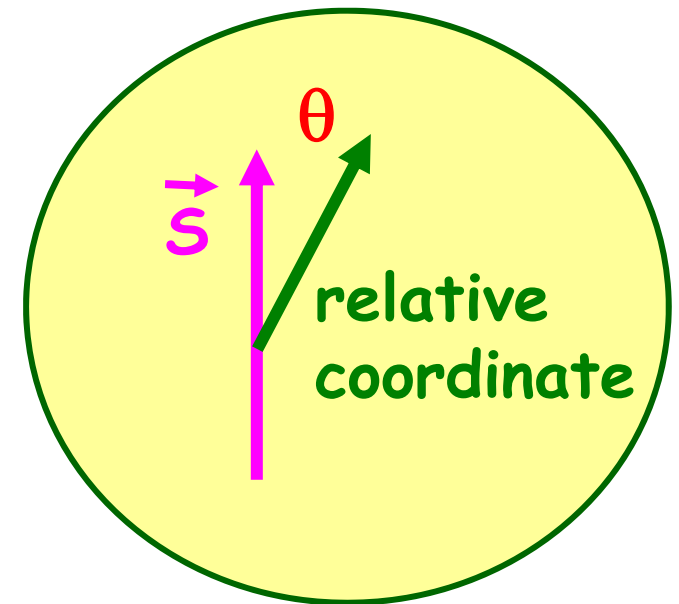
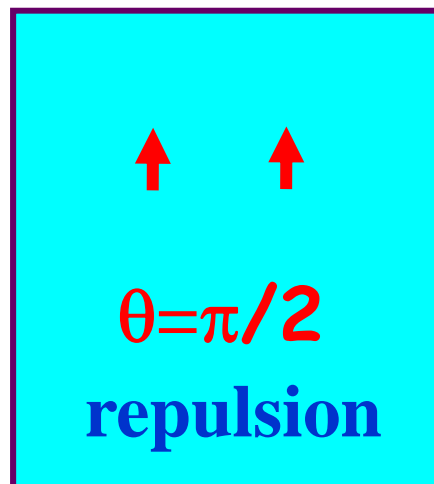
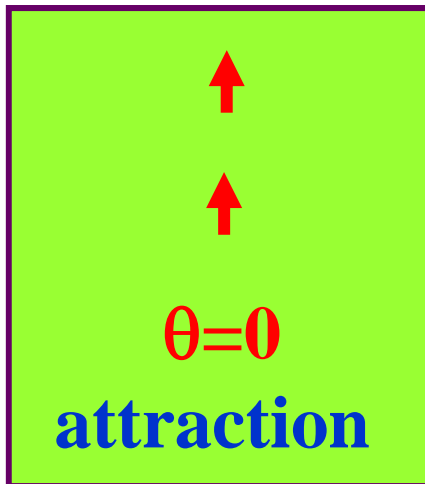
Ref: Osterfeld, Rev. Mod. Phys. 64, 491 (92)

How does the tensor force work ?

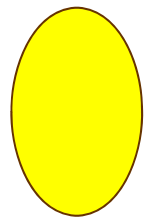
Spin of each nucleon \uparrow is parallel, because the total spin must be $S=1$

The potential has the following dependence on the angle θ with respect to the total spin S .

$$V \sim Y_{2,0} \sim 1 - 3 \cos^2 \theta$$



Monopole effects due to the tensor force - An intuitive picture -

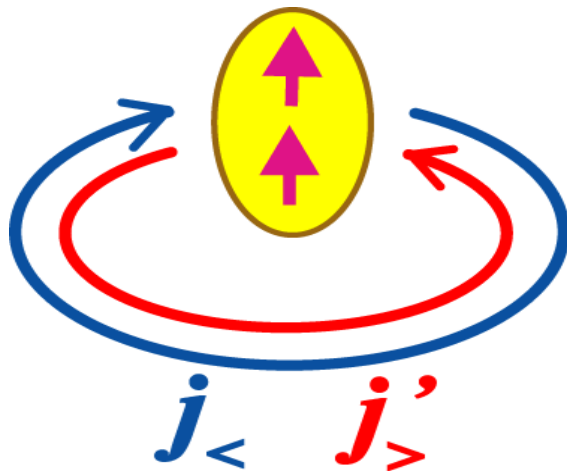


wave function of **relative motion**

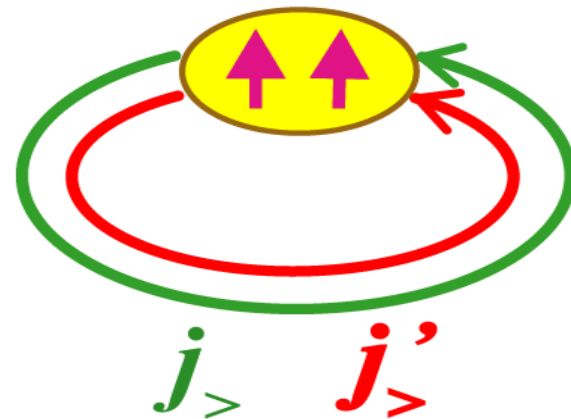
↑↑ spin of nucleon

large relative momentum

small relative momentum



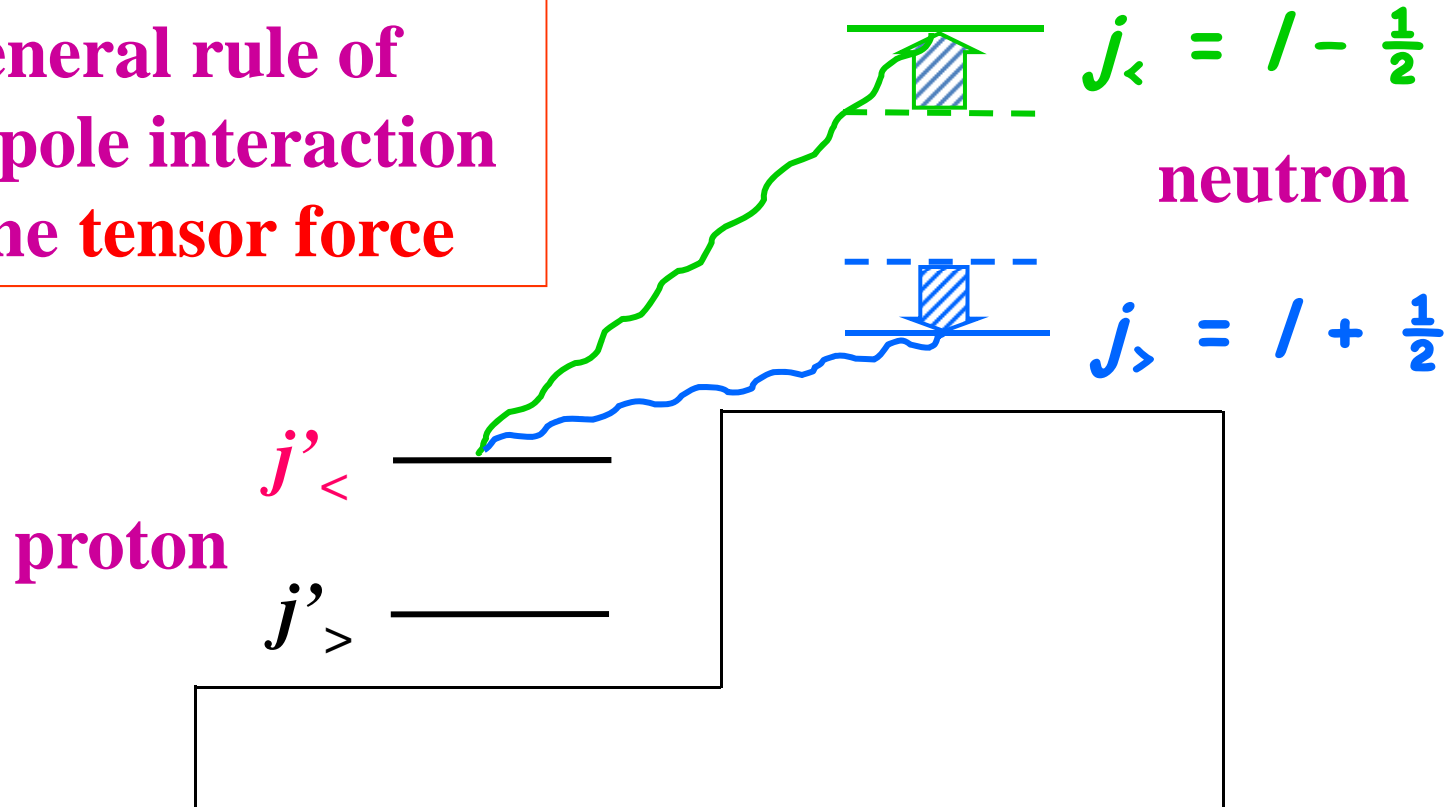
attractive



repulsive

$$j_{>} = l + \frac{1}{2}, \quad j_{<} = l - \frac{1}{2}$$

General rule of
monopole interaction
of the **tensor force**



Identity for tensor monopole interaction

$$(2j_> + 1) v_{m,T}^{(j' j_>)} + (2j_< + 1) v_{m,T}^{(j' j_<)} = 0$$

$v_{m,T}$: monopole strength for isospin T

An example with ${}_{51}\text{Sb}$ isotopes

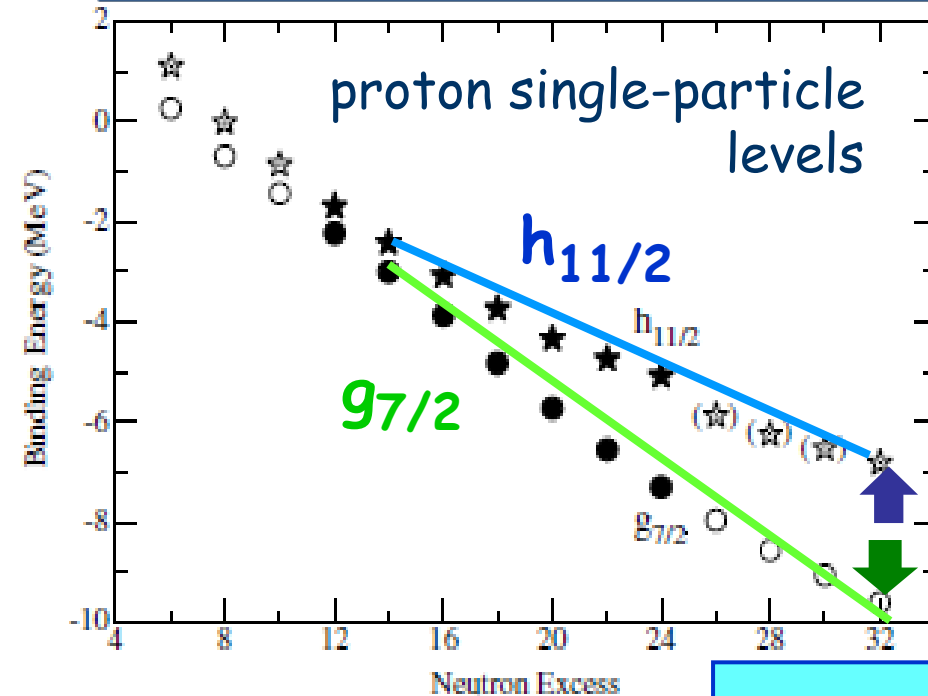
VOLUME 92, NUMBER 16

PHYSICAL REVIEW LETTERS

week ending
23 APRIL 2004

Is the Nuclear Spin-Orbit Interaction Changing with Neutron Excess?

J. P. Schiffer,¹ S. J. Freeman,^{1,2} J. A. Caggiano,³ C. Deibel,³ A. Heinz,³ C.-L. Jiang,¹ R. Lewis,³ A. Parikh,³ P. D. Parker,³
K. E. Rehm,¹ S. Sinha,¹ and J. S. Thomas⁴



$Z=51$ isotopes

change driven
by neutrons in $1h_{11/2}$

$h_{11/2} - h_{11/2}$ repulsive \uparrow

$h_{11/2} - 97/2$ attractive \downarrow

$\pi + \rho$ meson exchange tensor force
(splitting increased by ~ 2 MeV)

No mean field theory,
(Skyrme, Gogny, RMF)
explained this before.

What are the major monopole components
in the shell-model effective interaction
which are successful in systematic
description over many nuclei ?

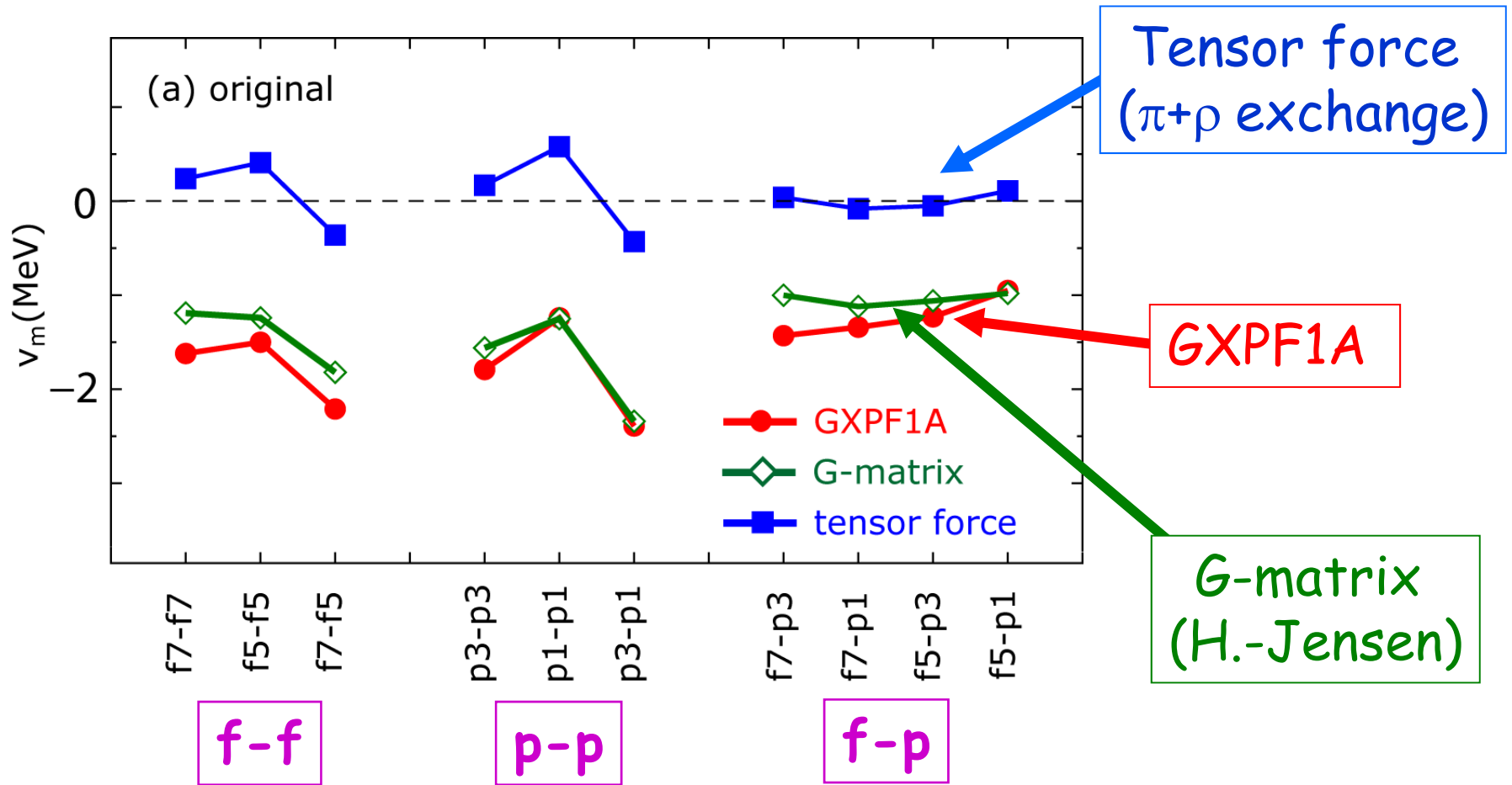
Example taken from pf-shell

GXPF1A interaction *

G-matrix obtained from Bonn-C potential + 3rd order Q_{box}
+ empirical refinement by χ^2 -fit to experimental data

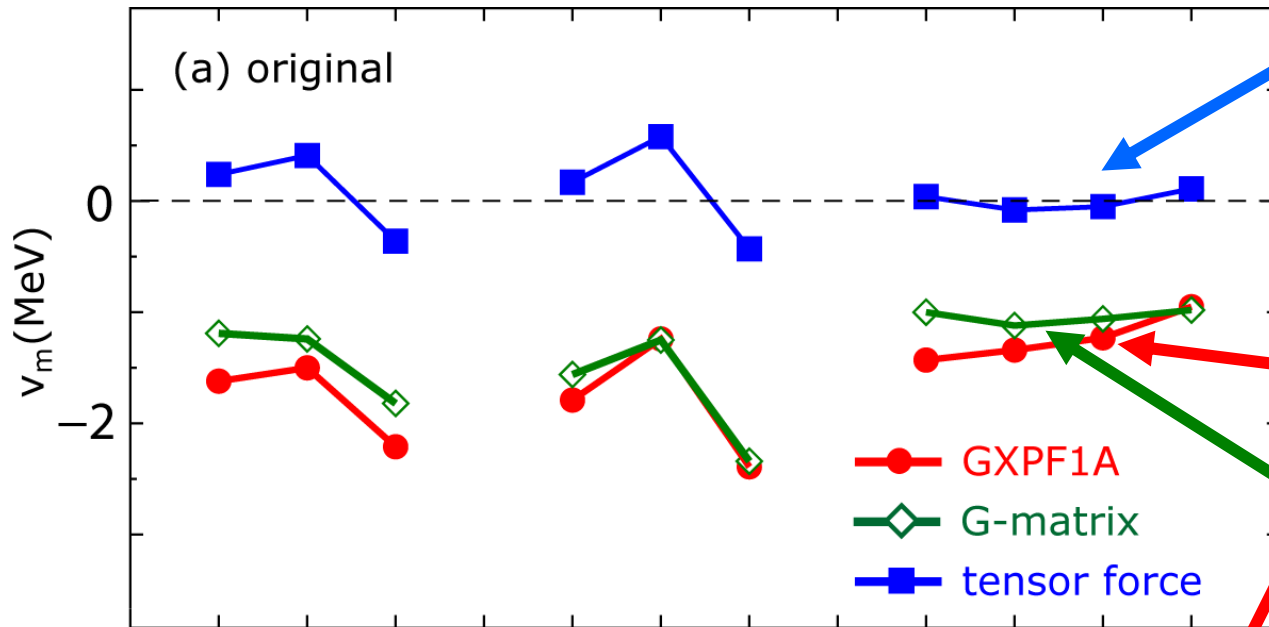
*) M. Honma et al., PRC65 (2002) 061301(R)

T=0 monopole interactions in the pf shell



"Local pattern" \leftarrow tensor force

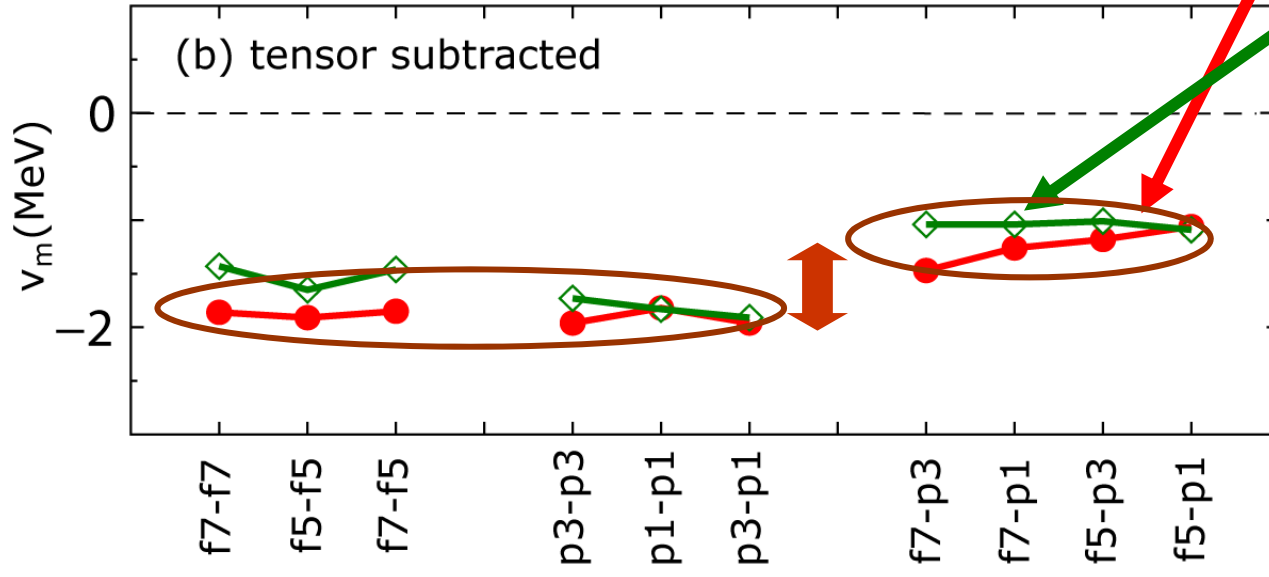
$T=0$ monopole interactions in the pf shell



Tensor force
($\pi+\rho$ exchange)

GXPF1A

G-matrix
(H.-Jensen)



Tensor component is subtracted

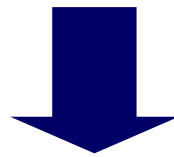
The central force is modeled by a Gaussian function

$$V = V_0 \exp(-r/\mu)^2 \quad (S, T \text{ dependences})$$

with $V_0 = -166 \text{ MeV}$, $\mu = 1.0 \text{ fm}$,

(S,T) factor	(0,0)	(1,0)	(0,1)	(1,1)

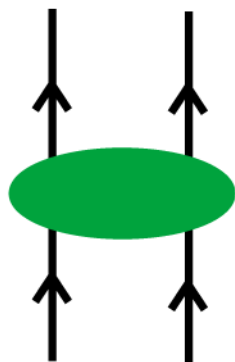
relative strength	1	1	0.6	-0.8



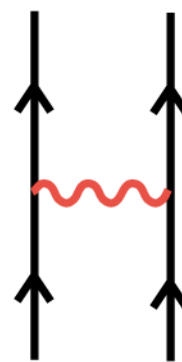
Can we explain the difference between f-f/p-p and f-p ?

Systematic description of monopole properties of exotic nuclei can be obtained by an extremely simple interaction as

(a) central force :
Gaussian
(strongly renormalized)



(b) tensor force :
 $\pi + \rho$ meson exchange



$$V_{MU} =$$

+

Parameters are fixed for all nuclei

The tensor part of the effective NN interaction for valence nucleons is similar to the bare tensor force.

bare tensor force : $\pi + \rho$ meson exchange

S. Weinberg,
PLB 251, 288 (1990)

Central force:
strongly renormalized

tive potential gives a local coordinate-space two-nucleon potential:

$$V_{2\text{-nucleon}} = 2(C_S + C_T \sigma_1 \cdot \sigma_2) \delta^3(\mathbf{x}_1 - \mathbf{x}_2) - \left(\frac{2g_A}{F_\pi}\right)^2 (\mathbf{t}_1 \cdot \mathbf{t}_2) (\sigma_1 \cdot \nabla_1) (\sigma_2 \cdot \nabla_2) Y(|\mathbf{x}_1 - \mathbf{x}_2|) - (1' \leftrightarrow 2'),$$

Tensor force is explicit

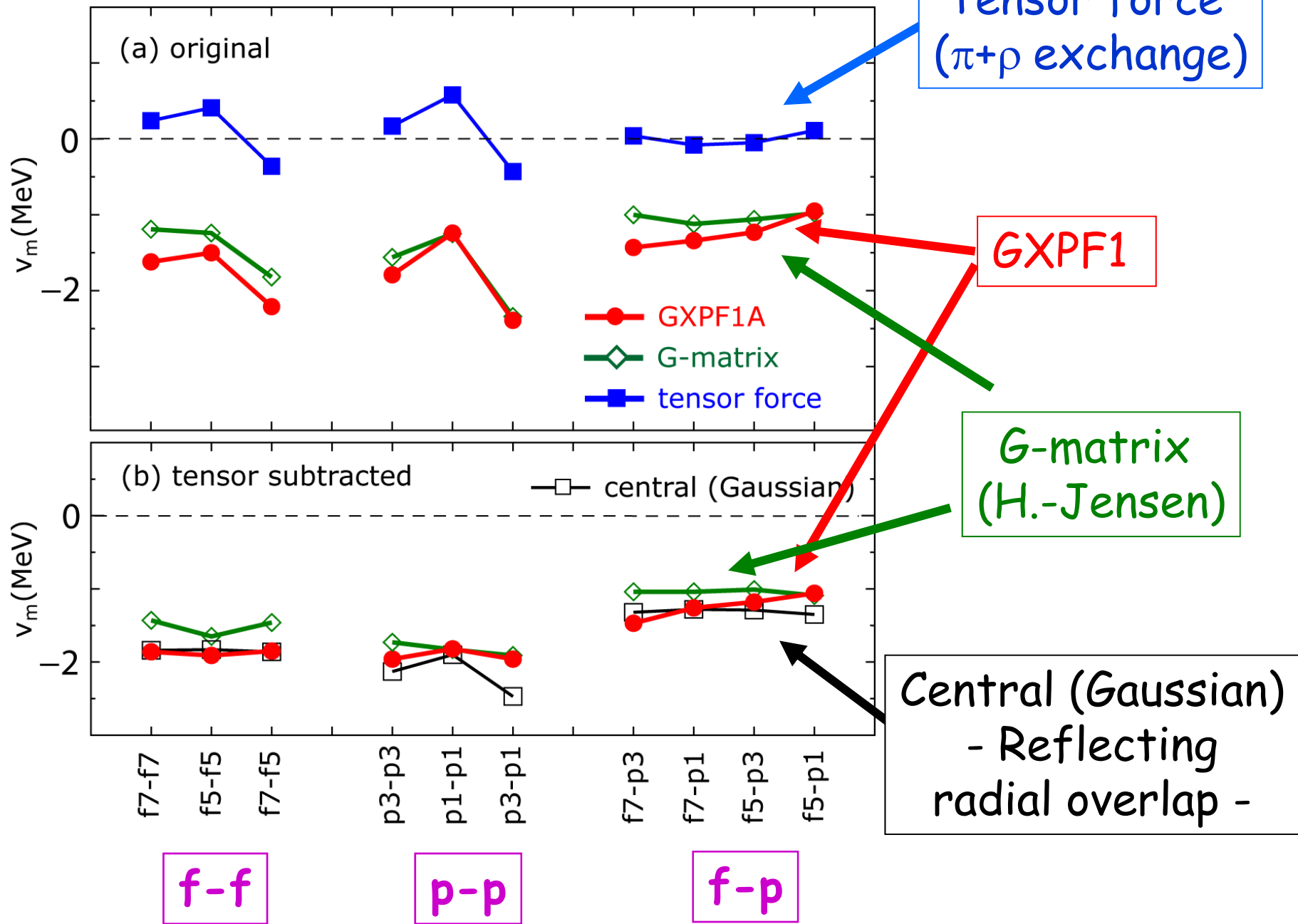
In nuclei

finite range

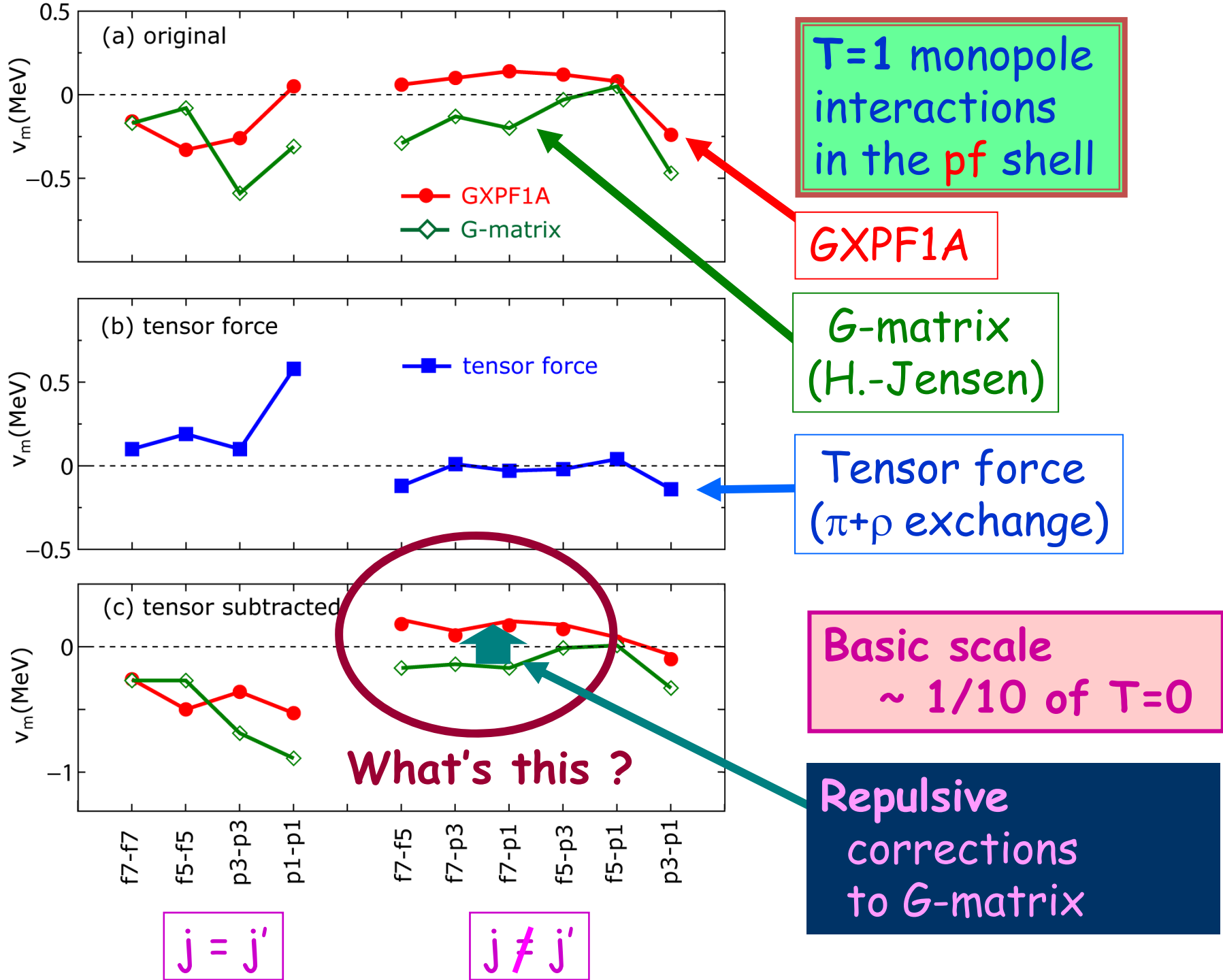
$\pi + \rho$ exchange

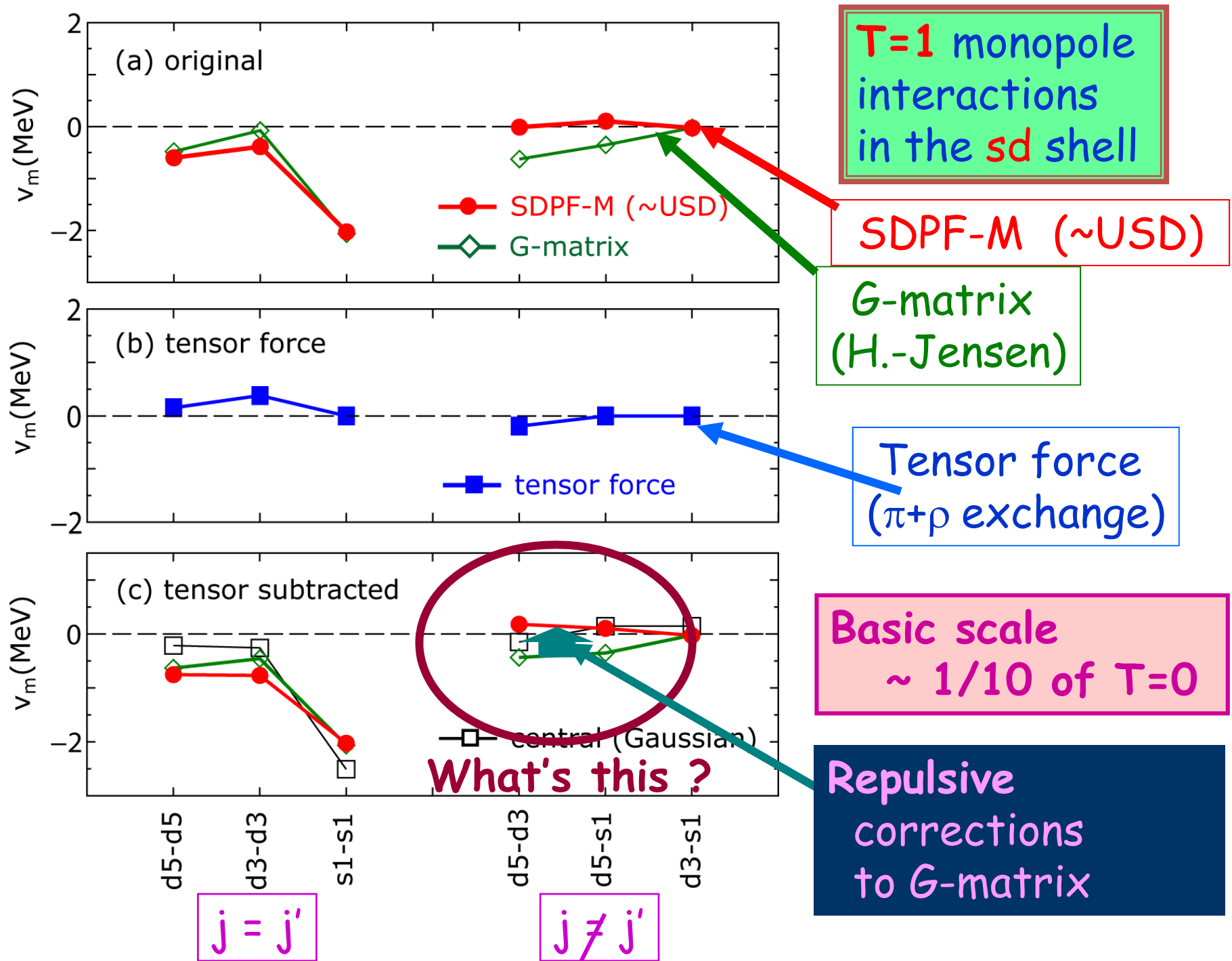
→ Chiral Perturbation of QCD

T=0 monopole interactions in the pf shell

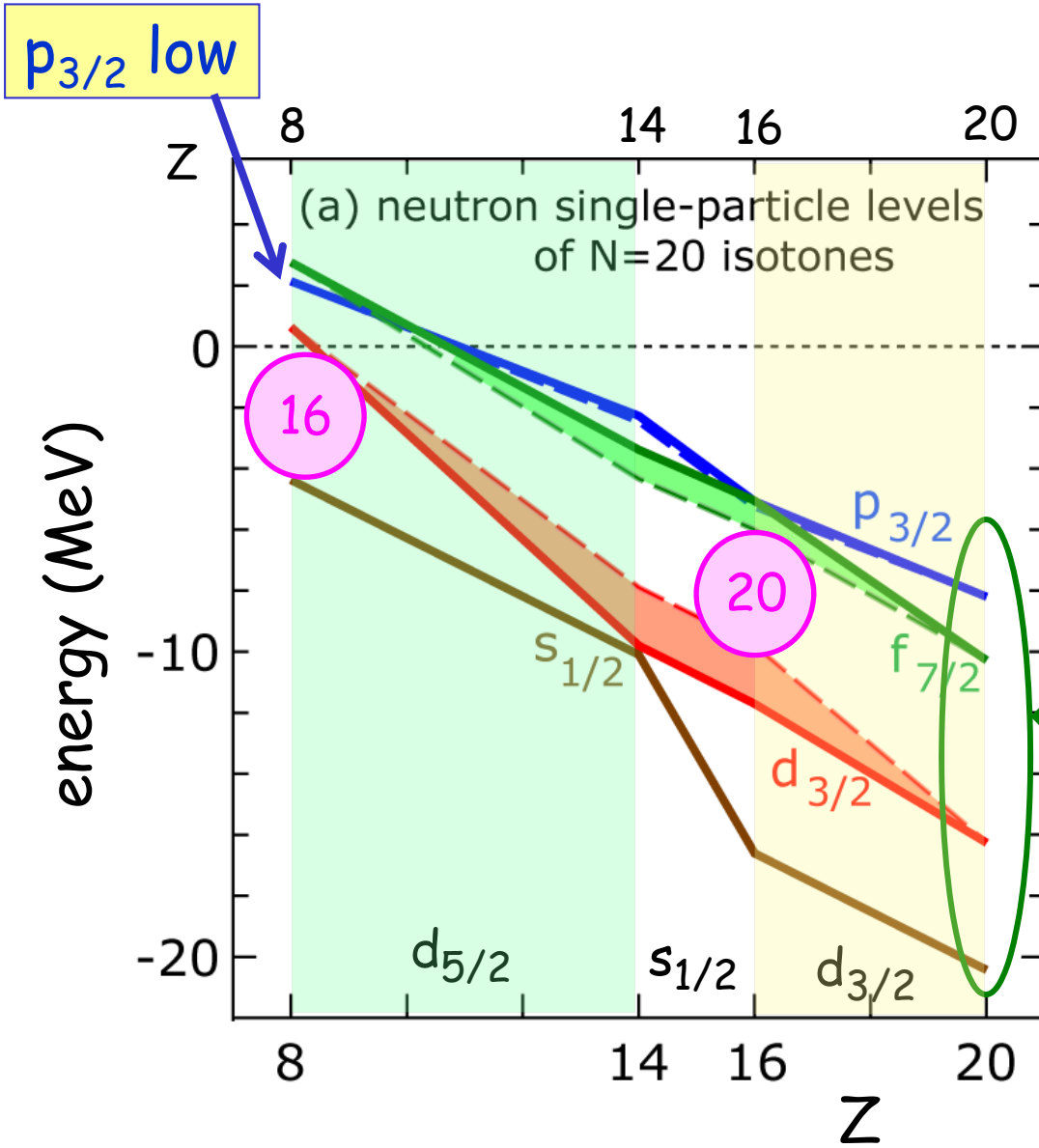


T=1 monopole interaction





Neutron single-particle energies at N=20 for Z=8~20



solid line : full
(central + tensor)

dashed line : central only

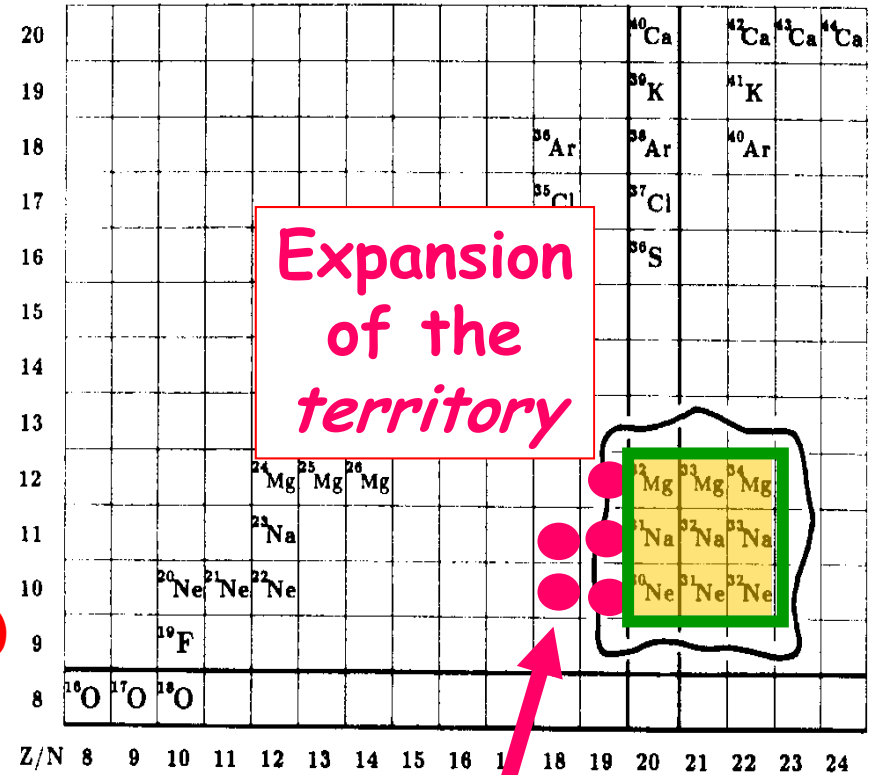
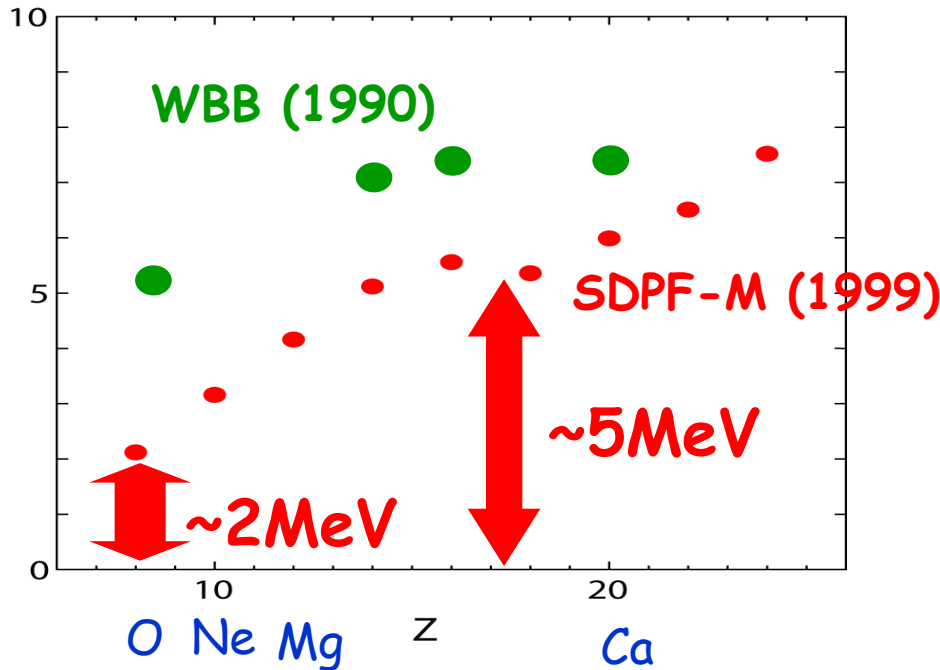
Tensor force makes
changes more dramatic.

These single-particle
energies are "normal"

$f_{7/2} - p_{3/2}$ 2~3 MeV
N=20 gap ~ 6 MeV

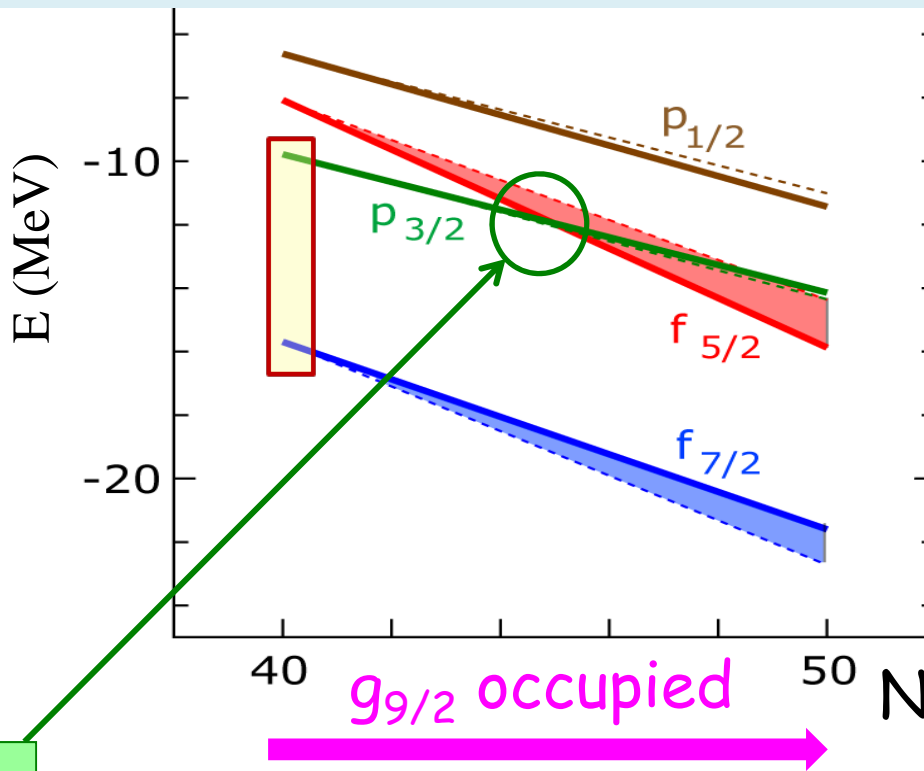
Monte Carlo Shell Model (MCSM) results have been obtained by the **SDPF-M** interaction for the full-*sd* + *f7/2* + *p3/2* space.

Effective N=20 gap between *sd* and *pf* shells



Neyens *et al.* 2005 Mg
 Tripathi *et al.* 2005 Na
 Dombradi *et al.* 2006 Ne
 Terry *et al.* 2007 Ne

Proton single-particles levels of Ni isotopes



Central Gaussian
+ Tensor

solid line:
full effect

dotted line:
central only

shaded area :
effect of
tensor force

From
Grawe,
EPJA25,
357

Crossing here
is consistent
with exp. on
Cu isotopes

PRL 103, 142501 (2009)

PHYSICAL REVIEW LETTERS

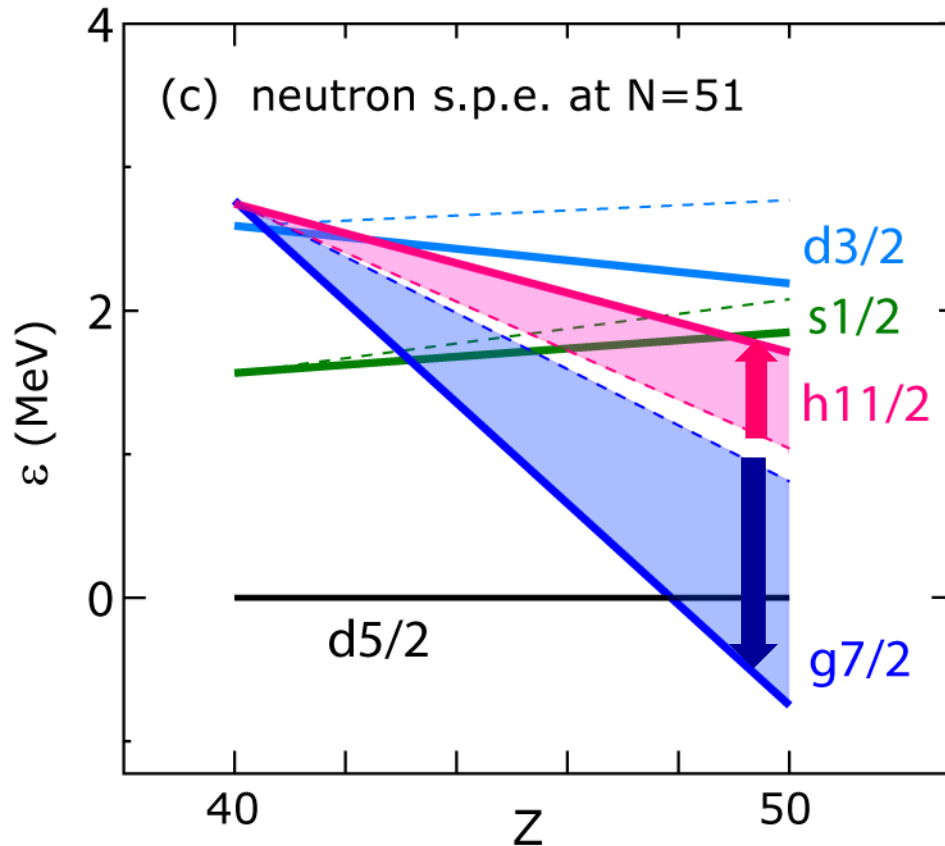
week ending
2 OCTOBER 2009

Nuclear Spins and Magnetic Moments of $^{71,73,75}\text{Cu}$: Inversion of $\pi 2p_{3/2}$ and $\pi 1f_{5/2}$ Levels in ^{75}Cu

K. T. Flanagan,^{1,2} P. Vingerhoets,¹ M. Avgoulea,¹ J. Billowes,³ M. L. Bissell,¹ K. Blaum,⁴ B. Cheal,³ M. De Rydt,¹ V. N. Fedosseev,⁵ D. H. Forest,⁶ Ch. Geppert,^{7,8} U. Köster,¹⁰ M. Kowalska,¹¹ J. Krämer,⁹ K. L. Kratz,⁹ A. Krieger,⁹ E. Mané,³ B. A. Marsh,⁵ T. Materna,¹⁰ L. Mathieu,¹² P. L. Molkanov,¹³ R. Neugart,⁹ G. Neyens,¹ W. Nörtershäuser,^{9,7} M. D. Seliverstov,^{13,16} O. Serot,¹² M. Schug,⁴ M. A. Sjoedin,¹⁷ J. R. Stone,^{14,15} N. J. Stone,^{14,15} H. H. Stroke,¹⁸ G. Tungate,⁶ D. T. Yordanov,⁴ and Yu. M. Volkov¹³

¹Instituut voor Kern- en Stralingsfysica, Katholieke Universiteit Leuven, B-3001 Leuven, Belgium

Shell structure of a key nucleus ^{100}Sn



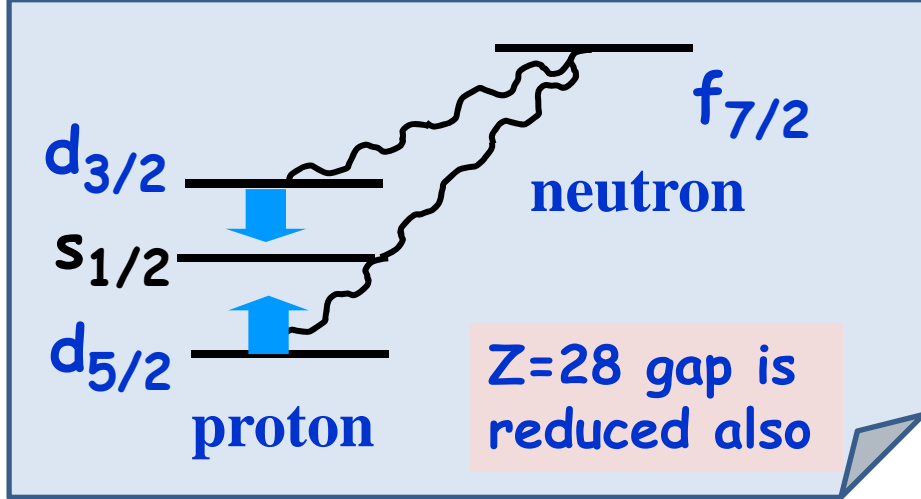
solid line : full
(central + tensor)

dashed line : central only

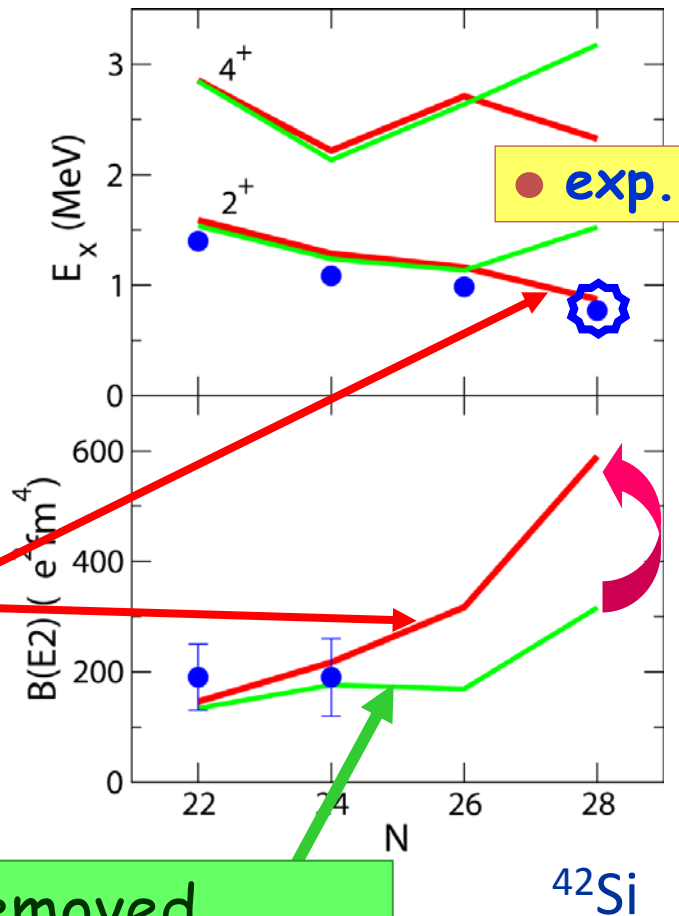
shaded area :
effect of tensor force

Exp. $d5/2$ and $g7/2$ should be close
Seweryniak et al.

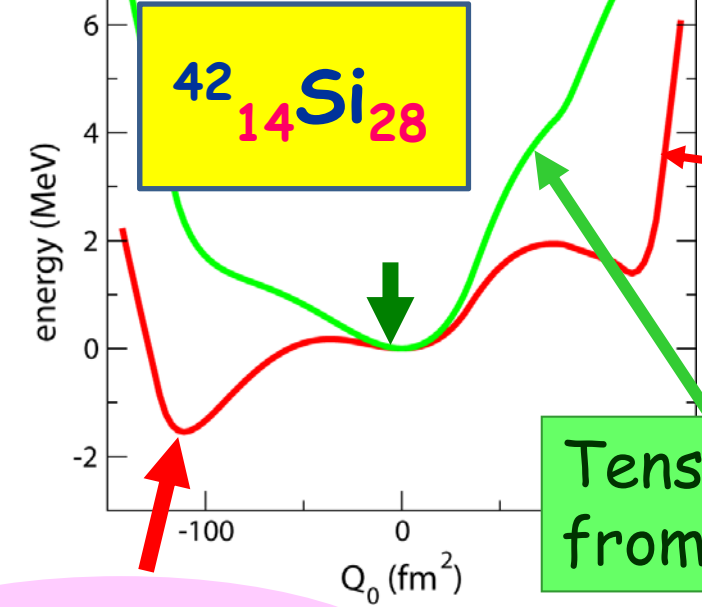
Phys. Rev. Lett. 99, 022504 (2007)
Gryzywacz et al.



Si isotopes
SM calc. by Utsuno et al.



Potential Energy Surface



full

Tensor force removed from cross-shell interaction

Strong oblate Deformation ?

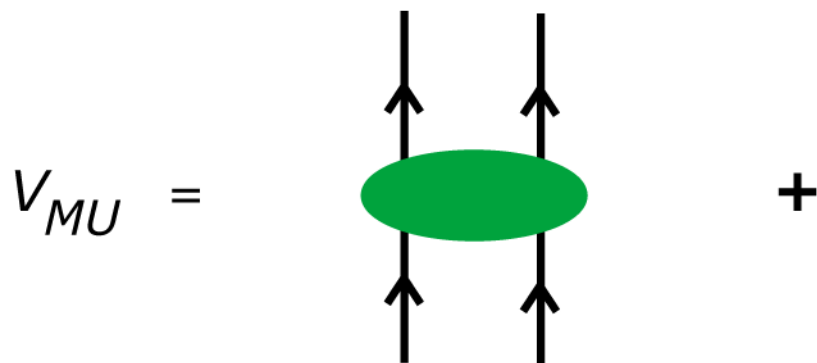
Other calculations show a variety of shapes.

Otsuka, Suzuki and Utsuno, Nucl. Phys. A805, 127c (2008)

^{42}Si : B. Bastin, S. Grévy et al., PRL 99 (2007) 022503

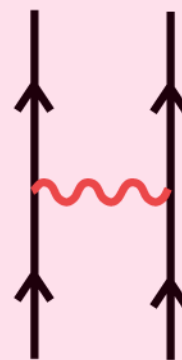
Systematic description of monopole properties of exotic nuclei can be obtained by an extremely simple interaction as

(a) central force :
Gaussian
(strongly renormalized)



+

(b) tensor force :
 $\pi + \rho$ meson
exchange



Can this be really the same as the bare tensor interaction ?

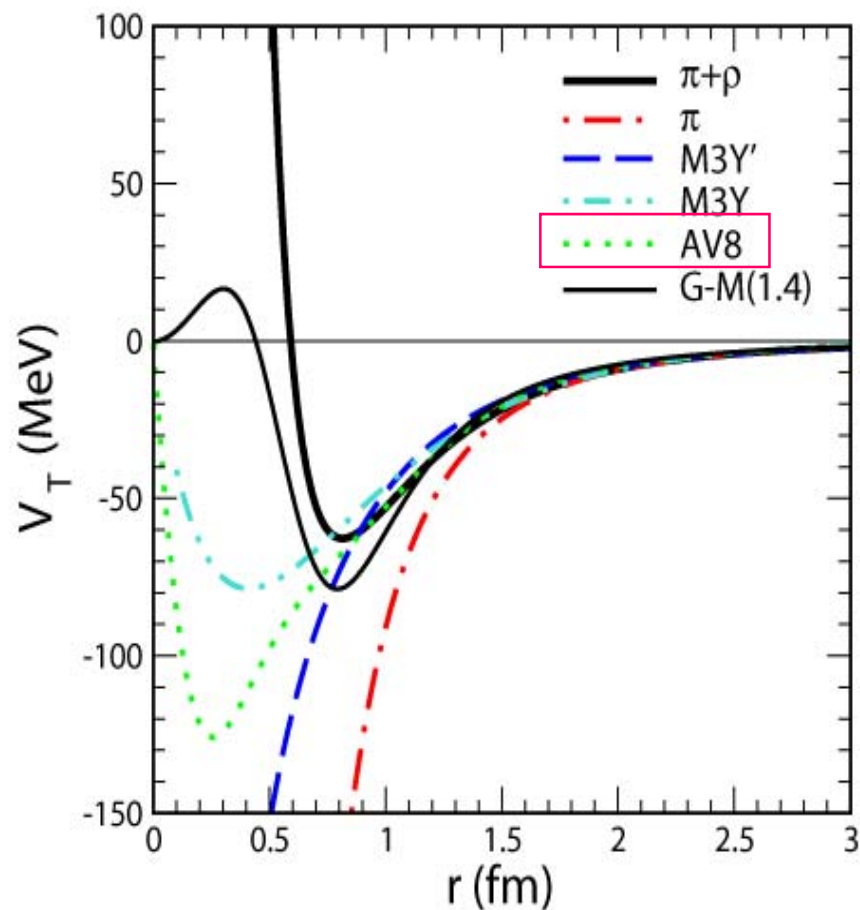
Question :

Can the tensor force survive after

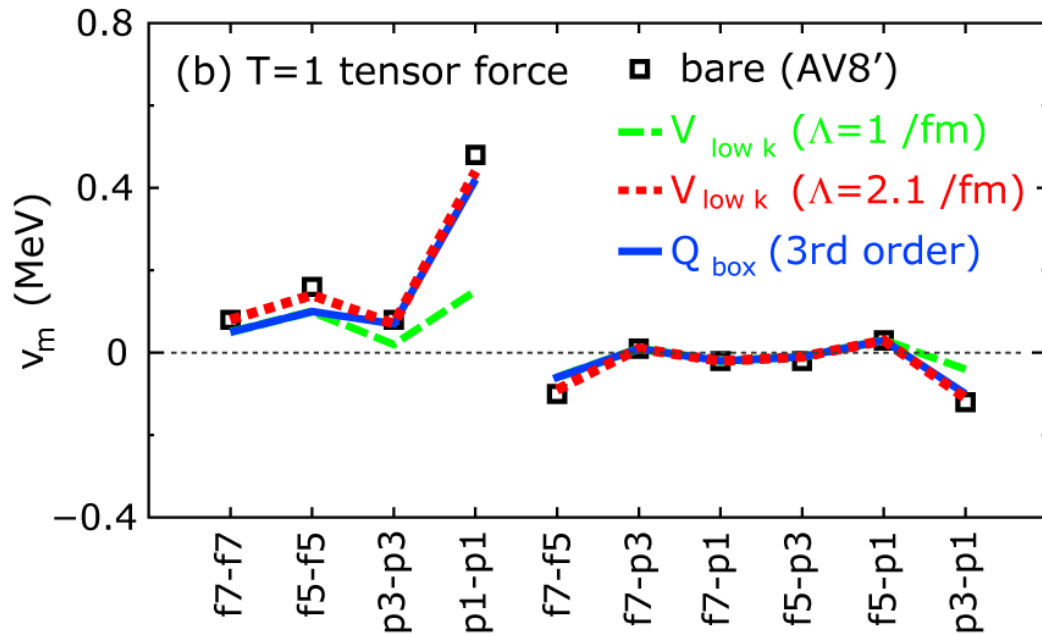
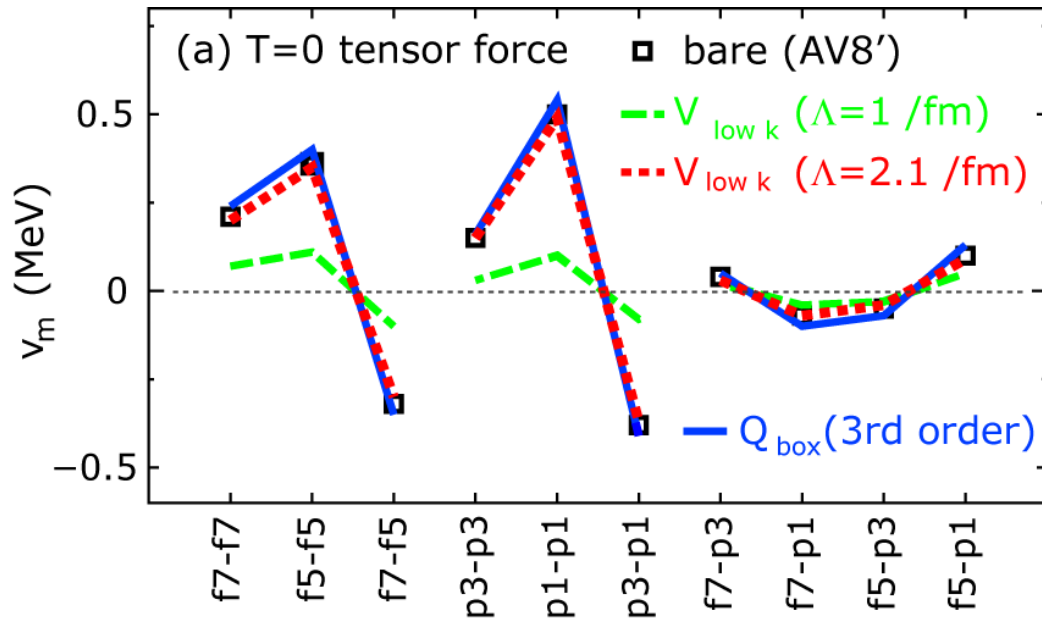
-treatment of short-range correlations (hard-core)

-treatment of core polarization

Tensor potential



Treatment of tensor force by $V_{\text{low } k}$ and Q box (3rd order)



Monopole component of **tensor** interactions in *pf* shell

Input: AV8'

Almost no cut-off (Λ) dependence, except for extremely low Λ

$V_{\text{low } k}$: Kuo, Bogner, et al.

Shell evolution in exotic nuclei due to tensor + central forces

PRL 104, 012501 (2010)

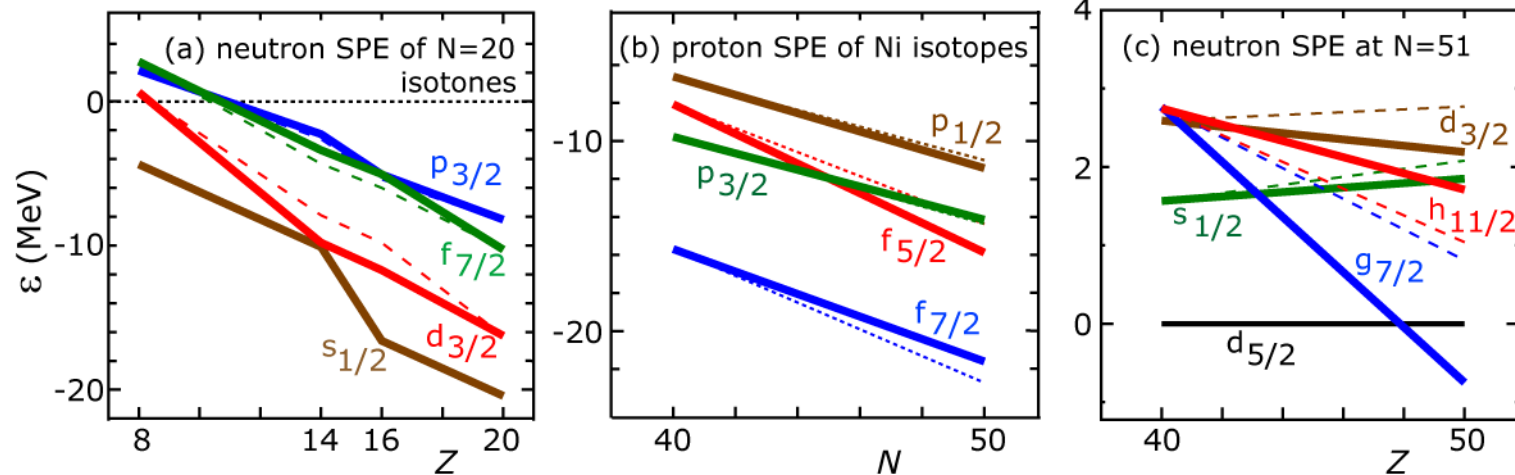
Selected for a Viewpoint in *Physics*
PHYSICAL REVIEW LETTERS

week ending
8 JANUARY 2010

Novel Features of Nuclear Forces and Shell Evolution in Exotic Nuclei

Takaharu Otsuka,^{1,2} Toshio Suzuki,³ Michio Honma,⁴ Yutaka Utsuno,⁵ Naofumi Tsunoda,¹
Koshiroh Tsukiyama,¹ and Morten Hjorth-Jensen⁶

Changes of single-particle properties due to these nuclear forces

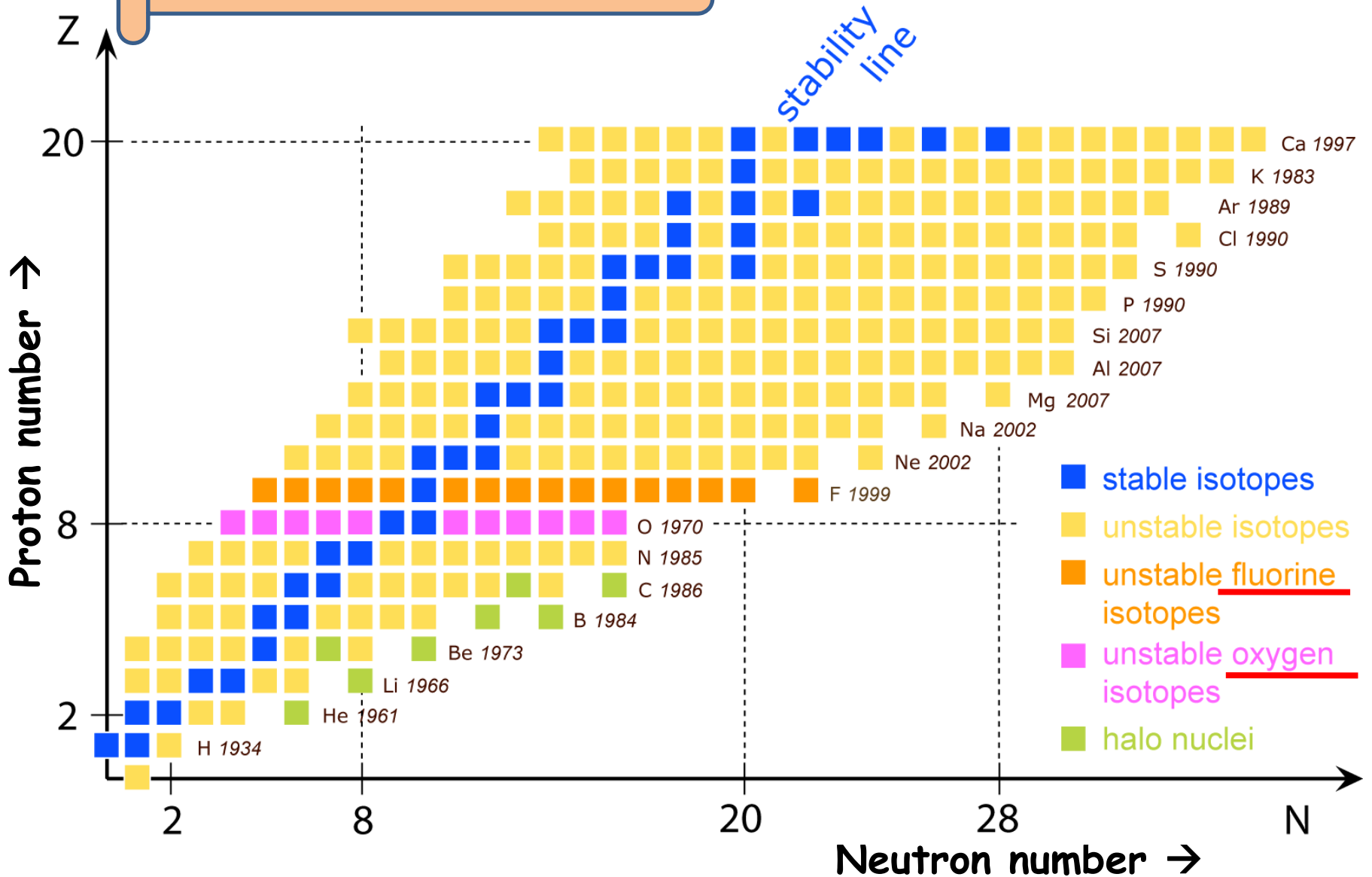


What is the next ?

Three-body force and exotic nuclei

Nuclear Chart - Left Lower Part -

Why is the drip line of Oxygen so near ?



Single-Particle Energy for Oxygen isotopes

by **microscopic** eff. int.

G-matrix+ core-pol. : Kuo, Brown

V_{low-k} : Bogner, Kuo, Schwenk

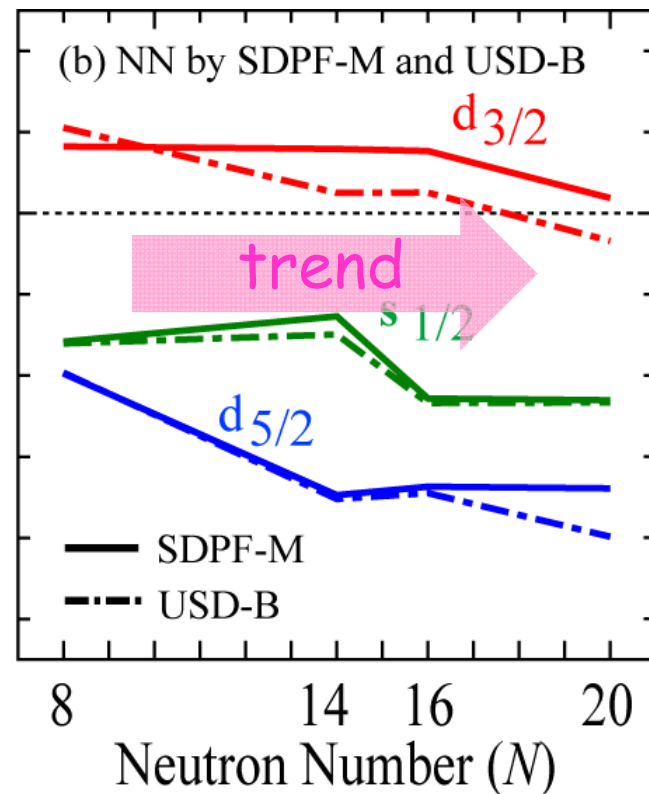
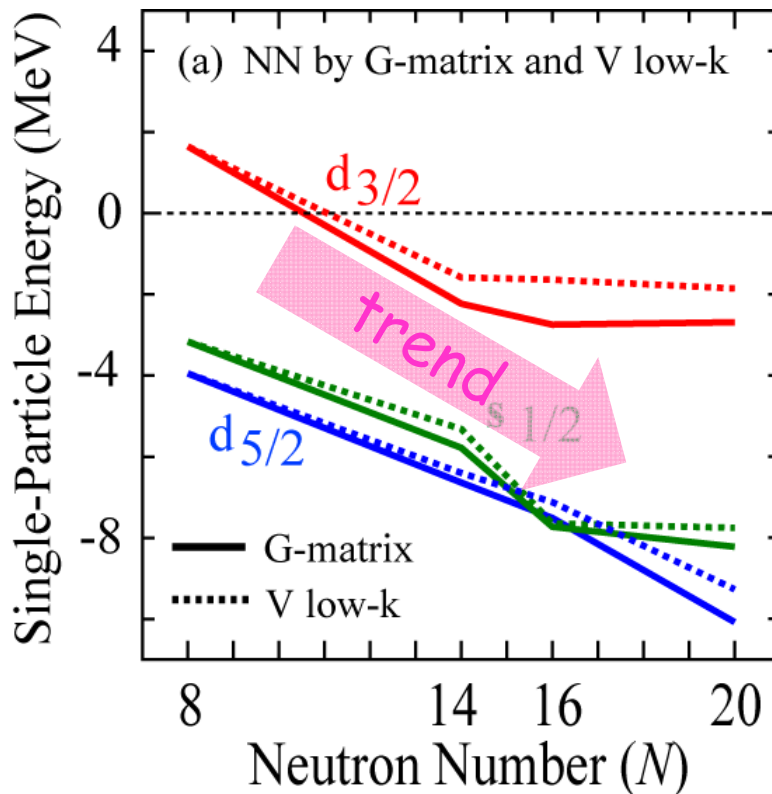
by **phenomenological** eff. int.
- **G-matrix + fit** -

SDPF-M

Utsuno, O., Mizusaki, Honma,
Phys. Rev. C **60**, 054315 (1999)

USD-B

Brown and Richter,
Phys. Rev. C **74**, 034315 (2006)



What is the origin of
the *repulsive modification* of
 $T=1$ monopole matrix elements ?

The same puzzle as in the pf shell

A solution within *bare* 2-body interaction
is very unlikely
(considering efforts made so far)

Zuker, Phys. Rev. Lett. 90, 042502 (2003)

→ 3-body interaction

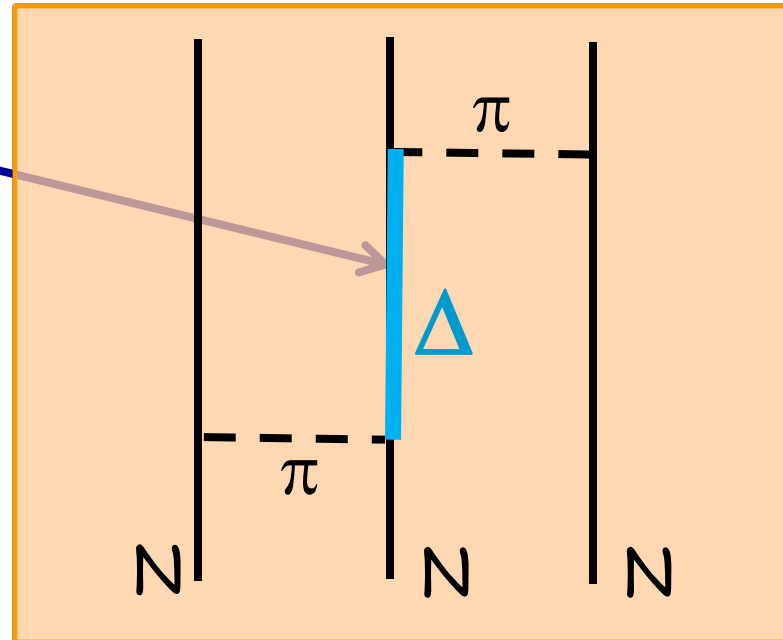
The clue : Fujita-Miyazawa 3N mechanism (Δ -hole excitation)

Progress of Theoretical Physics, Vol. 17, No. 3, March 1957

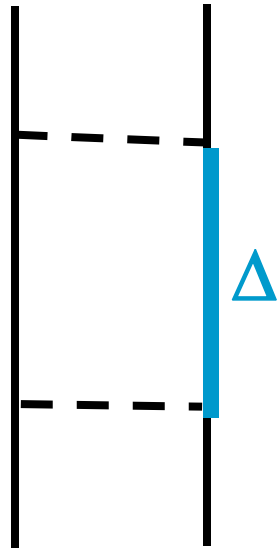
Pion Theory of Three-Body Forces

Jun-ichi FUJITA and Hironari MIYAZAWA

Δ particle
 $m=1232$ MeV
 $S=3/2, I=3/2$



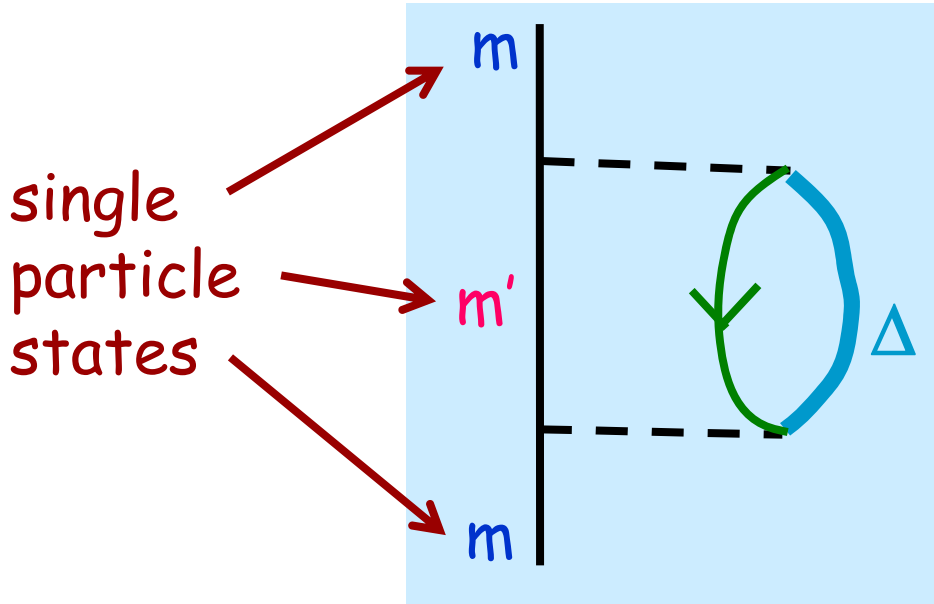
Renormalization of NN interaction due to Δ excitation in the intermediate state



Modification to
bare NN interaction
(for NN scattering)

$T=1$
attraction
between NN
effectively

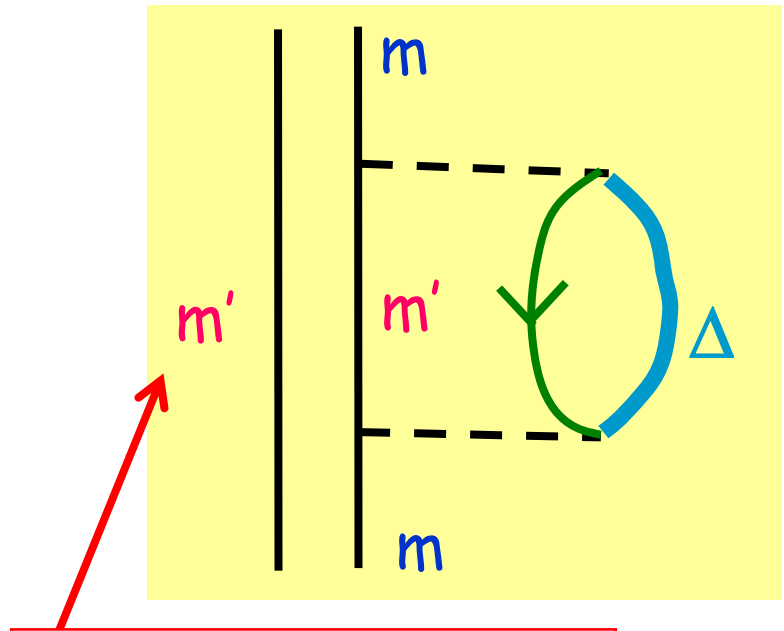
Pauli blocking effect on the renormalization of single-particle energy



Renormalization of single particle energy due to

Δ -hole excitation

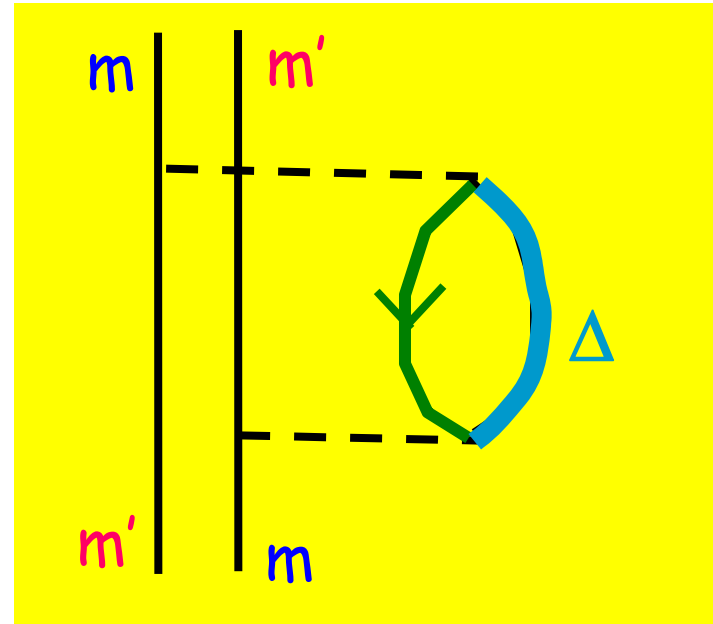
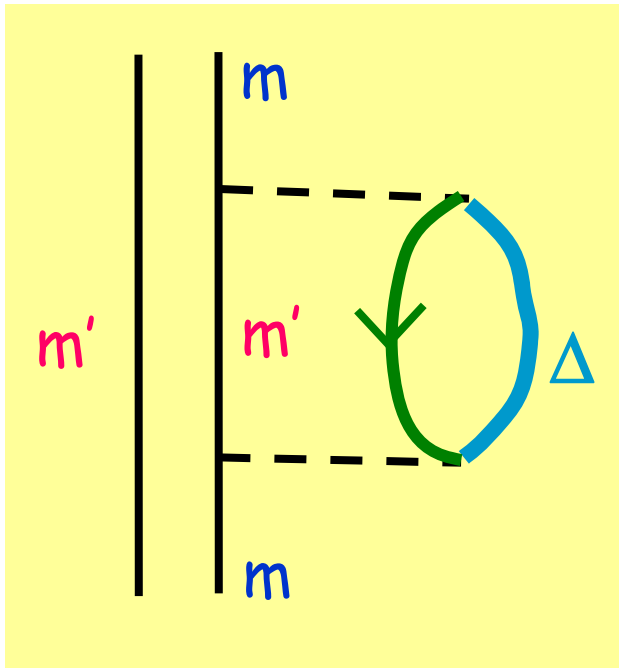
→ more binding (attractive)



Another valence particle in state m'

Pauli Forbidden
→ *The effect is suppressed*

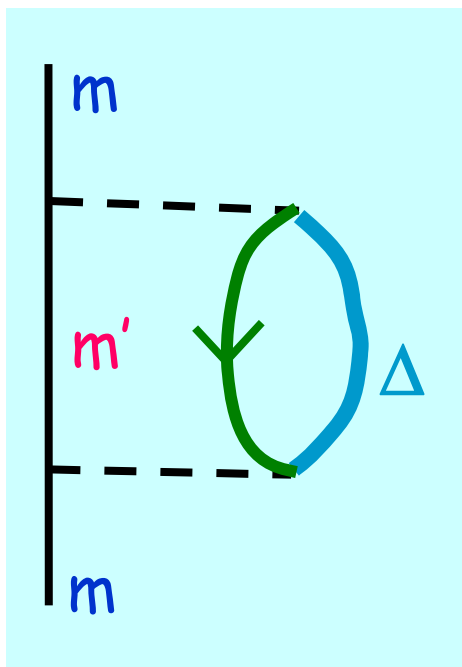
Inclusion of Pauli blocking



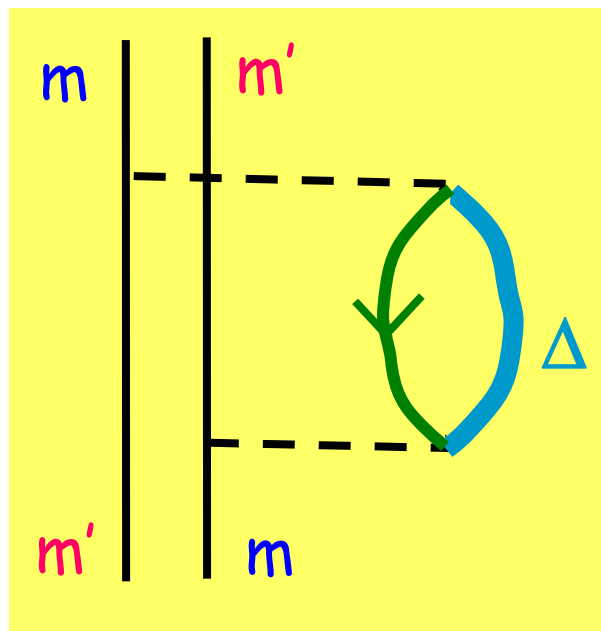
Pauli forbidden
(from previous page)

This Pauli effect is
included automatically
by the exchange term.

Most important message with Fujita-Miyazawa 3NF



+



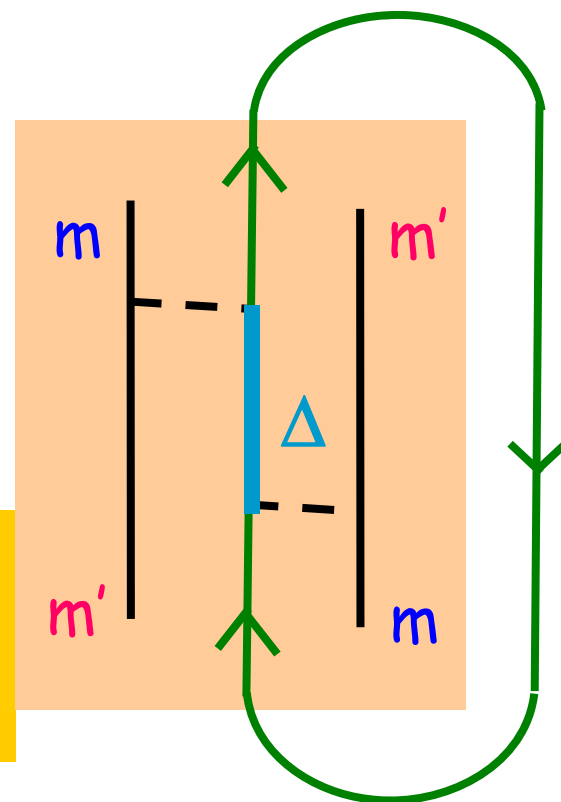
Pauli blocking

same



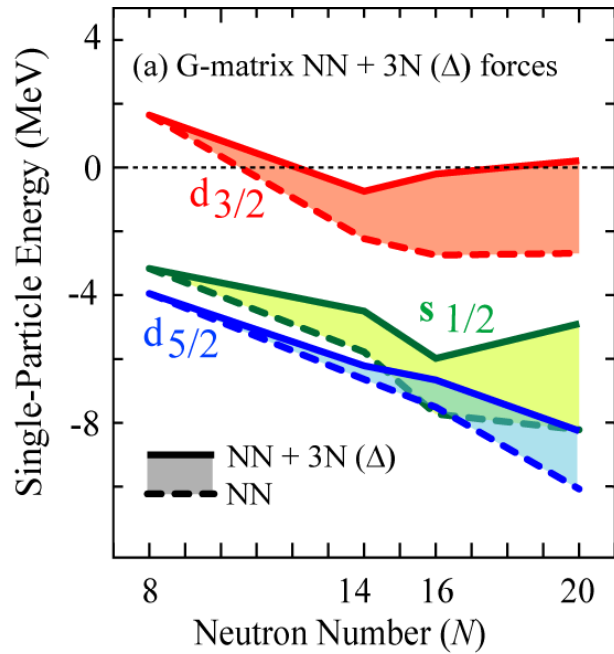
Monopole part of Fujita-Miyazawa 3-body force

Effective monopole repulsive interaction

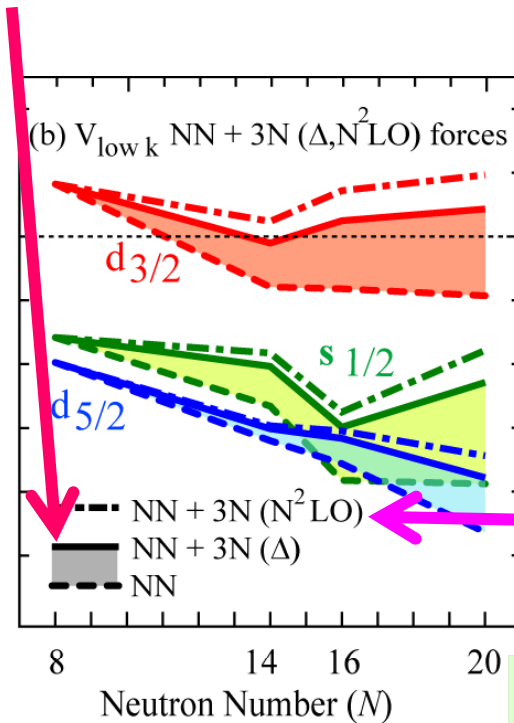


Renormalization of single particle energy

(i) Δ -hole excitation in a conventional way



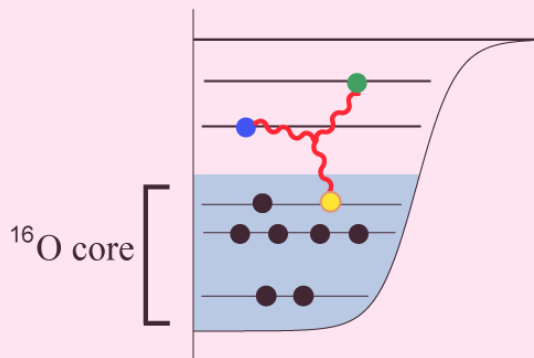
(ii) EFT with Δ



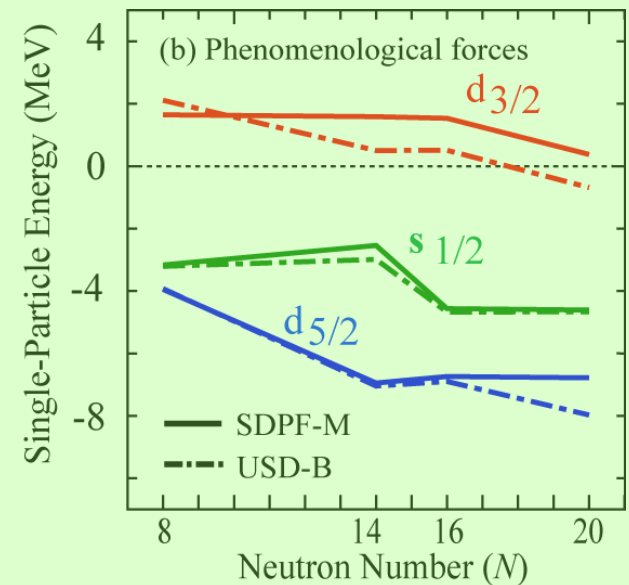
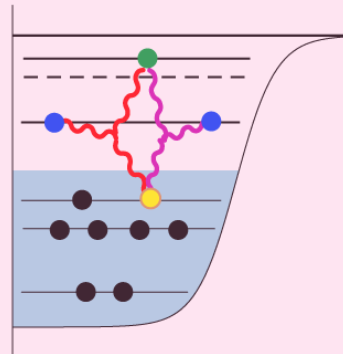
Δ -hole dominant role in determining oxygen drip line

(iii) EFT incl. contact terms ($N^2\text{LO}$)

(c) 3-body interaction



(d) 3-body interaction with one more neutron added to (c)



Conventional calculation with $\pi N\Delta$ coupling

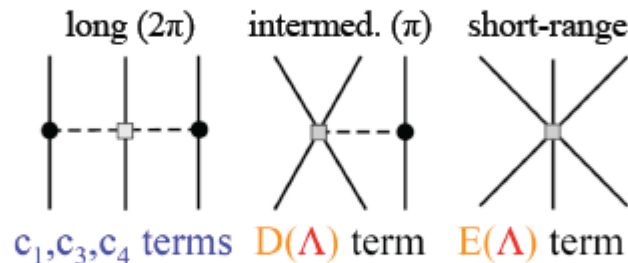
π exchange with radial cut-off at 0.5 fm , $\Delta E = 293$ MeV

$$f_{\{\pi N\Delta\}}/f_{\{\pi NN\}} = \sqrt{9/2}$$

A.M. Green, Rep. Prog. Phys. 39, 1109 (1976)

Low-momentum 3N interactions

from leading N^2LO chiral EFT $\sim (Q/\Lambda)^3$ van Kolck (1994), Epelbaum et al. (2002)



c_1 from πN , consistent with NN

Meissner (2007)

$$c_1 = -0.9^{+0.2}_{-0.5}, c_3 = -4.7^{+1.2}_{-1.0}, c_4 = 3.5^{+0.5}_{-0.2}$$

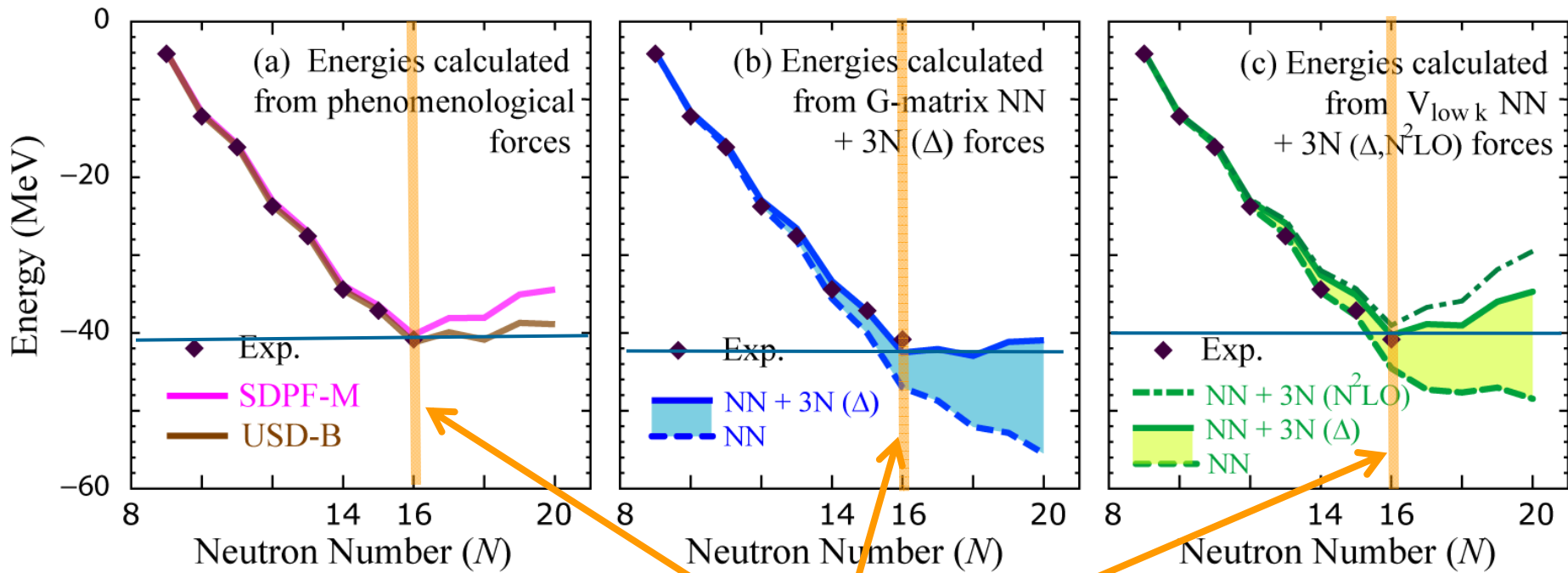
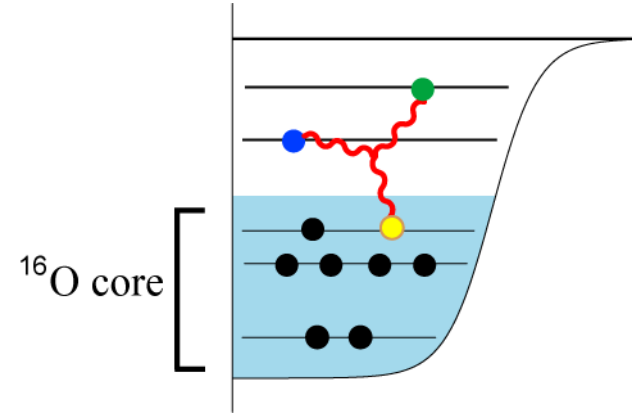
c_3, c_4 important for structure, large uncertainties at present

NN for smooth cutoff $V_{low k}$ ($n_{exp}=4$) from $N^3LO(500)$

D, E terms fitted to E(3H) and radius(4He)

Ground-state energies of oxygen isotopes

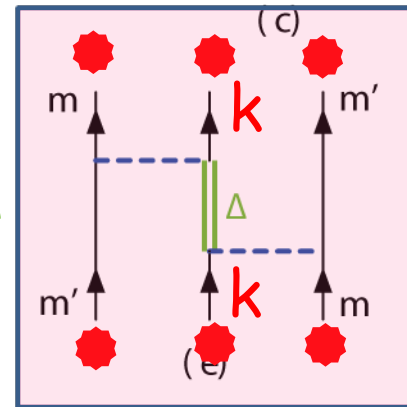
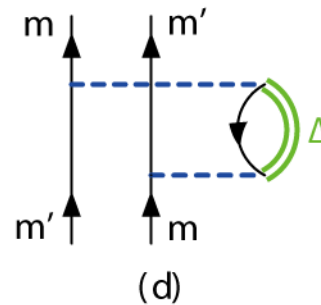
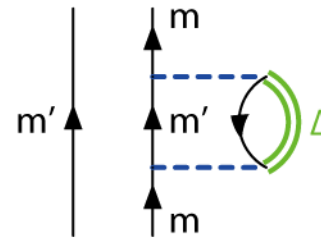
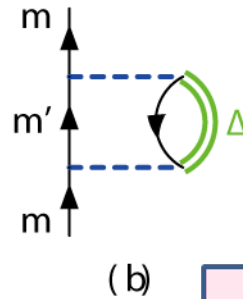
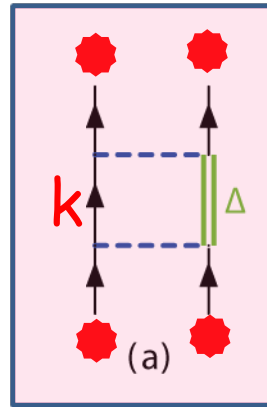
NN force + $3N$ -induced NN force
(Fujita-Miyazawa force)



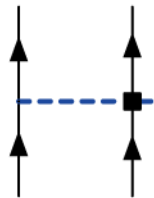
Drip line

For neutron matter: \bullet k states below Fermi level

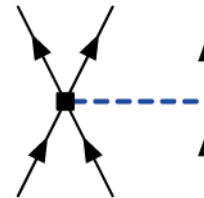
attractive



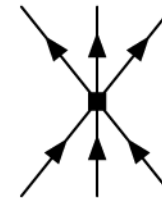
repulsive



(f)



(g)



(h)

Brown and Green, Nucl.Phys. A137, 1 (1969)

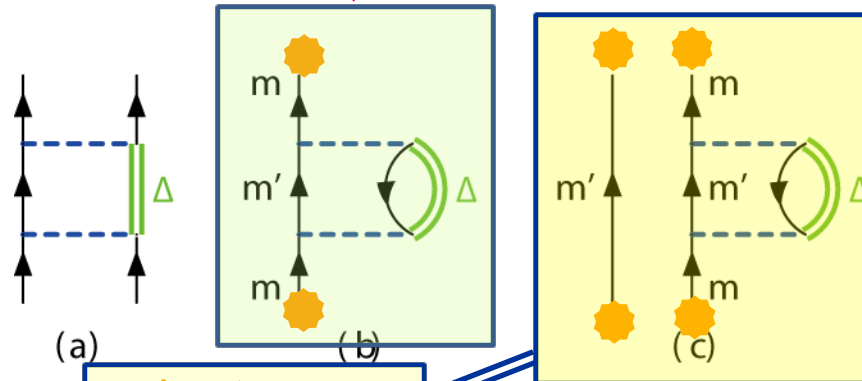
Fritsch, Kaiser and Weise, Nucl. Phys. A750, 259 (2005);


Tolos, Friman and Schwenk, Nucl.Phys. A806, 105 (2008);

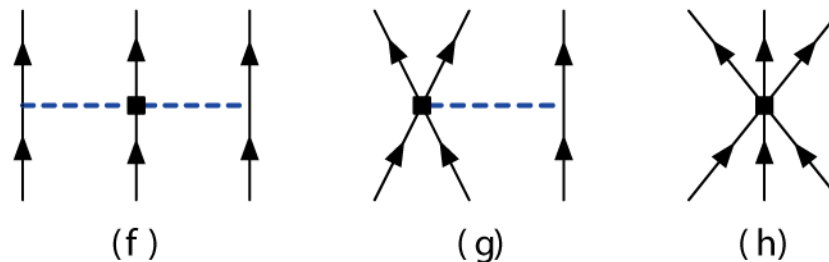
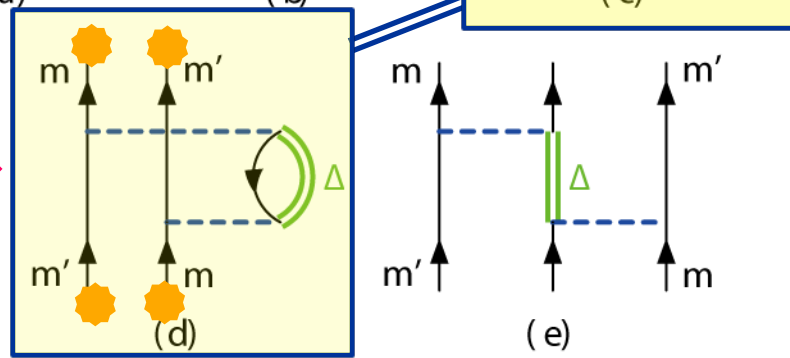
Hebeler and Schwenk, arXiv:0911.0483 [nucl-th]

For valence neutrons:  states outside the core

Attractive (single-particle energy renormalization)

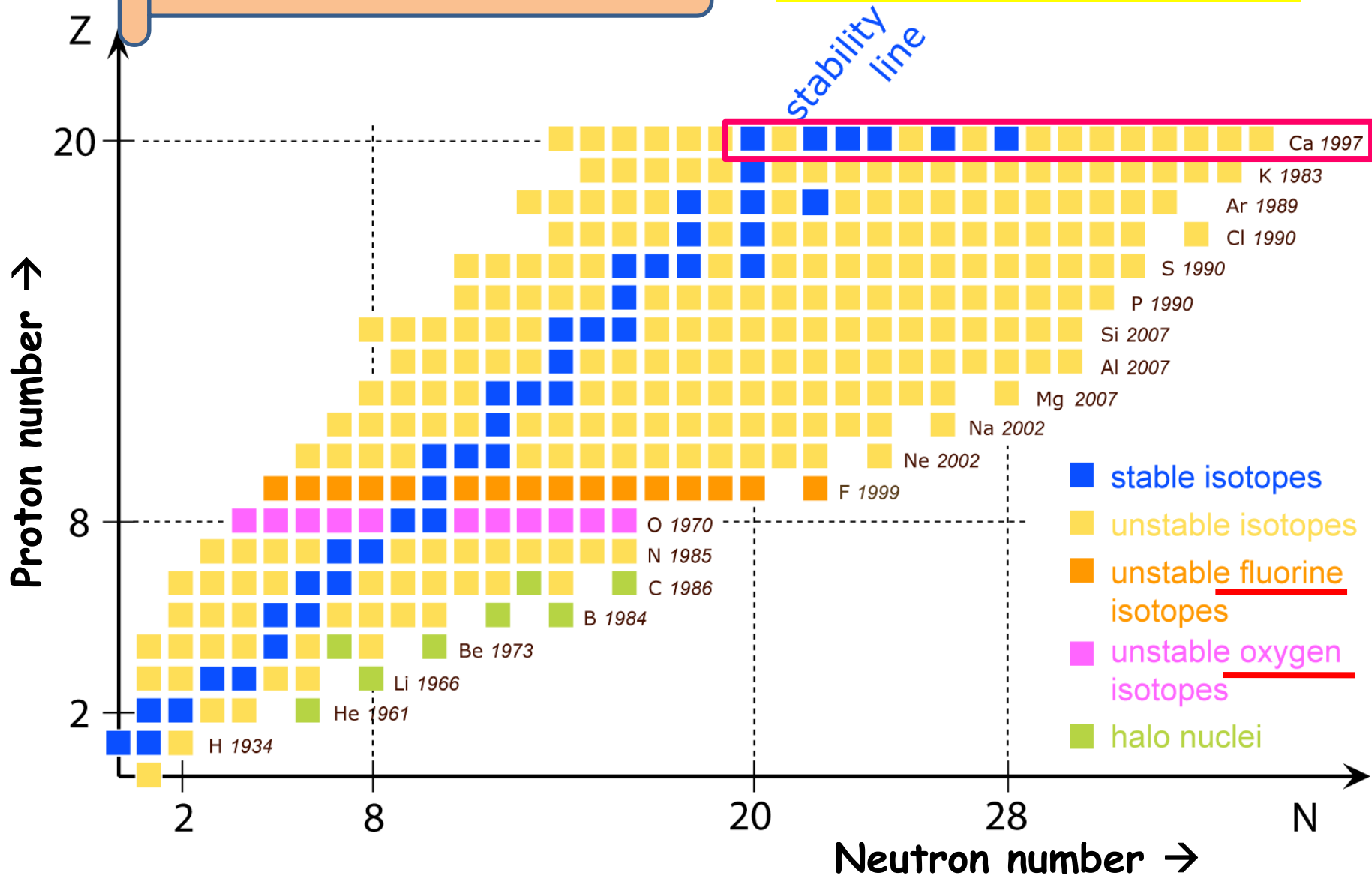


repulsive 
(valence neutron interaction)

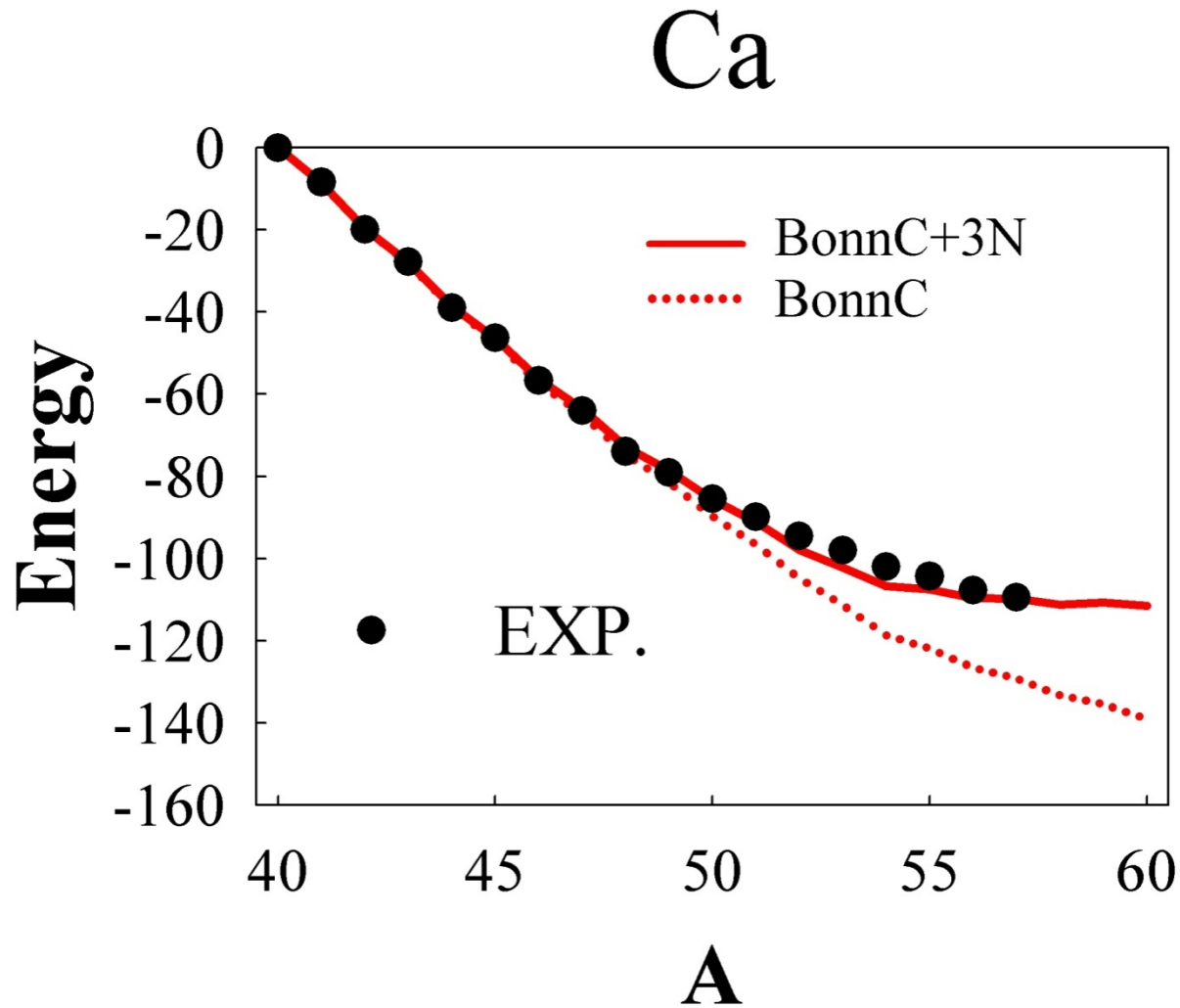


Nuclear Chart - Left Lower Part -

Ca isotopes



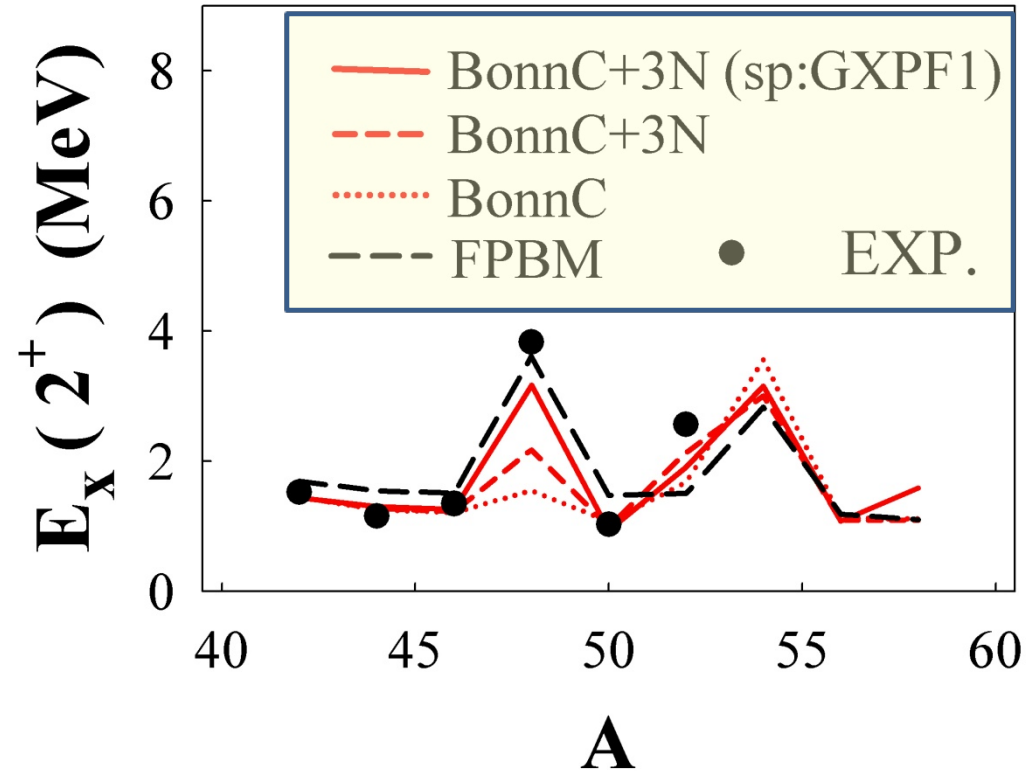
Ca ground-state energy *cont'd*



SPE : GXPF1 f7:-8.62 f5: -1.38 p3: -5.68 p1: -4.14

Ca 2^+ level systematics

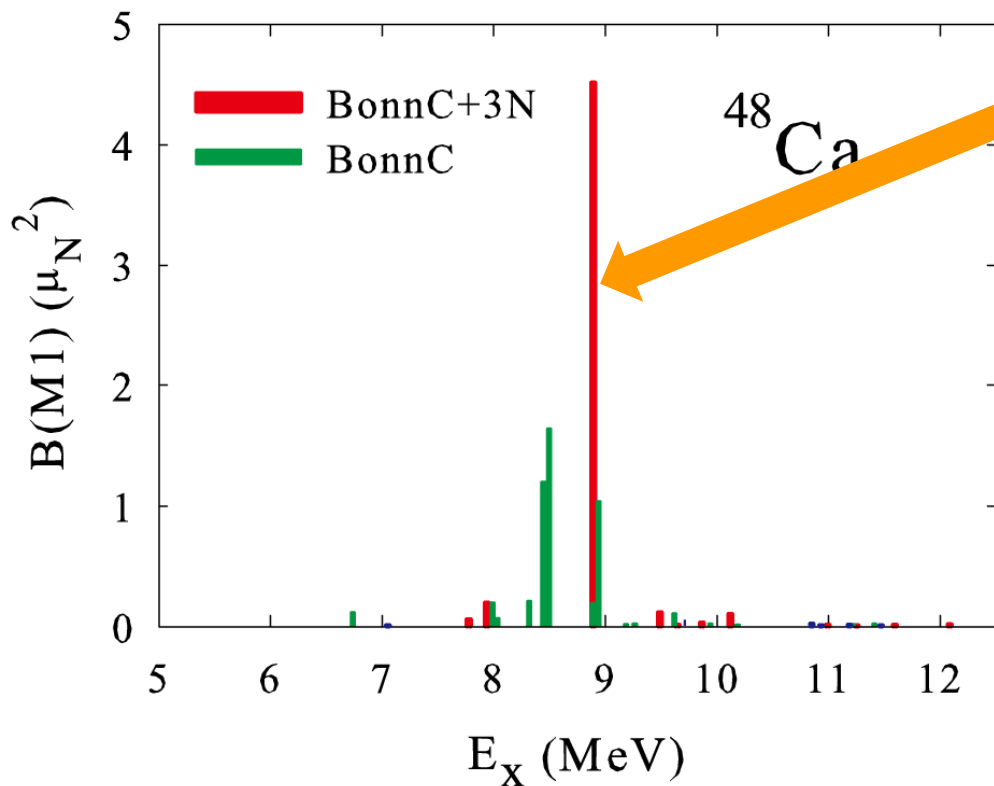
Ca



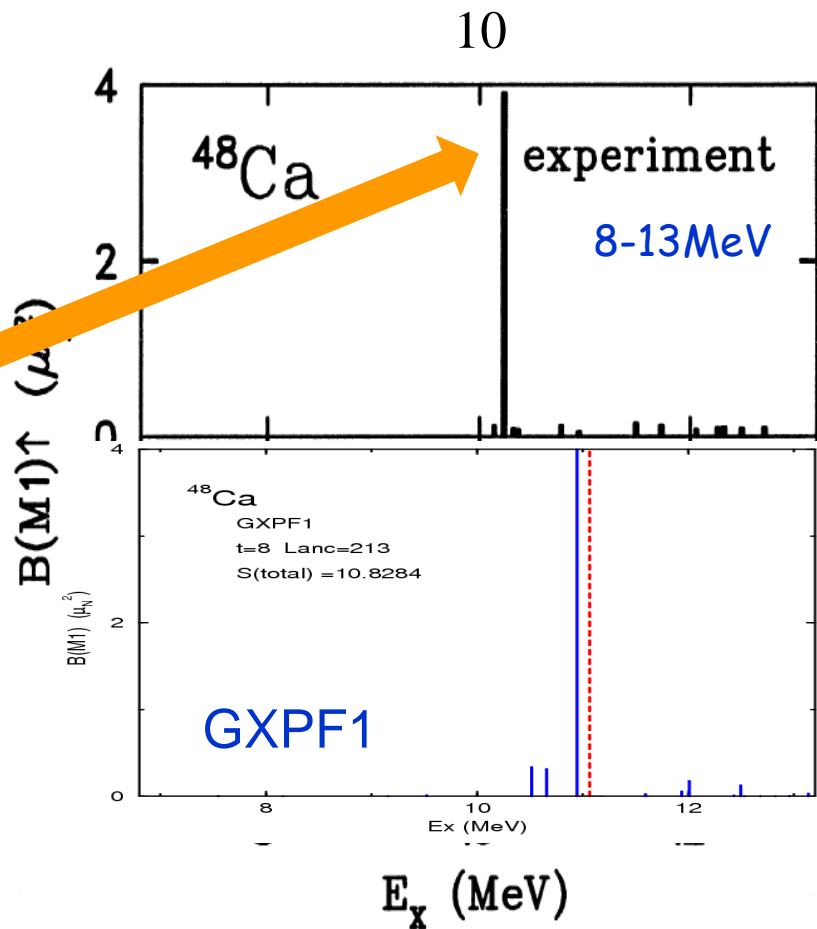
2^+ of ^{48}Ca rises by 3N
becomes about right by using GXPF1A SPE

N=32, 34 higher 2^+ levels

^{48}Ca M1 excitation



spe : GXPF1



Spin quenching factor 0.8

Summary

1. 核力のモノポール成分の効果は粒子数とともにリニアに増大
レンズのようなもので、中性子過剰核で核力を究めることを可能にする
2. メイヤー・イェンゼンの“原子核によらない普遍的魔法数・殻構造”のパラダイムは
エキゾチック核(不安定核)では崩れる。2体、3体の核力はその主な原動力。

殻構造は一粒子の性質を決めるだけでなく、Jahn-Teller効果に他ならない
基底状態近辺の変形を決め、スピンに関わる性質にも影響

3. テンソル力と中心力はワインバーグ型の有効力を構成。テンソル力は原子核中
でも bare と同じ (様々な現代風核力で同様、もちろん現代的理論)。

陽子-中性子間の相互作用をよく表す

N=20近傍の island of inversion、 ^{78}Ni , ^{100}Sn などの key issue を解く

4. 藤田-宮沢3体力は、バレンス中性子間に有効斥力を生じる。

中性子数を増やすと一粒子エネルギーの間は広がる

新たな魔法数の発生、shell quenching 仮説とは逆のこと

酸素のドリップラインの説明、殻模型有効相互作用の最後の謎を解明？

カルシウムなどで、ゆるく束縛された中性子物質を形成か？

5. 核力による多体問題が核物理であるならば、本当の核物理はこれから始まると
言っても過言でない (安定核はプロローグ？)

Summary

1. Effects of monopole component increase linearly as the particle number. This works like a lens, which enables us to study nuclear forces deeply in neutron-rich nuclei.
2. Meyer-Jensen's paradigm that magic number and shell structure remain (topologically) unchanged for all nuclei has been destroyed for exotic nuclei. 2- and 3-body forces are driving forces.

Shell structure determines not only single-particle properties but also deformations, which are nothing but Jahn-Teller effect, and spin properties as well.

3. Tensor force and central force constitute an effective force of Weinberg type. Tensor force appears, in its monopole channel, to be the same between in nuclei and in free space (as proved by modern theories).

Summary (continued)

Proton-neutron forces are well modeled.

N=20 island of inversion, key issues like ^{78}Ni , ^{100}Sn

4. Fujita-Miyazawa force generates effective repulsive forces between valence neutrons.

The spacing between single-particle levels gets wider as the neutron number increases.

new magic number; opposite to shell quenching

Drip line of oxygen

Last puzzle of SM eff. interaction solved ?

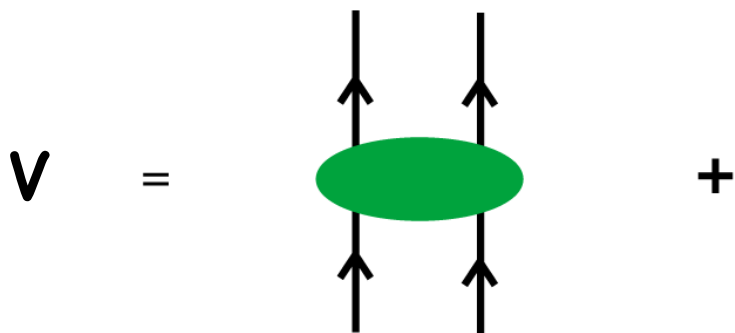
Loosely bound neutron matter in Ca isotopes ?

5. If nuclear physics implies many-body problems due to nuclear forces, it may be said that the real nuclear physics starts now (stable nuclei are prolog ?)

Summary-2

Dominant monopole forces are due to

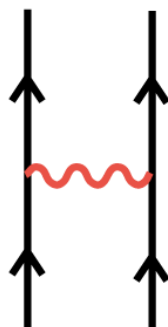
(a) central force :
Gaussian
(strongly renormalized)



$\mu = 1 \text{ fm}$

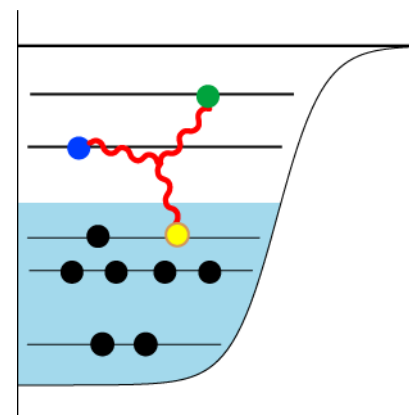
basic binding

(b) tensor force :
 $\pi + \rho$ meson
exchange



variation of
shell structure

FM 3NF

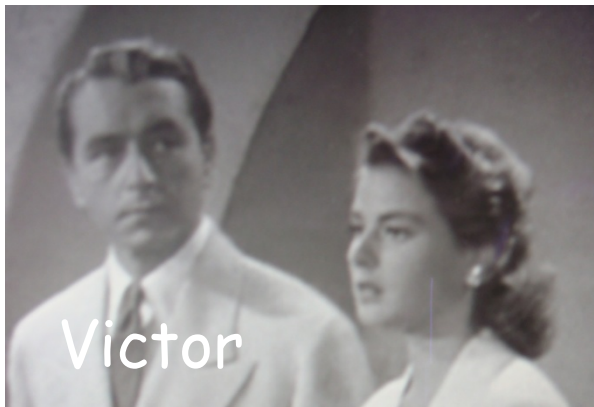


limit of
existence

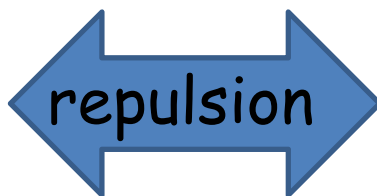
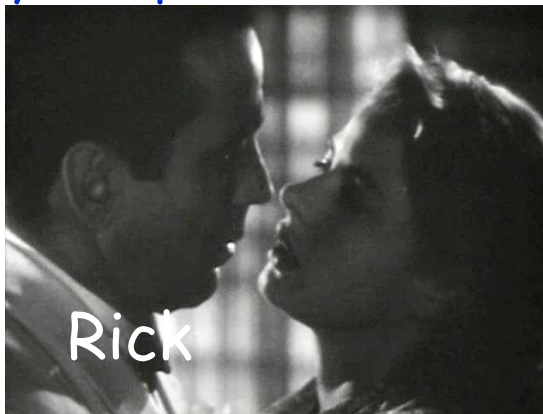
Casablanca mechanism

Love = attractive force

This love is reduced
by the presence of Rick



This love is reduced
by the presence of Victor



Collaborators

T. Suzuki	Nihon U.
M. Honma	Aizu
Y. Utsuno	JAEA
N. Tsunoda	Tokyo
K. Tsukiyama	Tokyo
M. H.-Jensen	Oslo
A. Schwenk	TRIUMF/Darmstadt
J. Holt	ORNL
K. Akaishi	RIKEN