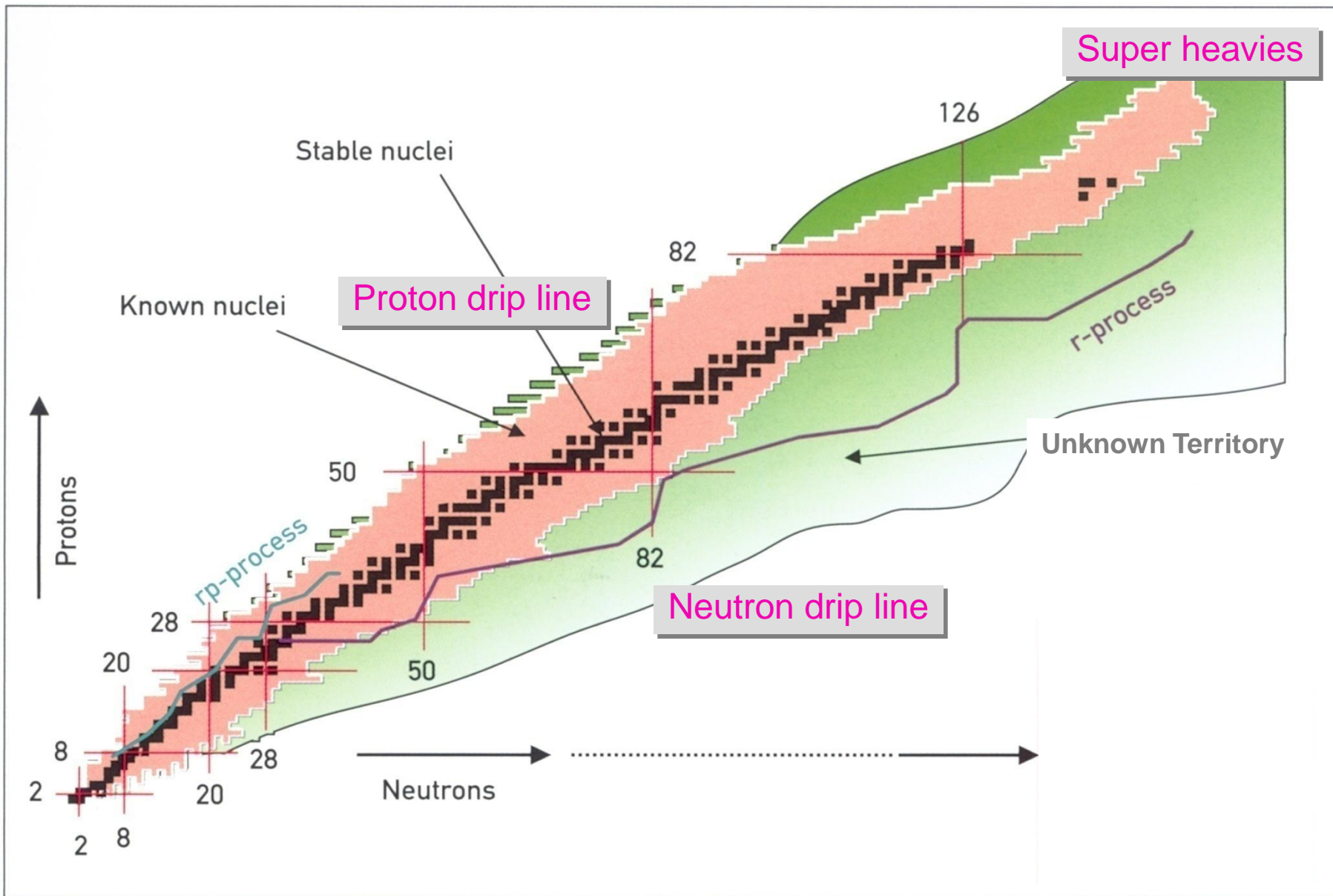
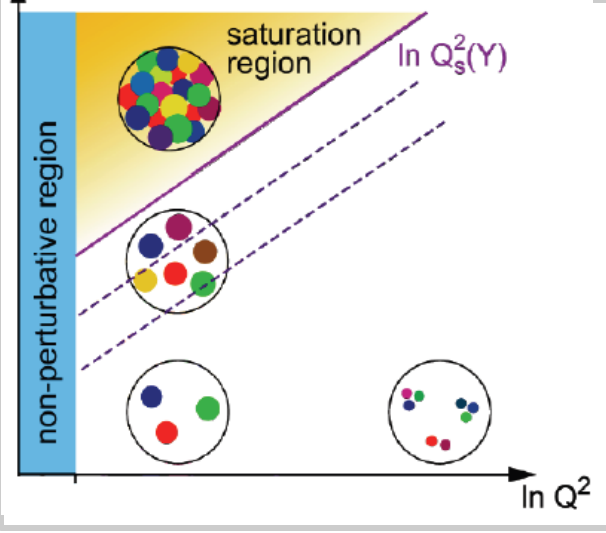
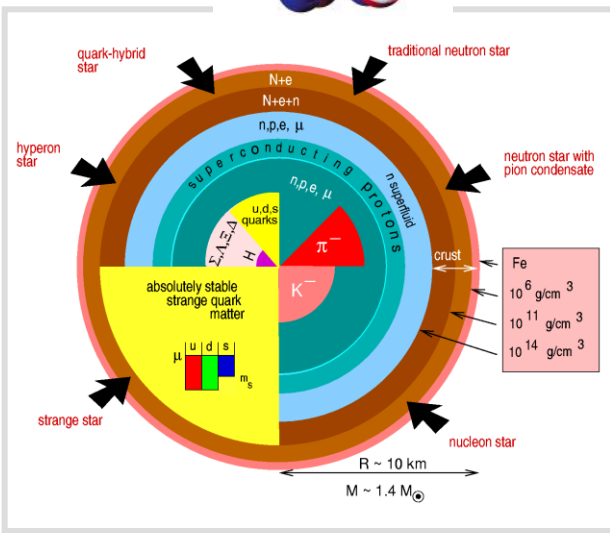
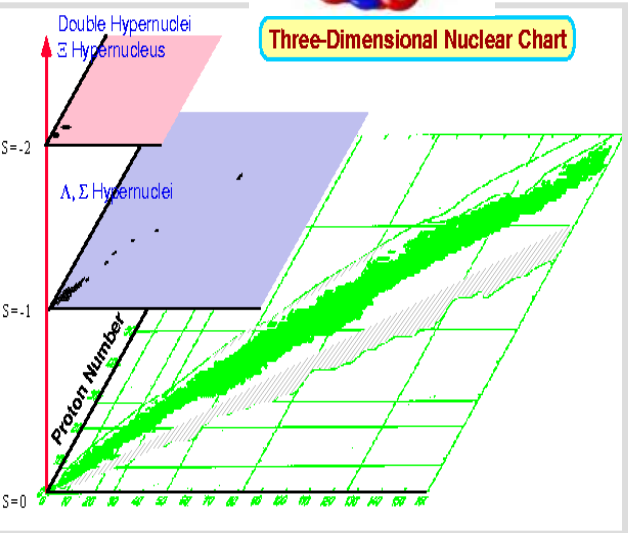
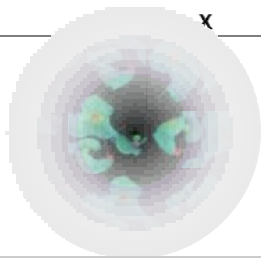
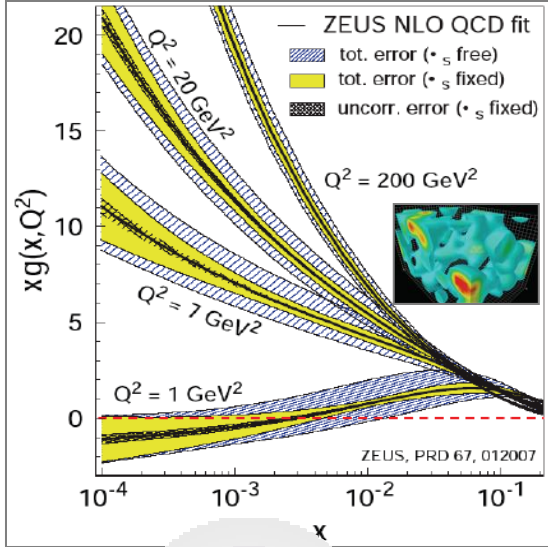
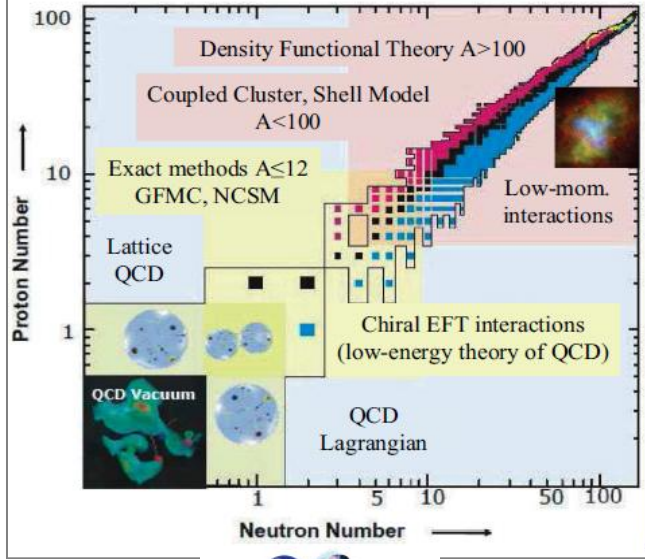
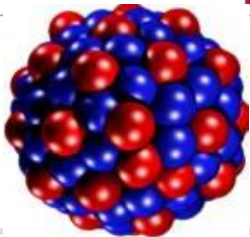
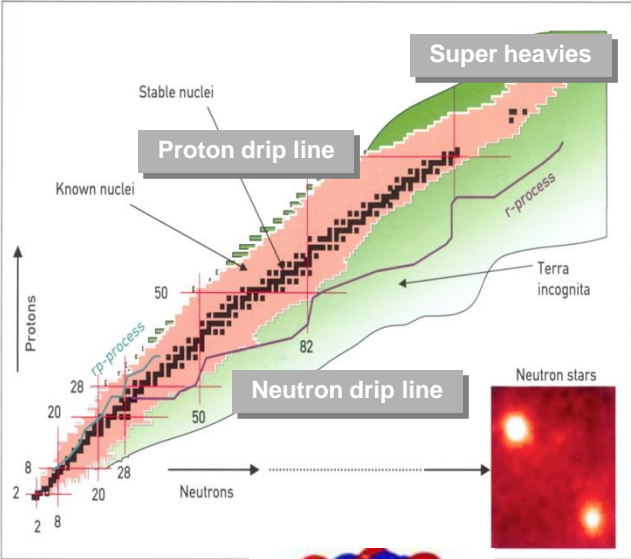


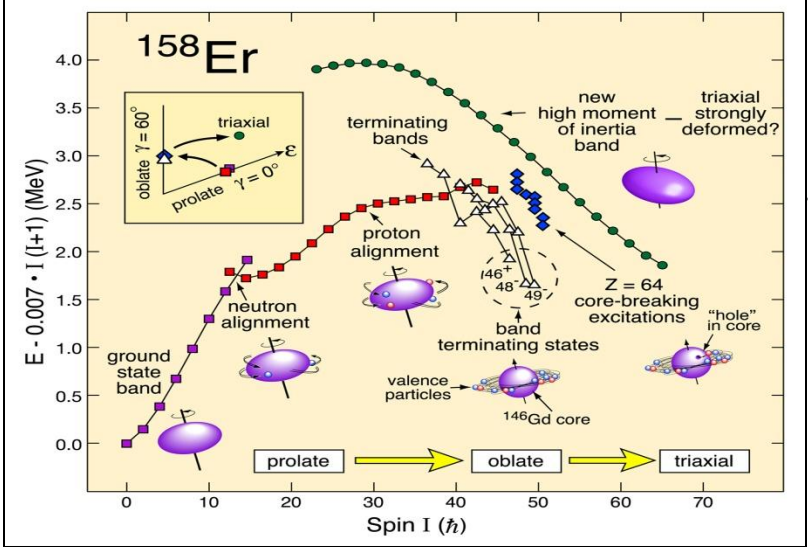
An aerial photograph of a vast, rugged canyon system. The landscape is characterized by deep, layered rock formations and a network of canyons. A prominent river winds through the center of the canyon, its path highlighted by a bright blue line. The overall scene is one of extreme geological complexity and scale.

Towards the Extremes of the Nuclear Landscape

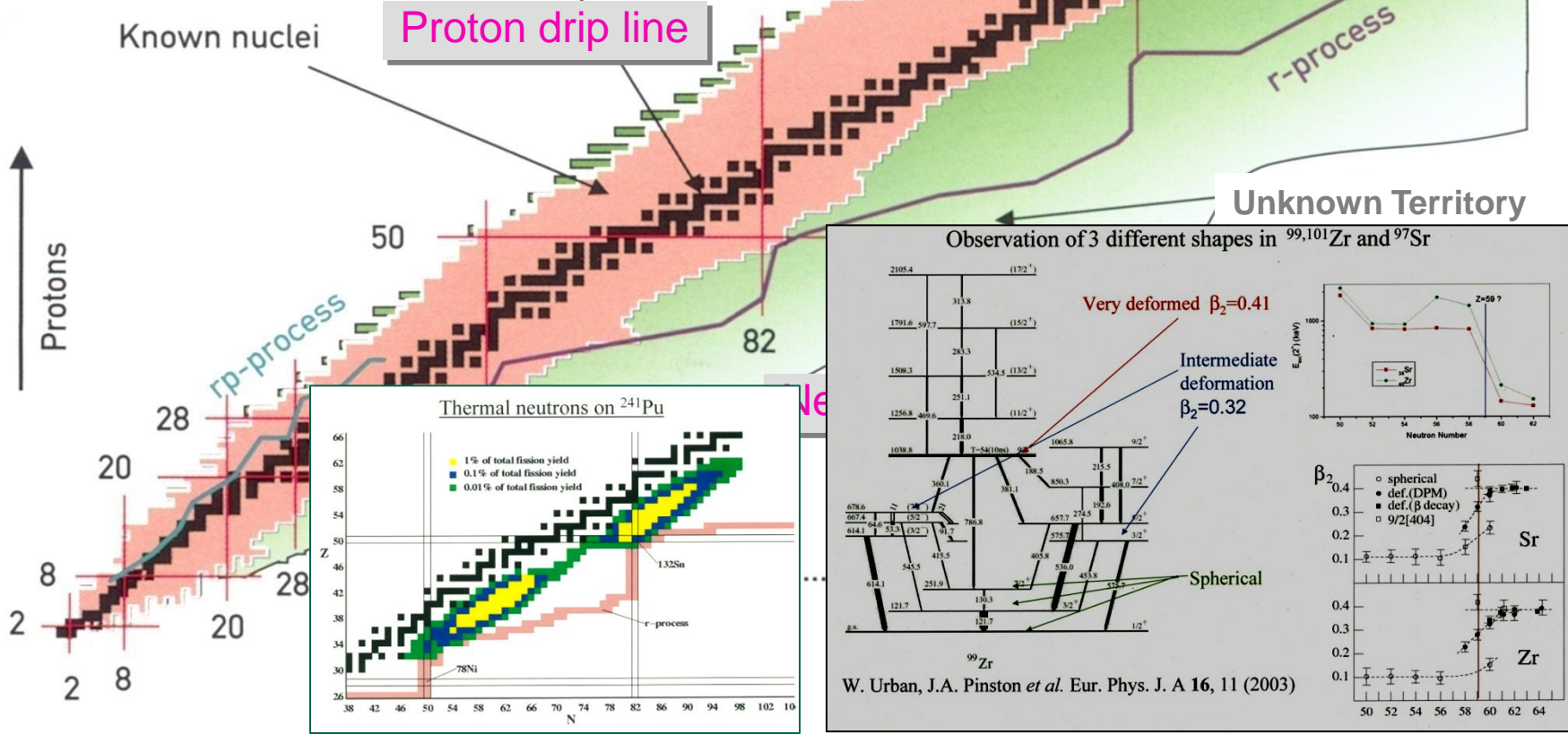
Walter F. Henning - RIKEN Nishina Center





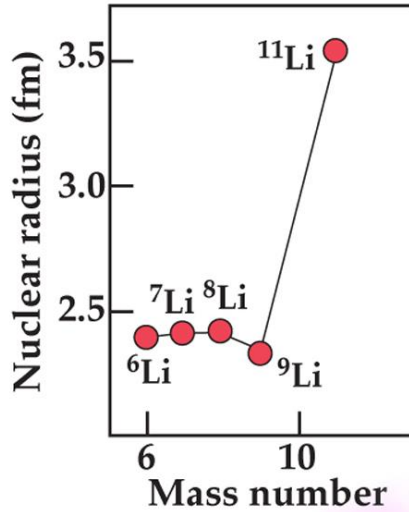


Super heavies

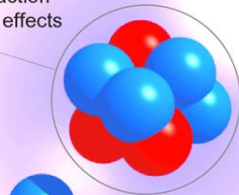


I. Tanihata et al., Phys. Rev. Lett. 55 (1985) 2676
(LBL Bevelac)

Transmission & Interaction Cross Section Measurement



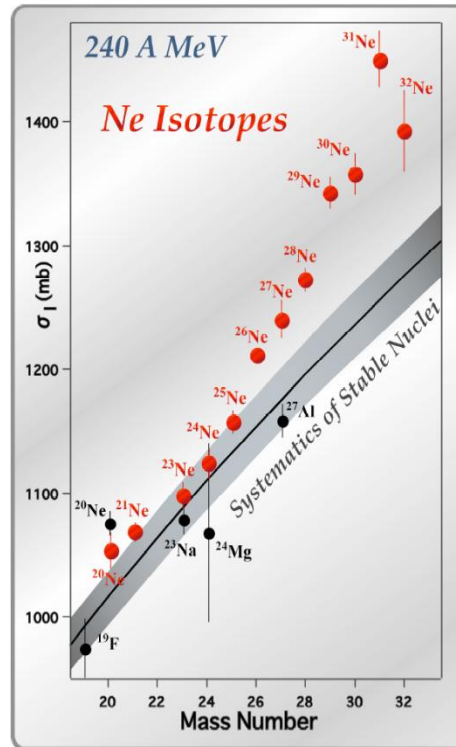
effective NN interaction
strong in-medium effects



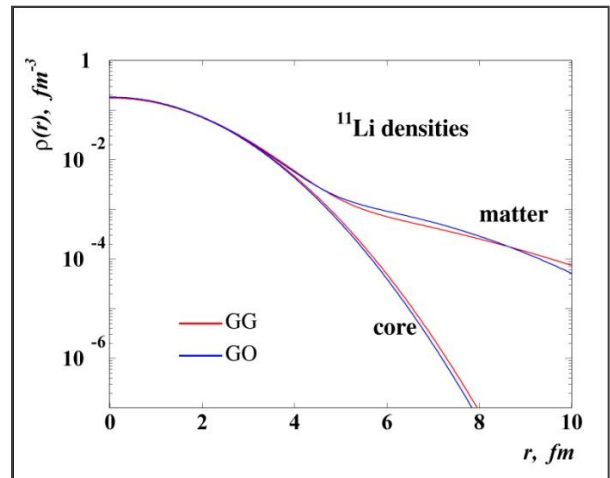
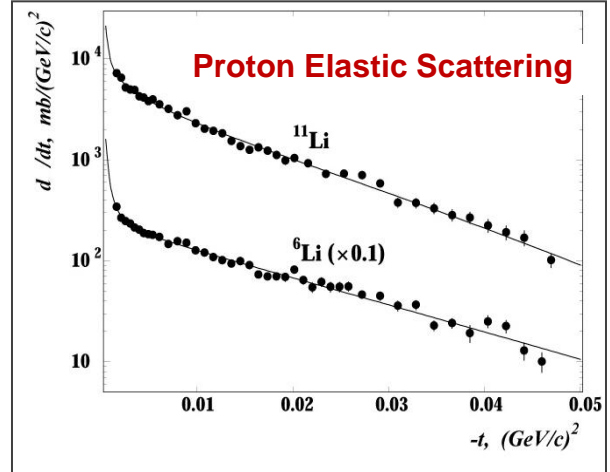
(almost) bare NN interaction
weak in-medium effects

M. Takechi et al., Nucl. Phys. A
834 (2010) 412
(RIKEN RIBF)

Transmission & Interaction Cross Section Measurement



IKAR Collaboration (GSI SIS18):
G.D. Alkharov et al., Nucl. Phys. A 712
(2002) 269
P. Egelhof et al., Eur. Phys. J. A 15
(2002) 27
A. Dobrovolsky et al., Nucl. Phys. A 766
(2006) 1



Pairing

Full shells

Collectivity

Empty shells

Pairing

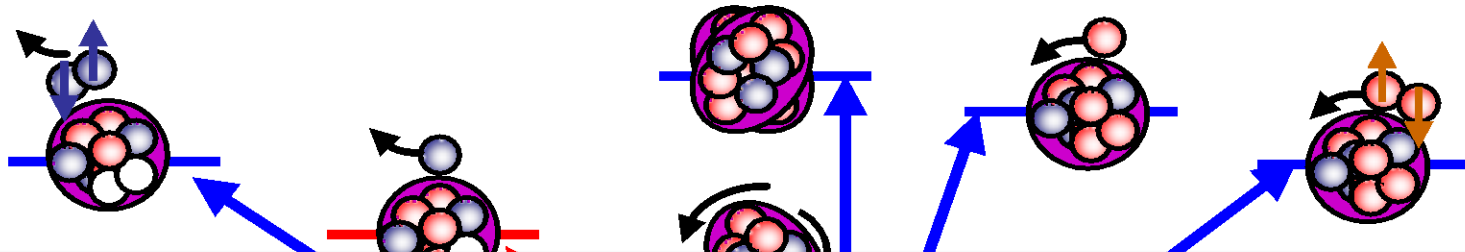
Removing correlated pair

Removing a particle

Rotations & Vibrations

Adding a particle

Adding a correlated pair



Reactions: - Key tool for spectroscopy
- Measurement of reaction cross sections (e.g. nuclear astrophysics etc.)

A-2

A-1

A

A+1

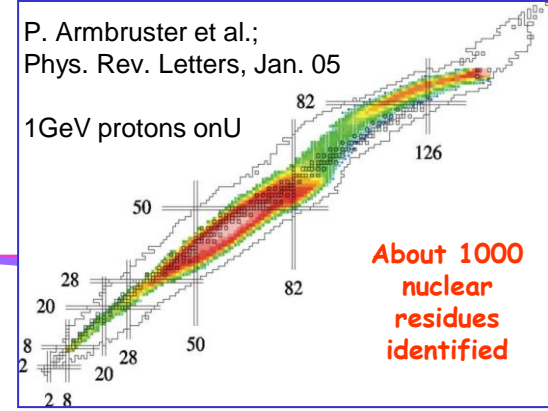
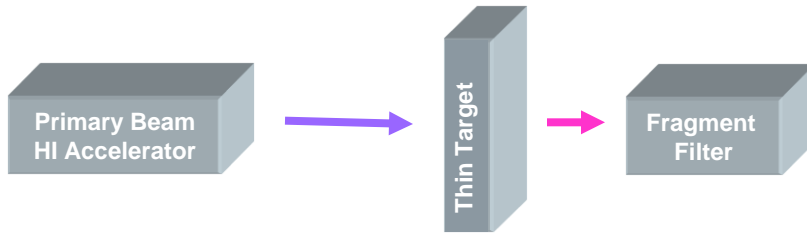
A+2

↑ Fast beams or Reaccelerated beams

Reaccelerated beams ↑

Production Methods for Radioactive Beams

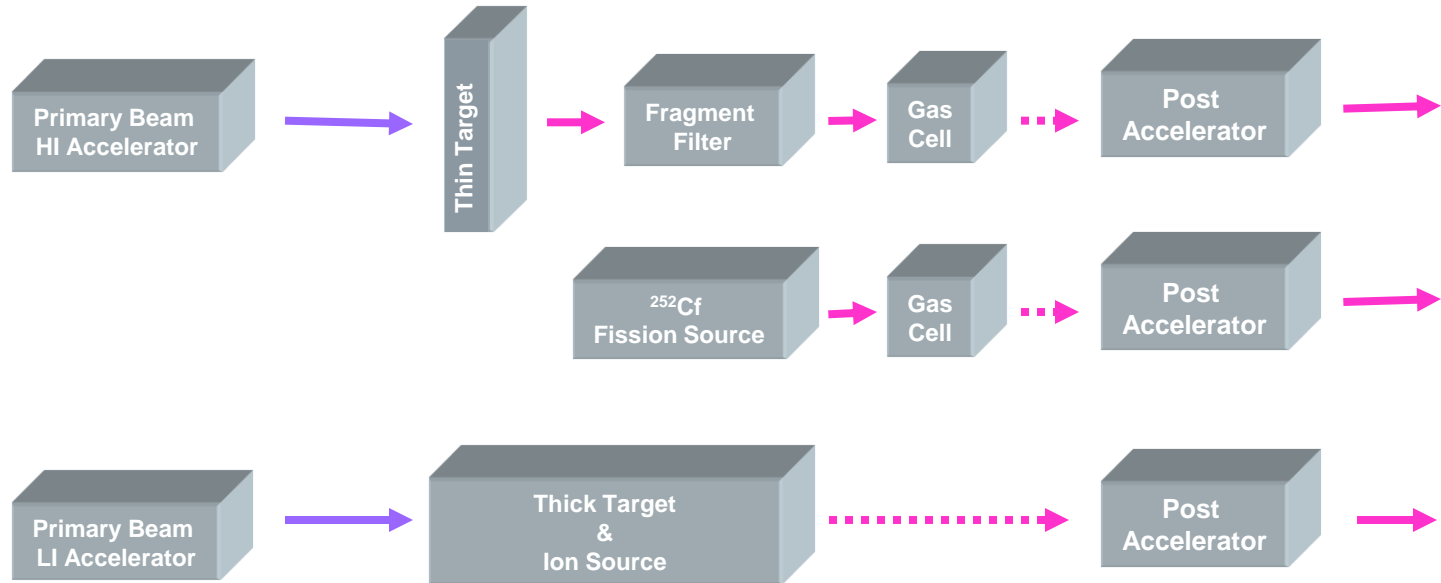
In-Flight Reactions



- High-energy fragmentation
- Low-energy reactions

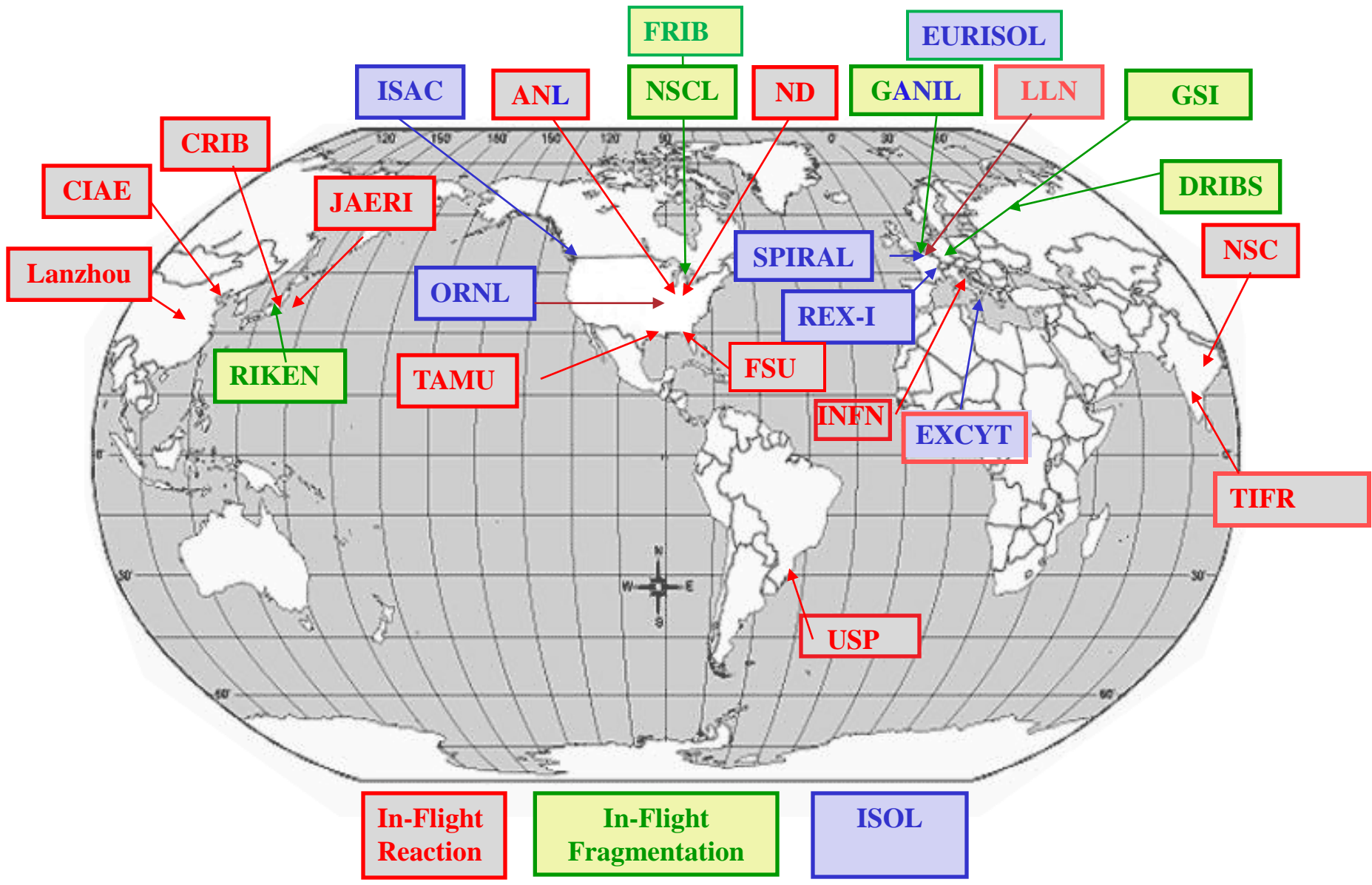
Cross section	Target	Yield curve	Reaction channels	Phase space reduction
large (b)	thick (g)	saturates	many (1000)	passive
medium (~ few mb)	thin (mg)	linear	few (~5-10)	active

Stopping & Re-acceleration (ISOL)

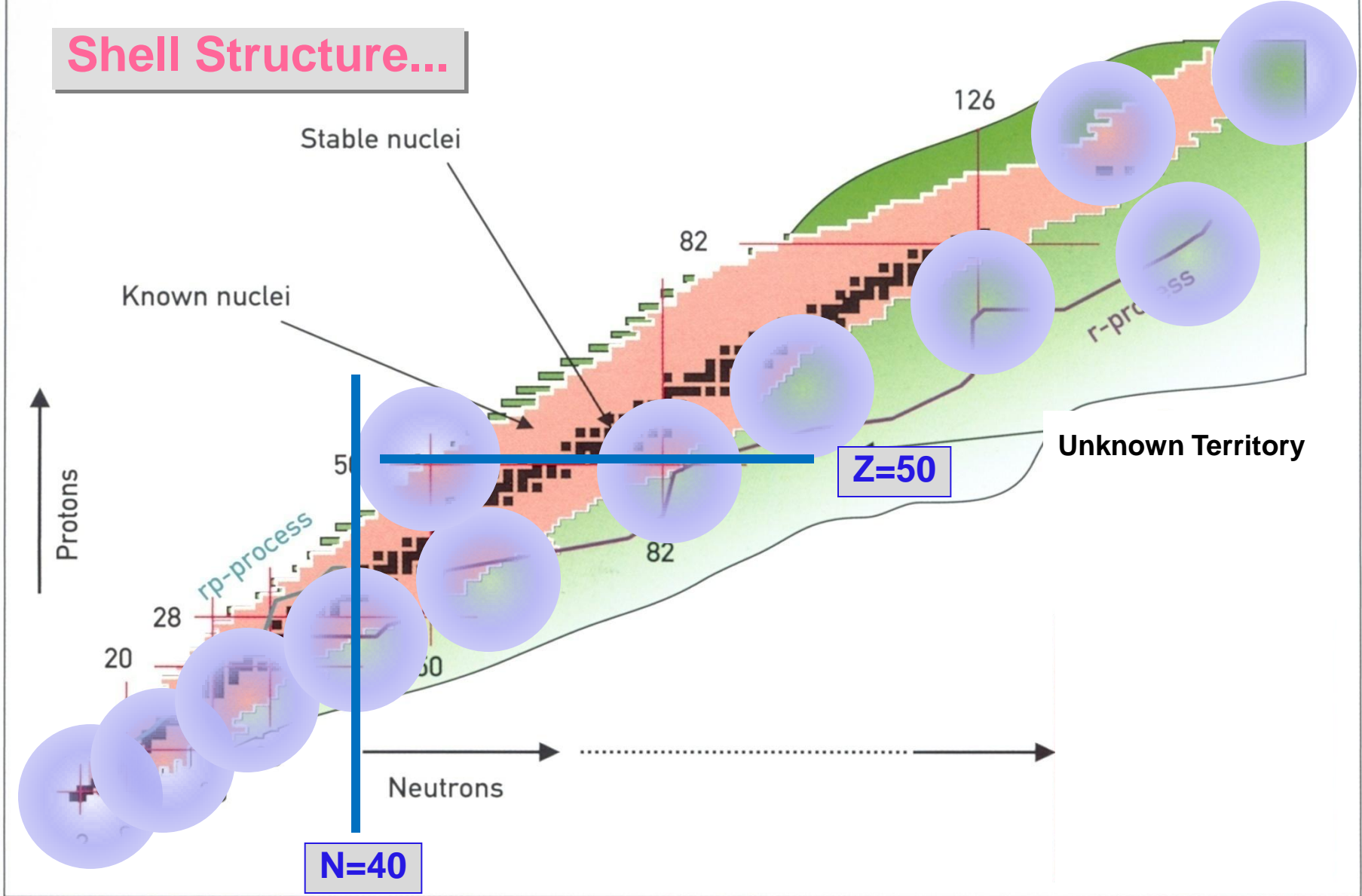


EXPERIMENT

RNB Production Facilities



Shell Structure...

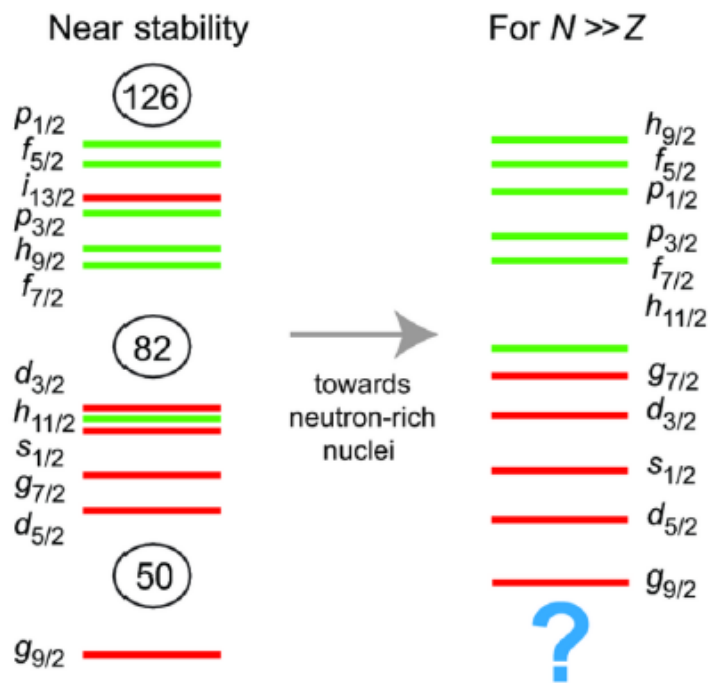


5 next slides:

A. Gade
INPC2010

Driving forces for shell evolution in neutron-rich nuclei

Nuclear Shell Structure



- Mean field near stability
- Strong spin-orbit term

- Mean field for $N \gg Z$?
- Reduced spin-orbit
- Diffuse density
- Tensor force

- Spin-isospin parts of the NN interaction

T. Otsuka et al., PRL 87, 082502 (2001)

- Particularly the monopole parts of the tensor force

T. Otsuka et al., PRL 95, 232502 (2005)

Also

- Central part
- 3N force

play important roles

T. Otsuka et al., PRL 104, 012501 (2010)

T. Otsuka, et al., PRL, in press arXiv:0908.2607

Weak binding:

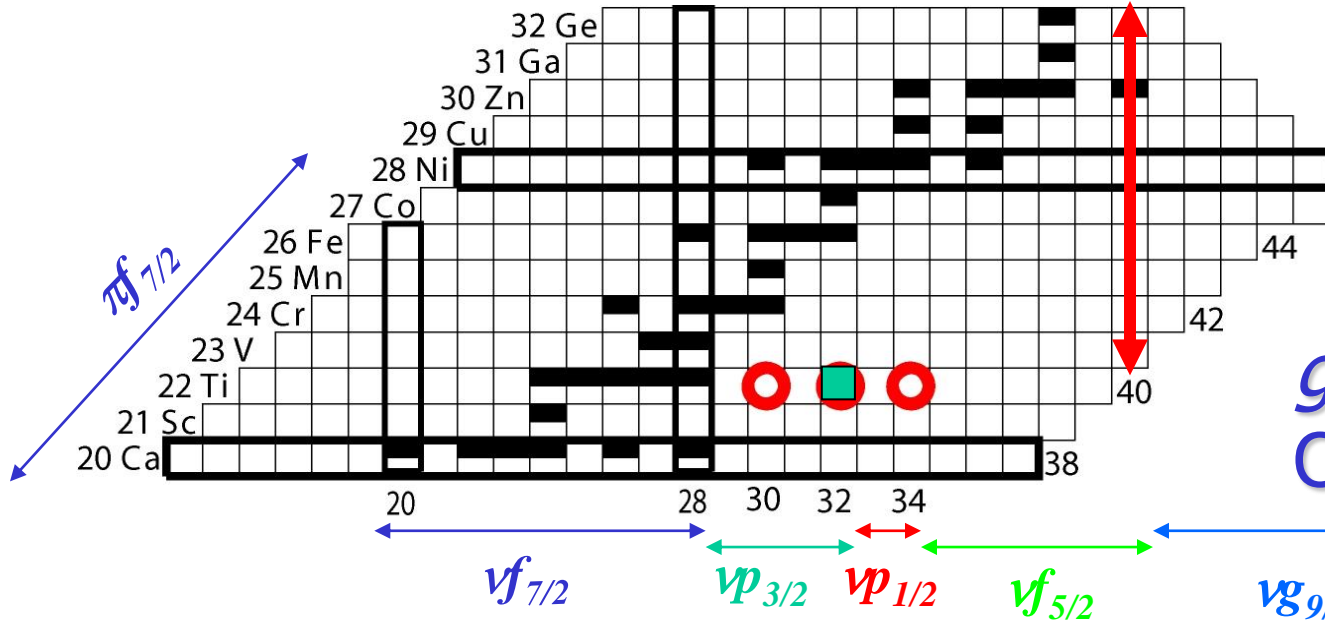
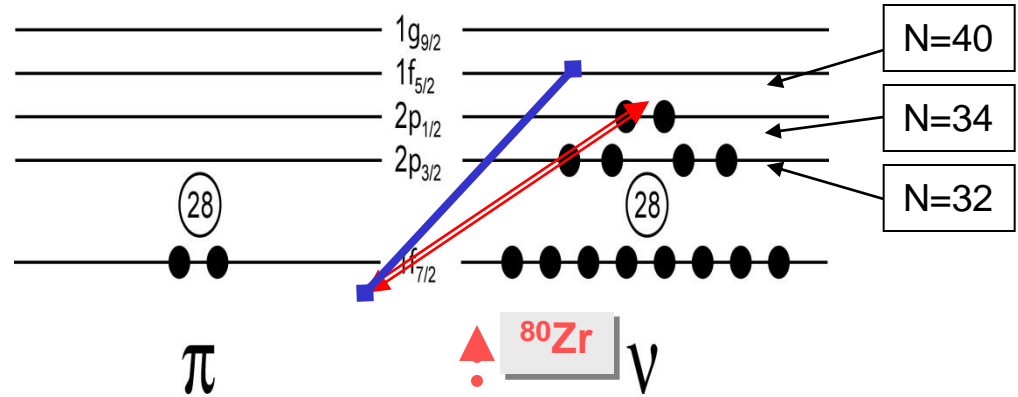
- Density dependent spin-orbit force
- Coupling to the continuum

J. Dobaczewski et al., PNP 59, 432 (2007)

Impacts of tensor force

As protons are removed from the $\pi f_{7/2}$ shell, the monopole pairing interaction weakens:

- $\pi f_{7/2} - \nu f_{5/2}$: attractive, $\nu f_{5/2}$ pushes up in energy
- $\pi f_{7/2} - \nu g_{9/2}$: repulsive, $\nu g_{9/2}$ comes down in energy

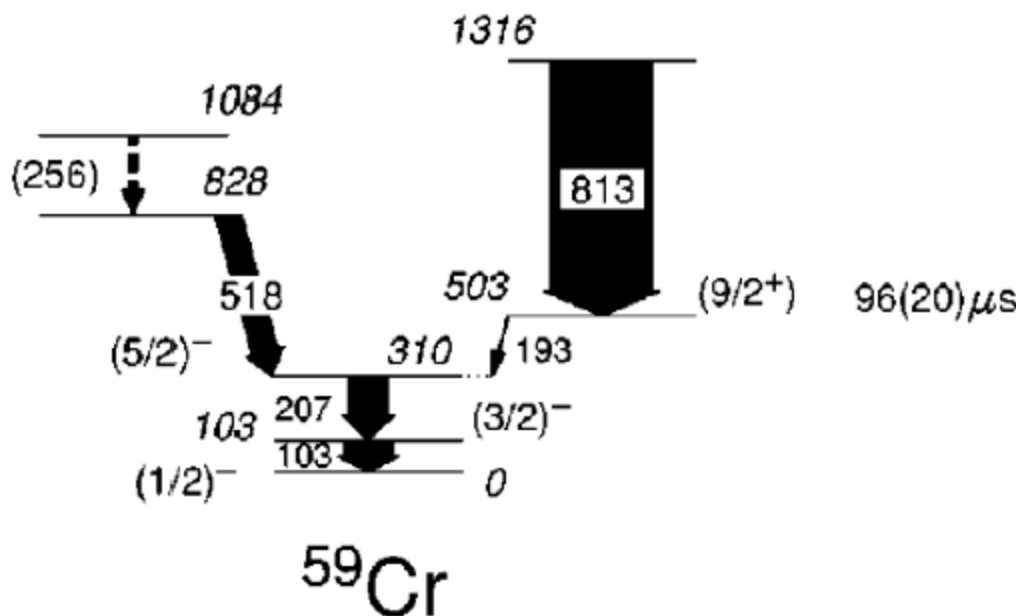


$g_{9/2}$ orbital:
Collectivity

- 1) T. Otsuka *et al.*, Phys. Rev. Lett. 87 (2001),
- 2) T. Otsuka *et al.*, Phys. Rev. Lett. 104 (2010), 012501.

Neutron-rich $N=40$

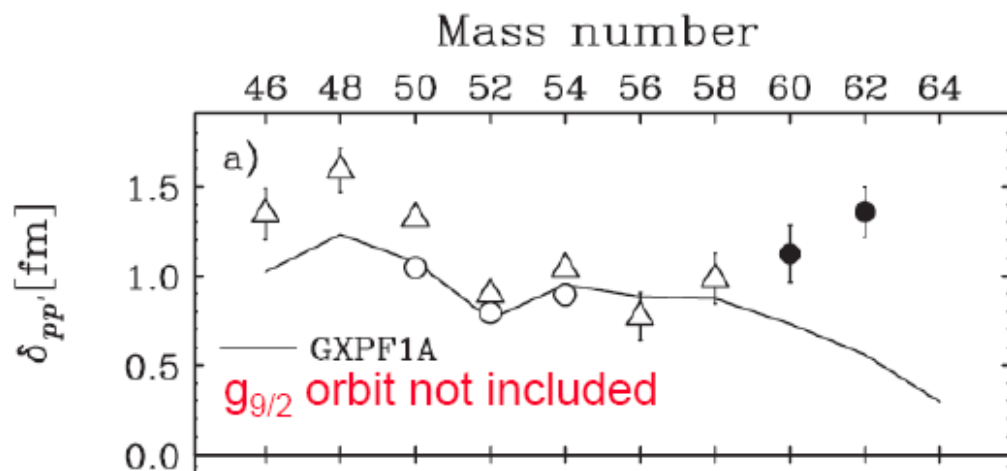
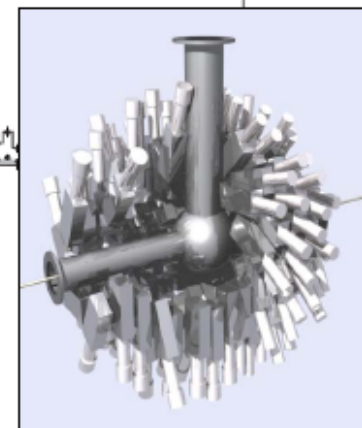
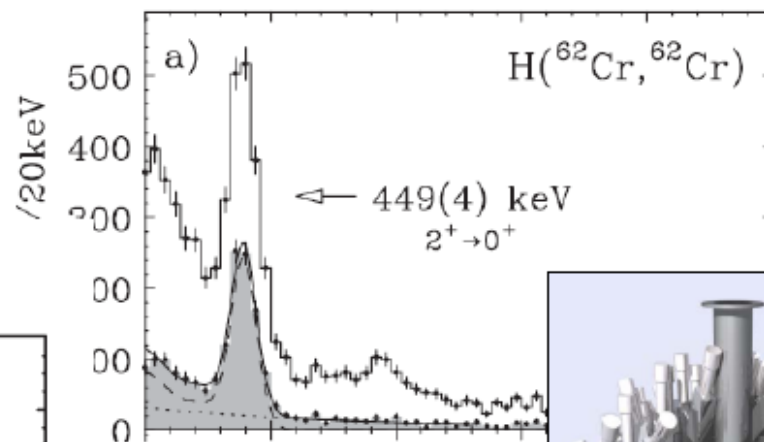
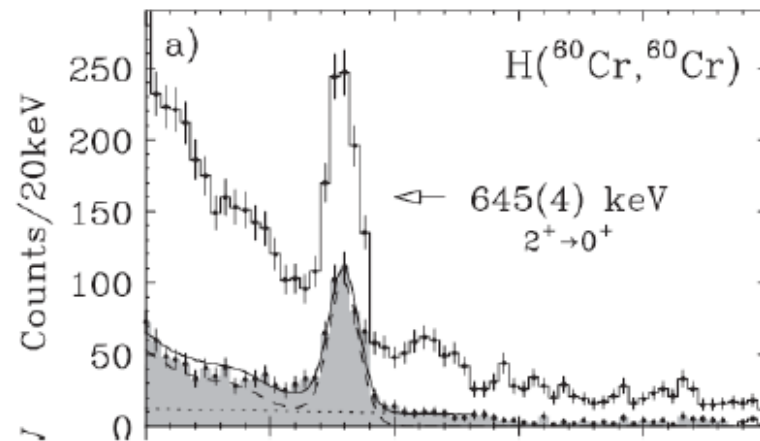
- Beta decay to $^{64,66}\text{Fe}$ at CERN/ISOLDE [M. Hannawald *et al.*, PRL 82, 1391 (1999)]
- Beta decay to $^{60-63}\text{Cr}$ (GANIL) [O. Sorlin *et al.*, EPJ A 16, 55 (2003)]
- Rotational band built on $9/2^+$ in $^{55,57}\text{Cr}$ (ANL) [A. Deacon *et al.*, PLB 622, 151 (2005)]
- Low-lying $9/2^+$ isomer in ^{59}Cr (ANL) at 503 keV with possibly oblate deformation [S. J. Freeman *et al.*, PRC 69, 064301 (2005)]
- Large deformation of ^{62}Cr (RIKEN) [N. Aoi *et al.*, J. Phys.: Conf. Ser. 49, 190 (2006), PRL 102, 012502 (2009)]



Collectivity in neutron-rich Cr

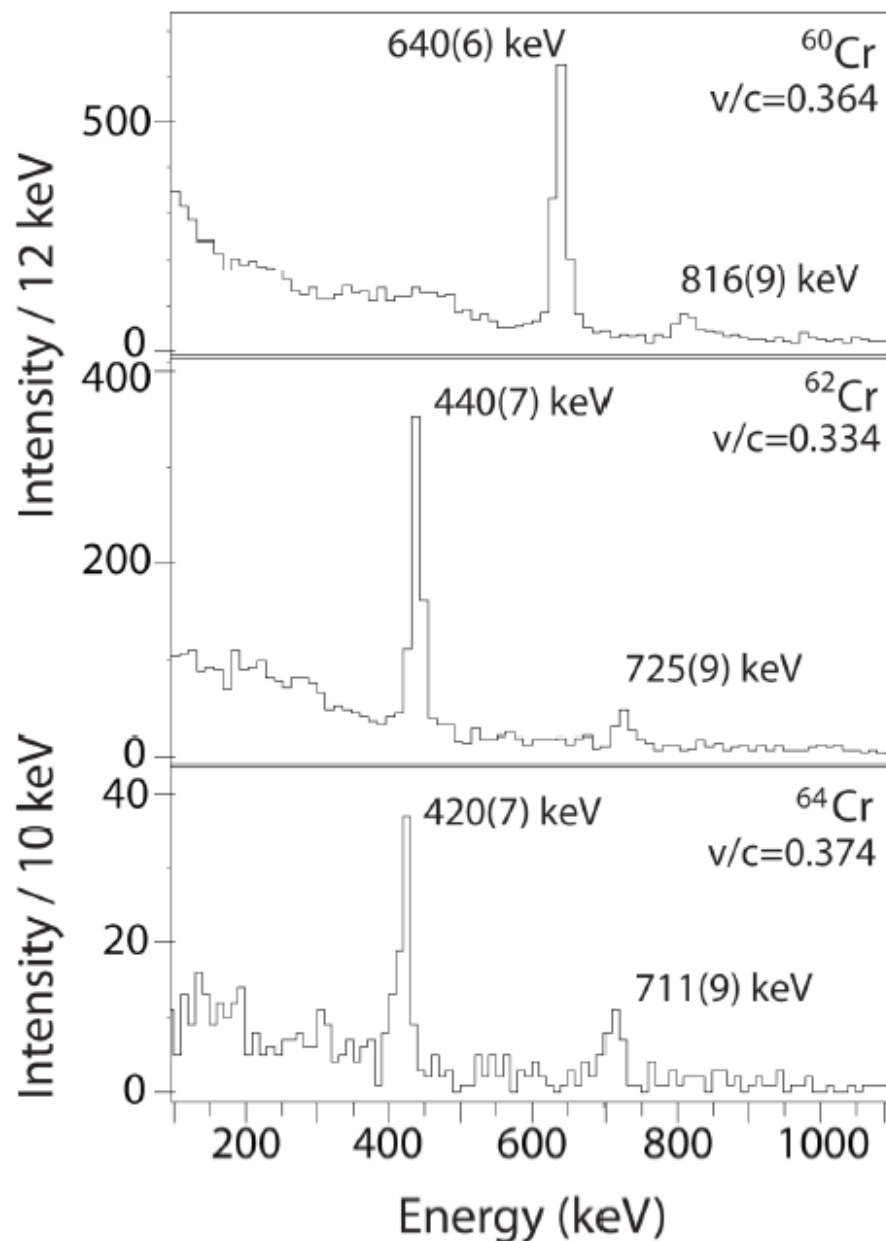
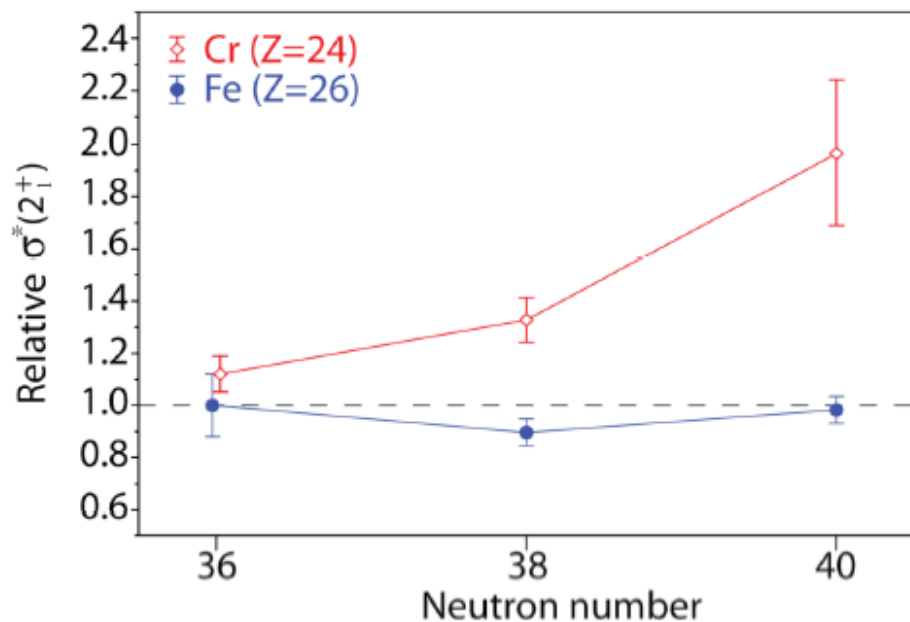
- $^{60,62}\text{Cr}$ produced by fragmentation of stable ^{70}Zn beam at 63 MeV/u
- Inelastic proton scattering induced by a LH_2 target, γ -ray detected in NaI array DALI2 and scattered Cr in a TOF spectrometer (TOMBEE)
- Deformation length deduced from angle-integrated cross sections, $\sigma(2^+)$, in comparison to CC calculations
- Data not described by models without neutron $g_{9/2}$ orbit

N. Aoi et al., PRL 102, 012502 (2009) - RIKEN



Inelastic scattering of $^{62,64,66}\text{Fe}$ and $^{60,62,64}\text{Cr}$ at NSCL

- Gamma-ray spectroscopy of excited states populated in $^{60,62,64}\text{Cr}$ and $^{62,64,66}\text{Fe}$ via inelastic scattering off ^9Be
- Relative cross section for the population of the 2^+ state – including observed and unobserved feeding – confirms change in structure between Fe and Cr



Deformation in Co and Ni approaching N=40:

D. Pauwels *et al.*, PRC 78, 041307(R) (2008); PRC 79, 044309 (2009)

N=40

	^{62}Fe		^{64}Fe	^{66}Fe		^{68}Fe
	^{60}Cr		^{62}Cr	^{64}Cr		

Neutron-rich Cr up to ^{62}Cr :

N. Aoi *et al.*, PRL 102, 012502 (2009)

S. Zhu *et al.*, PRC 74, 064315 (2006)

Structure of iron up to ^{68}Fe :

J. Ljungvall *et al.*, PRC 81, 061301(R) (2010)

P. Adrich *et al.*, PRC 77, 054306 (2008)

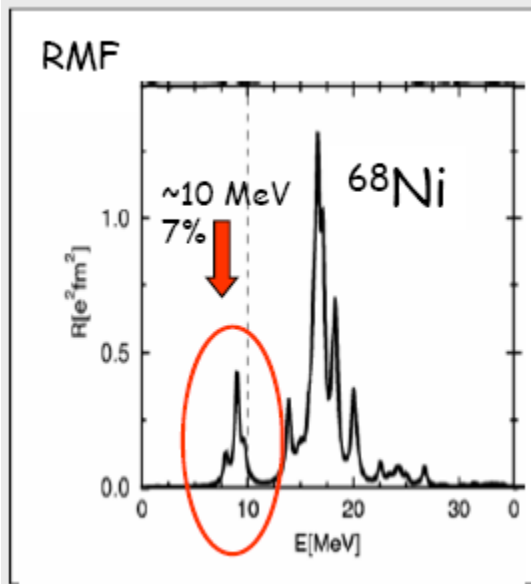
S. Lunardi *et al.*, PRC 76, 034303 (2007)

Structure of Mn up to ^{63}Mn :

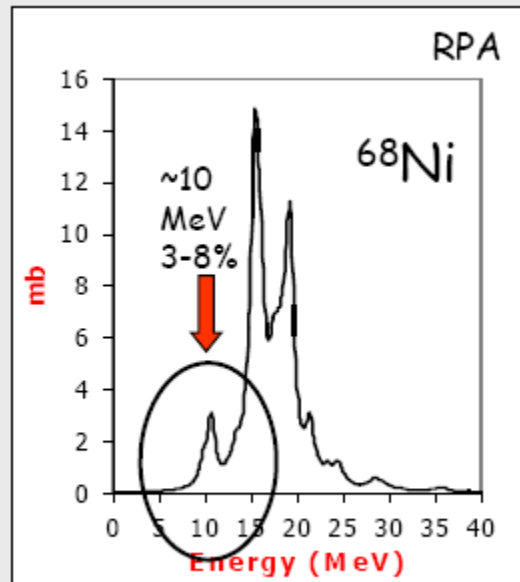
D. Steppenbeck, *et al.*, PRC 81, 014305 (2010)

H. L. Crawford, *et al.*, PRC 79, 054320 (2009)

J. J. Valiente-Dobón *et al.*, PRC 78, 024302 (2008)

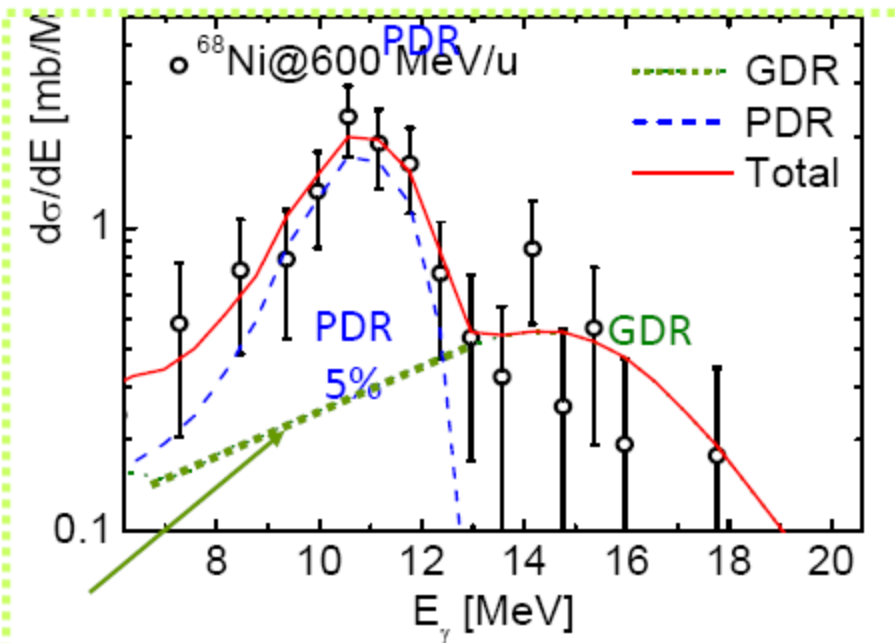


D. Vretnar et al. NPA 692(2001)496



G. Colo private communications

+ J. Liang et al.,
PRC75(2007)
fRPA: 7-8%:



O. Wieland et al., PRL102(2009)092502

Pygmy in ^{68}Ni at 11 MeV

Width $\approx 2\text{ MeV}$ mainly due to Doppler Broadening

5 (p/m 1.5) % of the EWSR

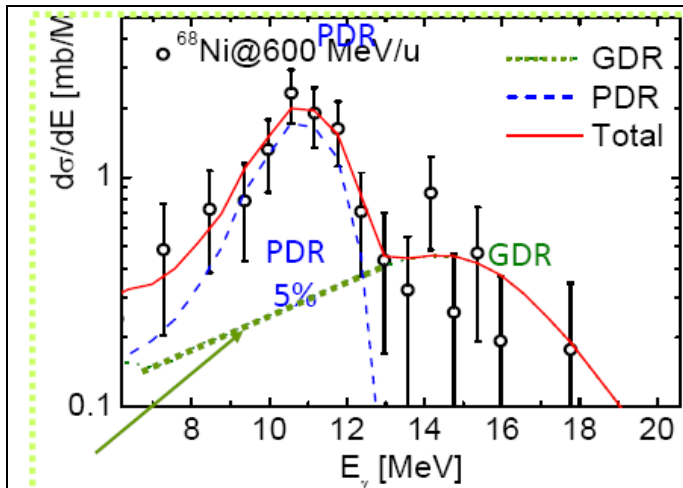
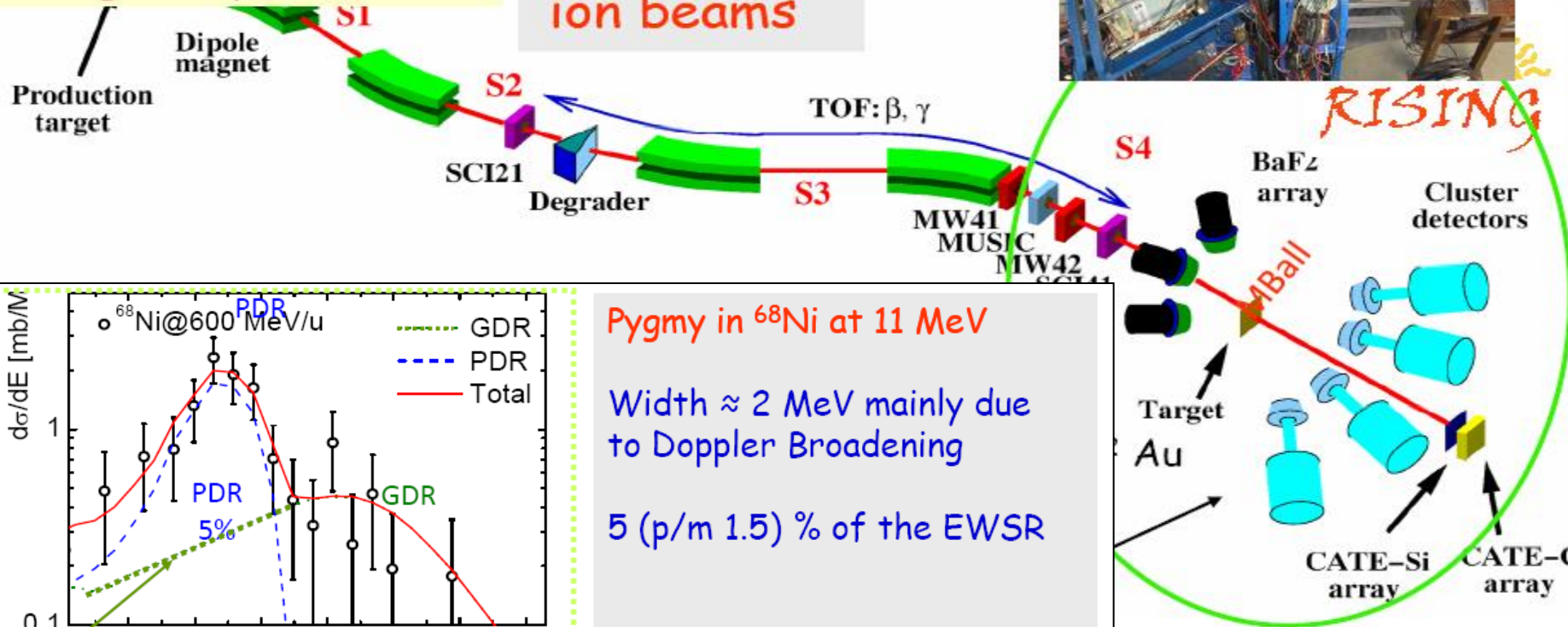
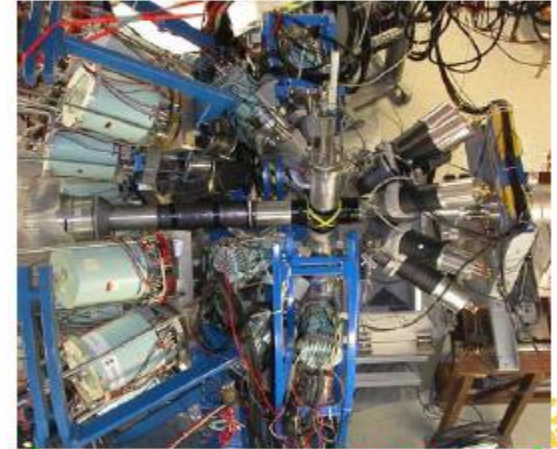
$B(E1) = 1.2\text{ e}^2\text{fm}^2$

High resolution γ -spectroscopy at the FRS of GSI

□ ^{68}Ni beam by fragmentation of ^{86}Kr @ 900 MeV/u on Be target ($4\text{g}/\text{cm}^2$):

□ 10^{10} pps pill ^{86}Kr , Spill length 6s, period 10 s

FRS provides secondary radioactive ion beams



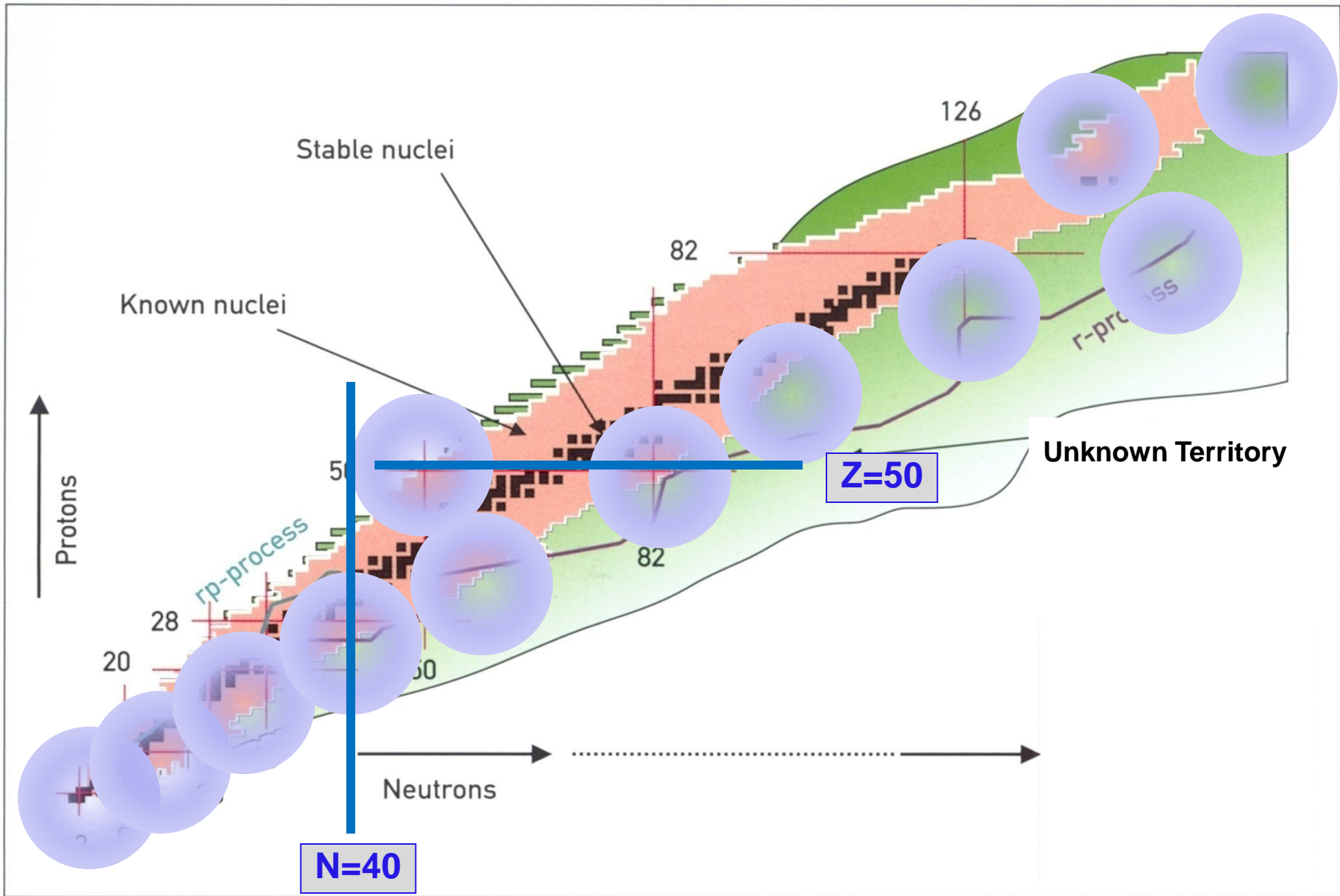
Pygmy in ^{68}Ni at 11 MeV

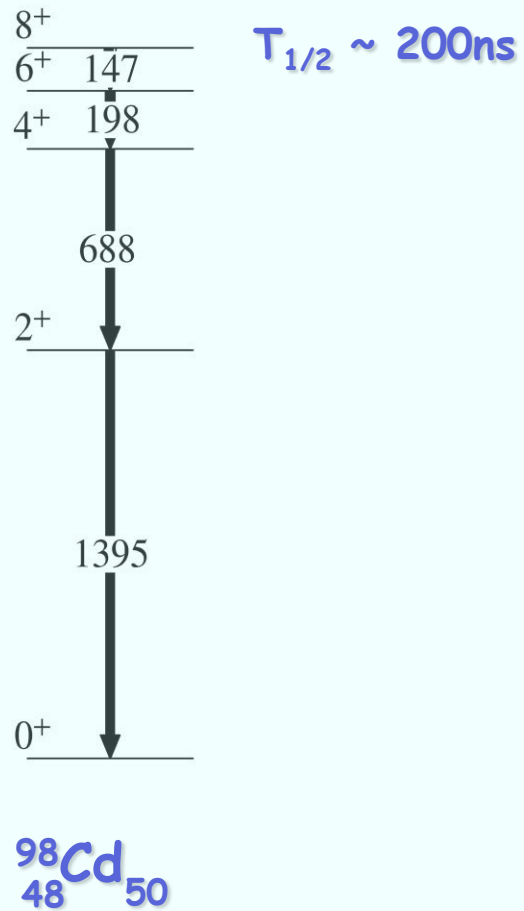
Width ≈ 2 MeV mainly due to Doppler Broadening

5 (p/m 1.5) % of the EWSR

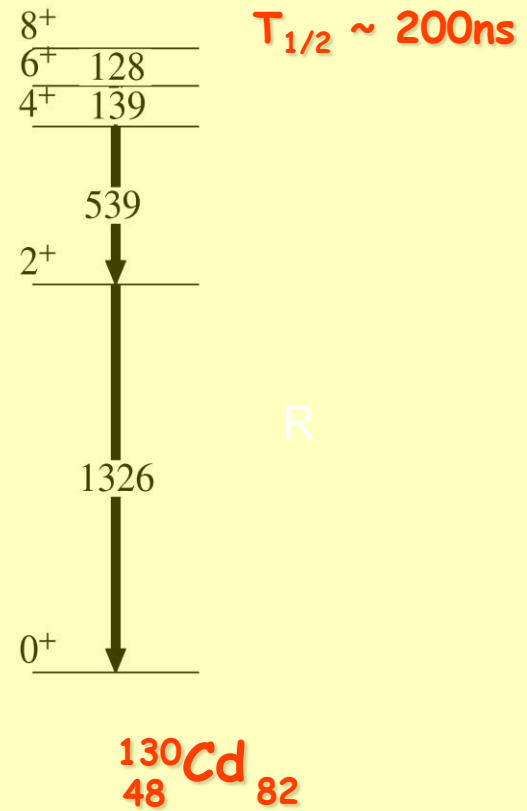
$$B(E1) = 1.2 e^2\text{fm}^2$$

O. Wieland et al., PRL102(2009)092502

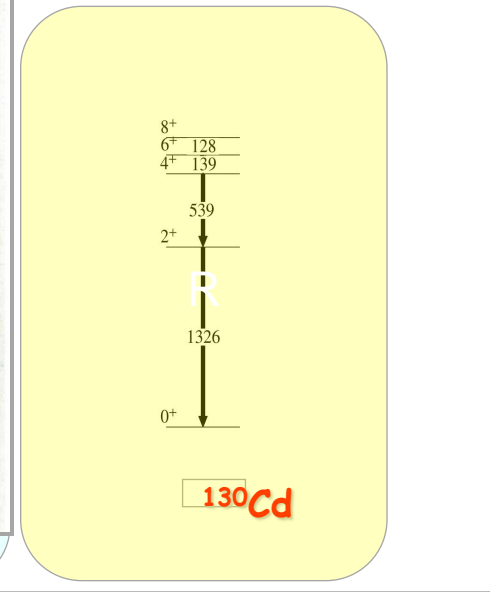
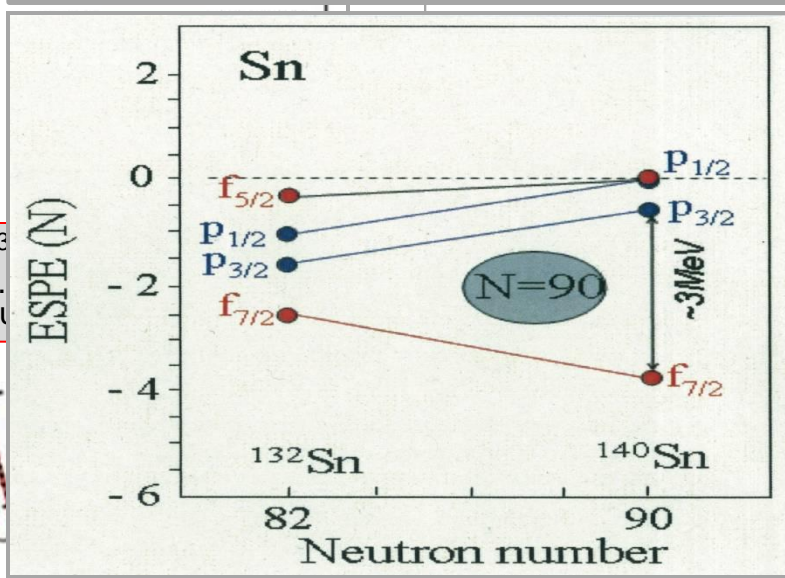
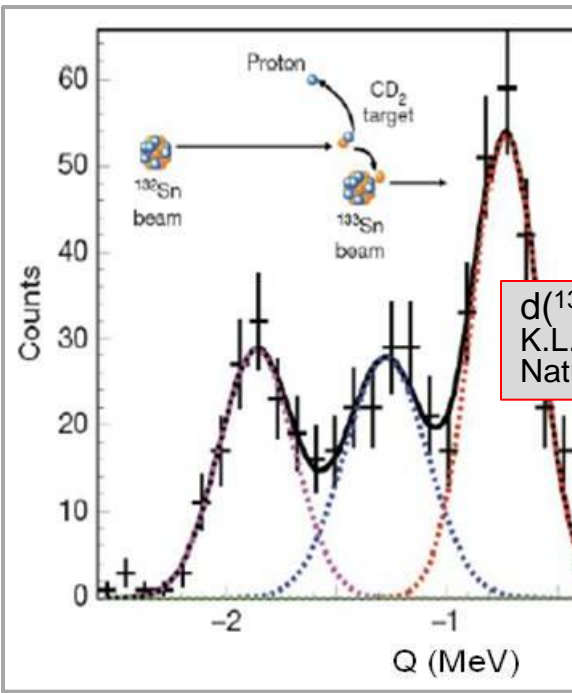
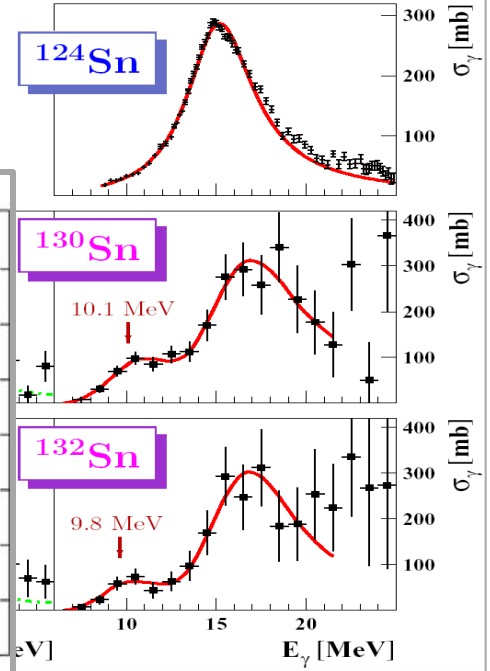
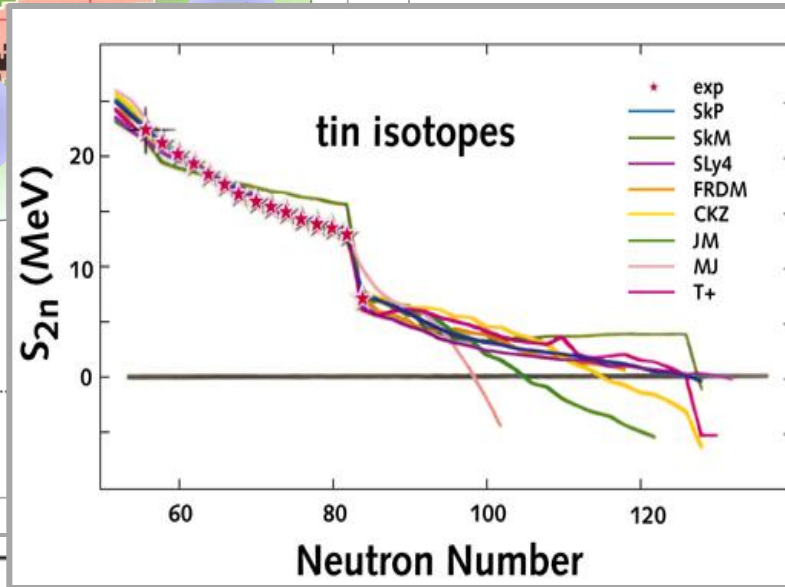
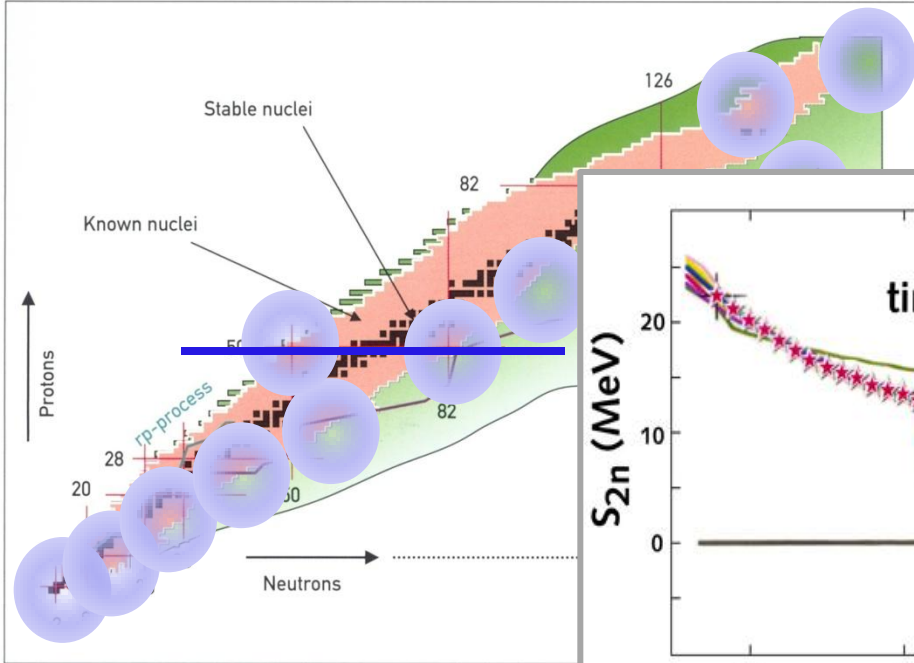




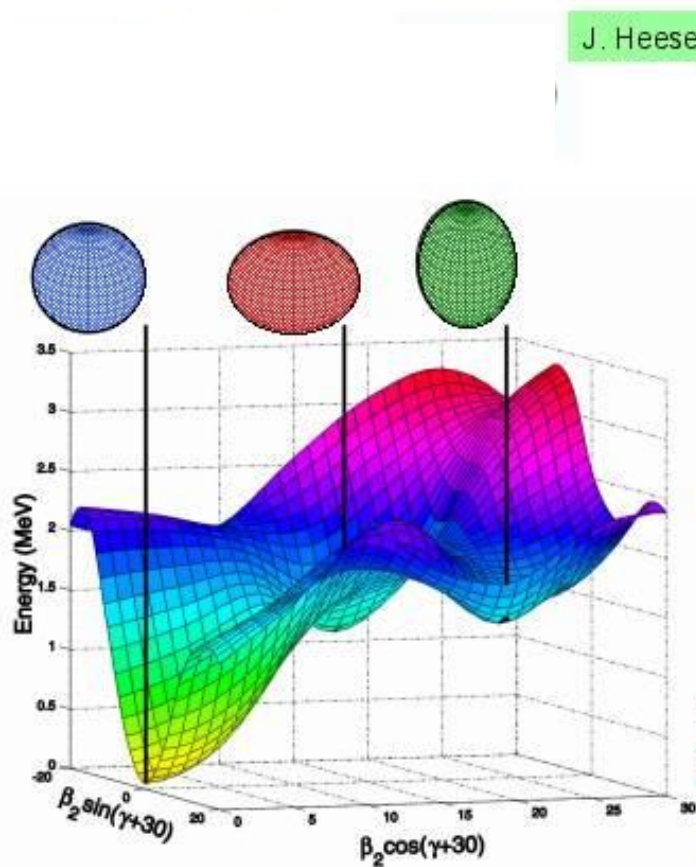
M. Górska et al., PRL 79 (1997) 2415



A. Jungclaus et al., PRL 99 (2007) 132501

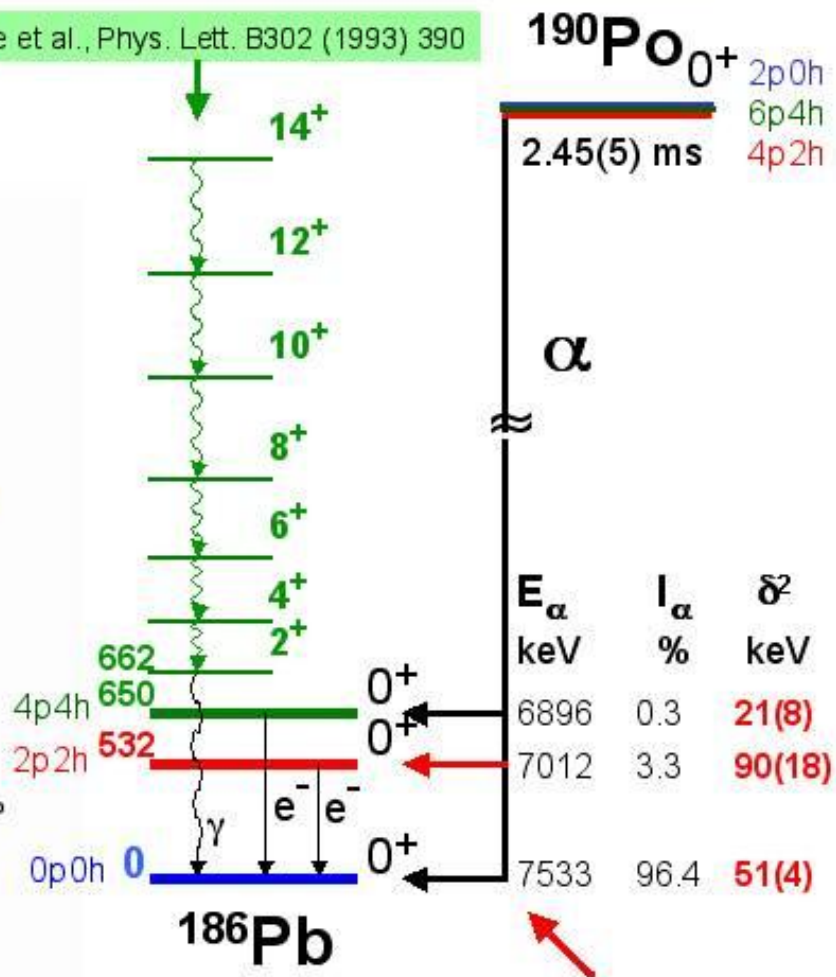


Shape Co-existence and Quantum Phase Transitions



Potential Energy Surface for ^{186}Pb

J. Heese et al., Phys. Lett. B302 (1993) 390



A. Andreyev et al., Nature 405 (2000) 430

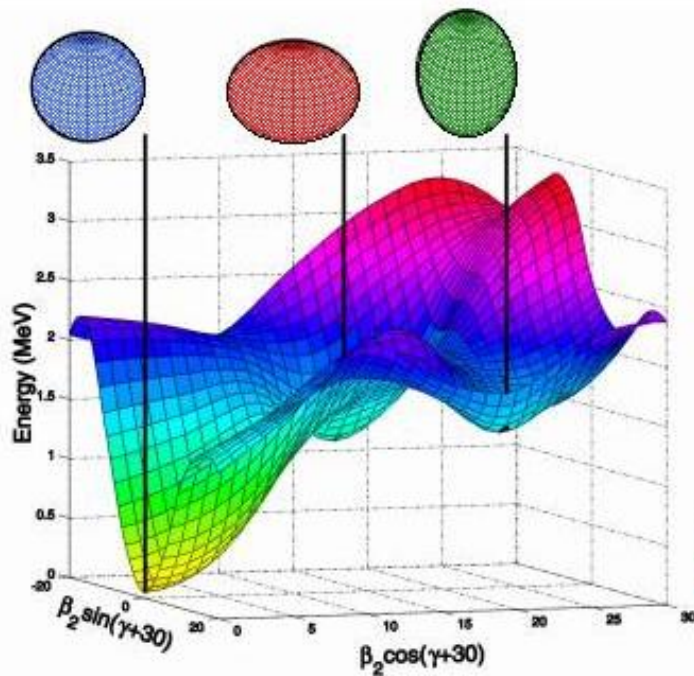
Quantum Phase Transitions QPT

QPT defined for $N \rightarrow \infty$

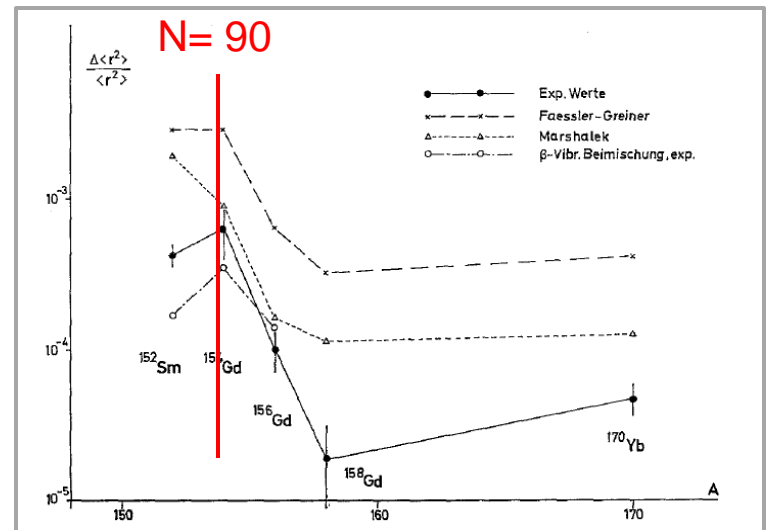
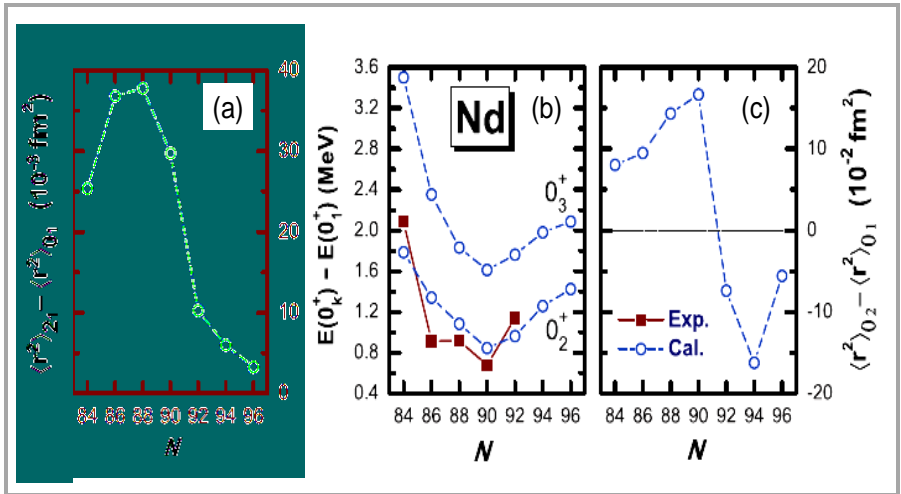
Iachello and Zamfir (PRL 92 (2004) 212501) show that main features persist down to $N \sim 10$

Z.P. Li et al. PRC 80 (2009) 061301 present calculations on observables \longrightarrow

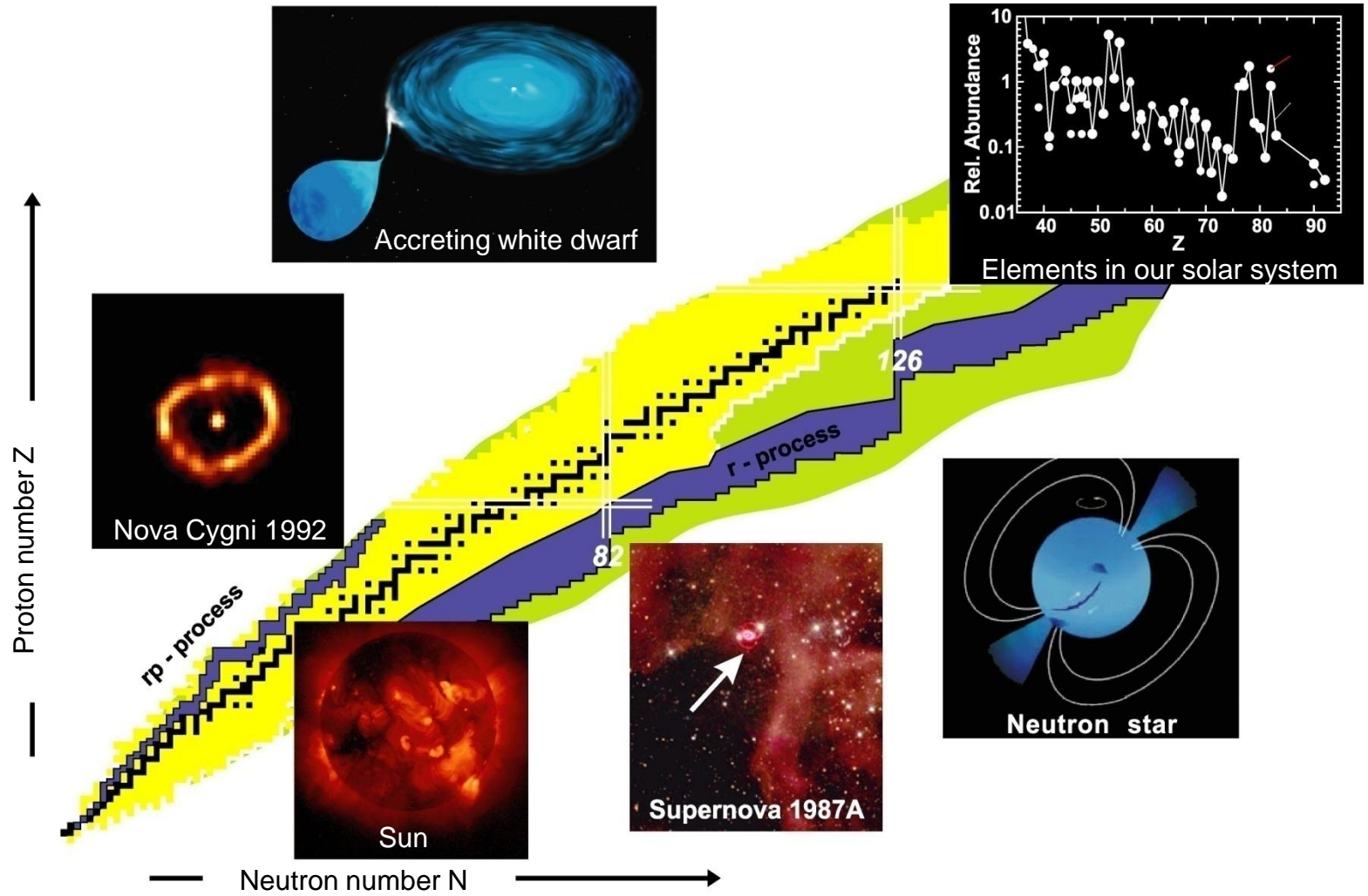
ISOLDE: QPT in masses of n-rich Kr isotopes



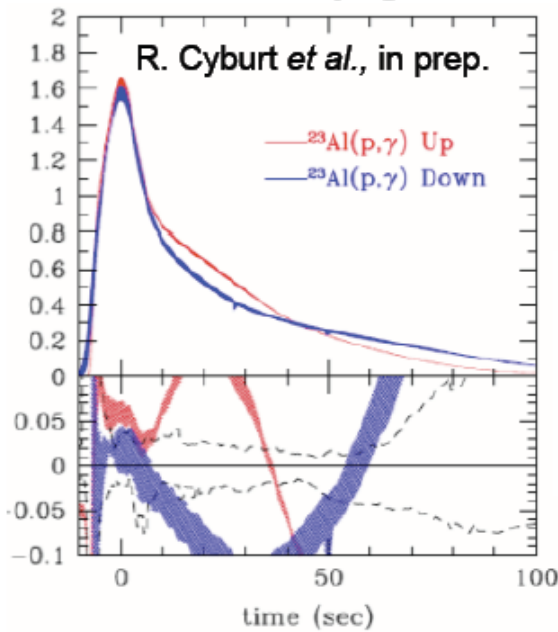
Potential Energy Surface for ^{186}Pb



Nuclear astrophysics with RI Beams:



Calculated X-ray light curve



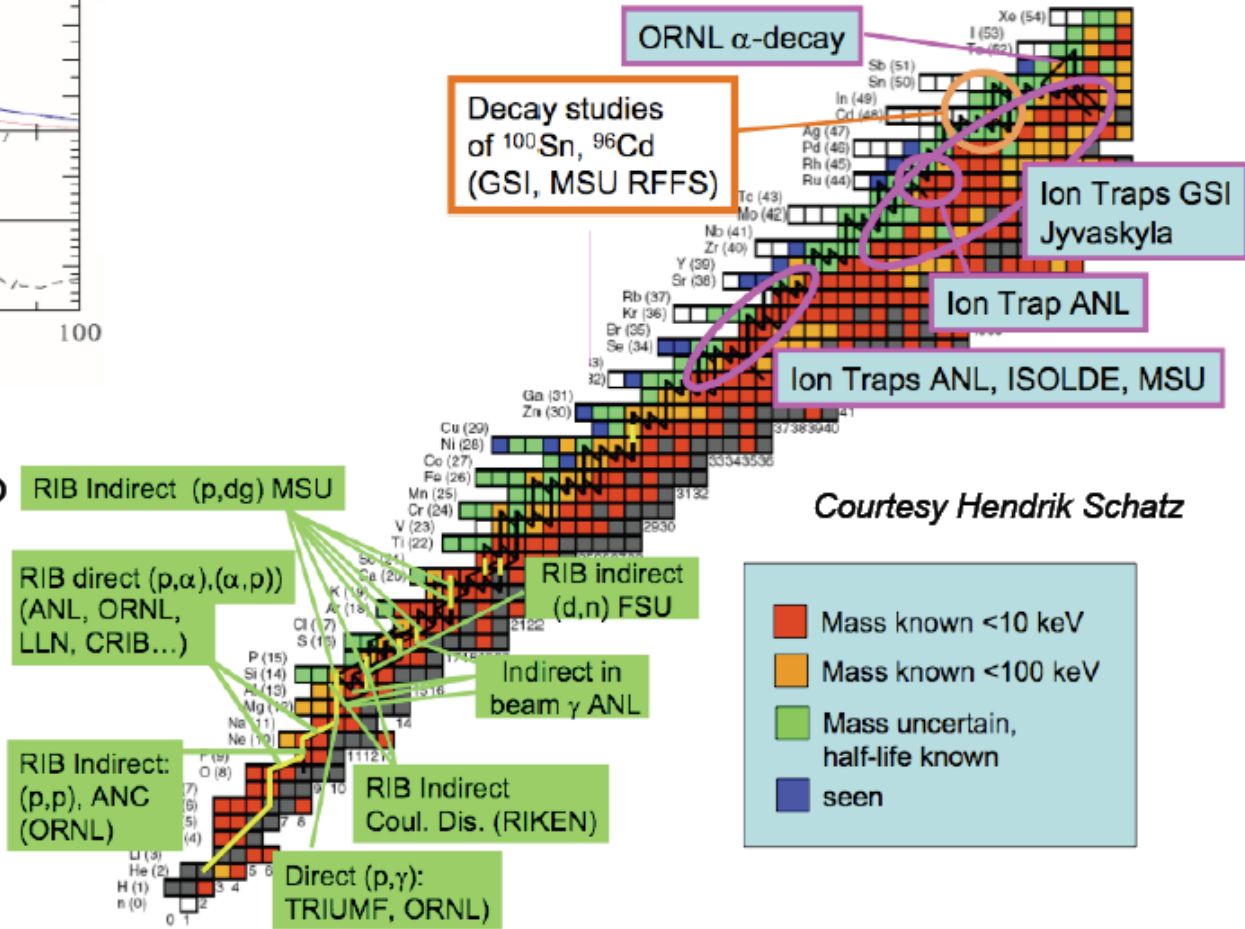
X-ray bursts

- Reaction rates are important, but few direct measurements have been possible so far

- Sensitivity studies allow experiments to be focused on most important rates
- Must improve data for statistical rates

$$\langle \sigma v \rangle \propto e^{-Q/kT}$$

➔ Masses
- Decay properties

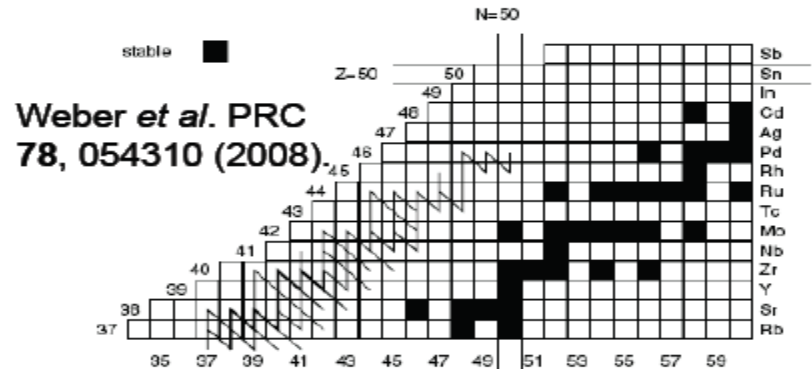


Courtesy Hendrik Schatz

vp process

Fröhlich *et al.*, PRL **96**, 142502 (2006).

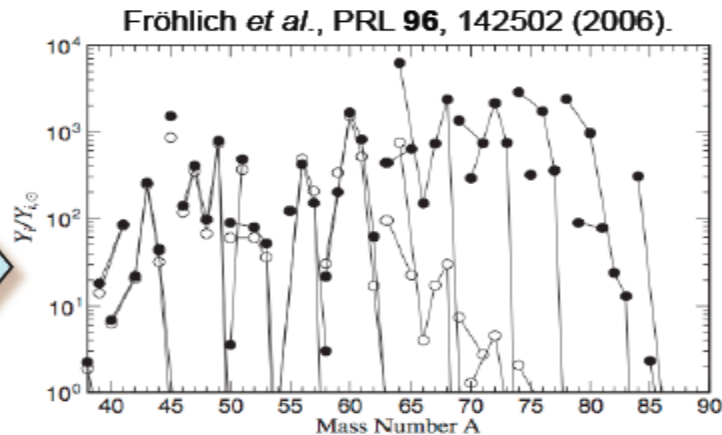
- The intense neutrino/antineutrino flux in a supernova can strongly influence the formation of elements
- Nuclei that form in the “alpha-rich freezeout” are expected to experience a high anti-neutrino flux
- $\bar{\nu} + p \rightarrow n + e^+$ reactions produce neutrons that bypass the decay of long-lived nuclei
- Reaction sequence starting on iron-group nuclei that is similar to the rp process



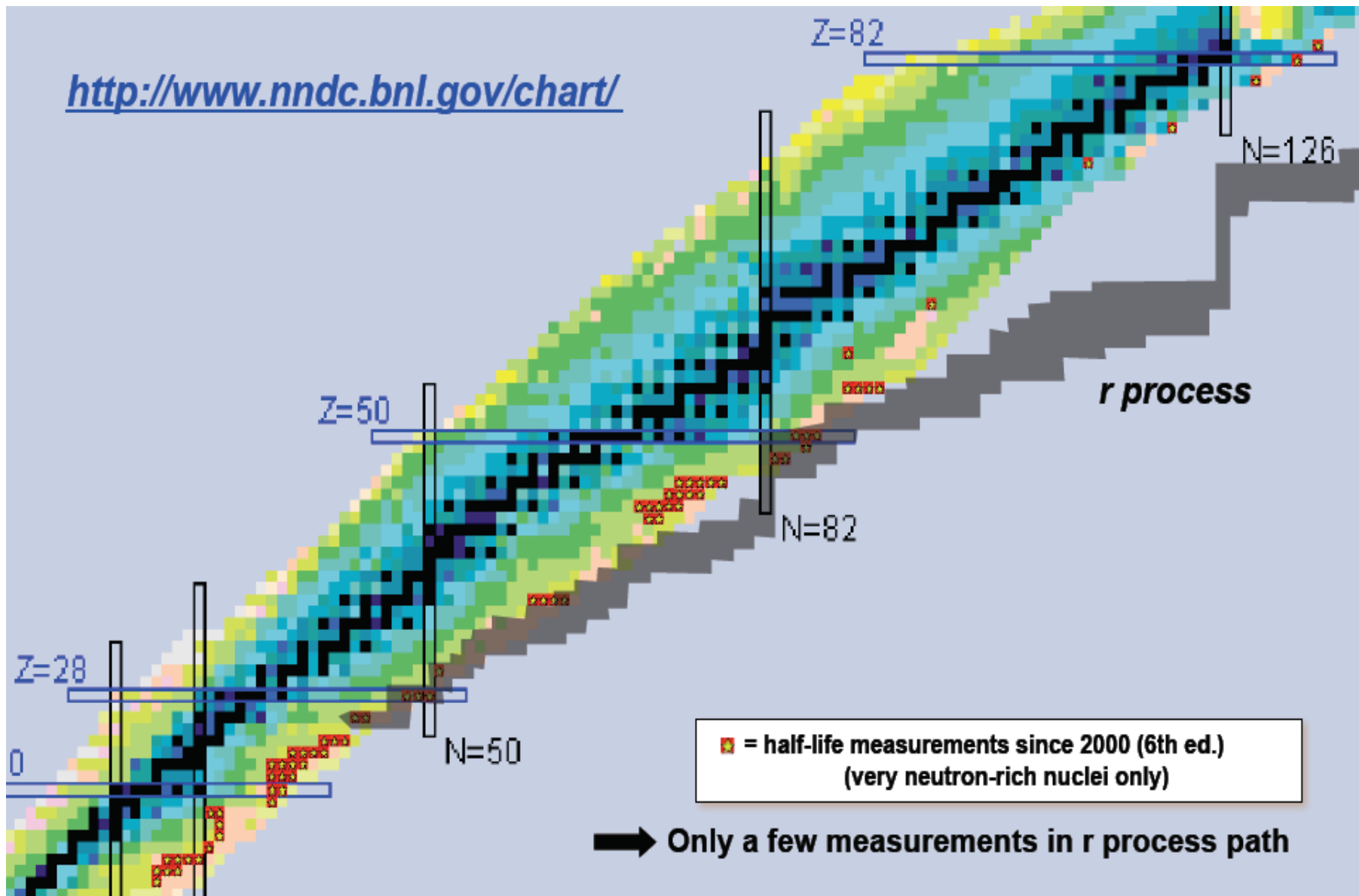
See Suzuki
5:00pm Thur

Abundances relative to solar

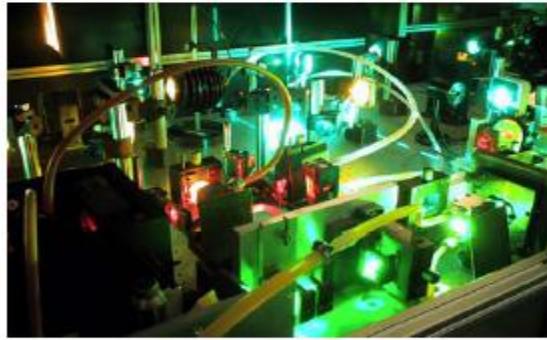
- with ν reactions
- without ν reactions



<http://www.nndc.bnl.gov/chart/>



Groundstate Properties of Nuclei: Lasers, Traps, and Storage Rings



What do we learn?

Charge Radii

nuclear structure: nuclear charge distribution, deformation

Spins, Moments

microscopic nuclear structure: wave functions, coupling of nucleons, configuration mixing, shell structure

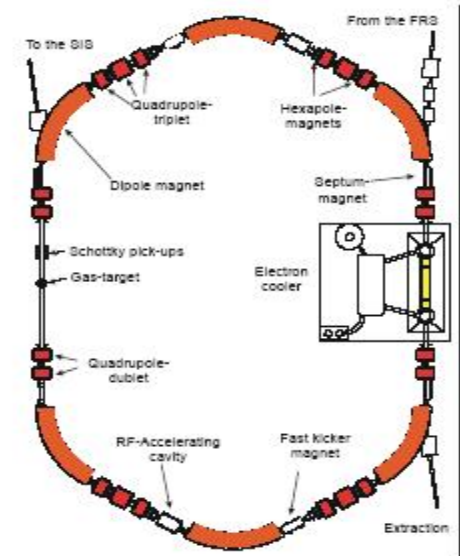
macroscopic nuclear structure: deformation

Masses

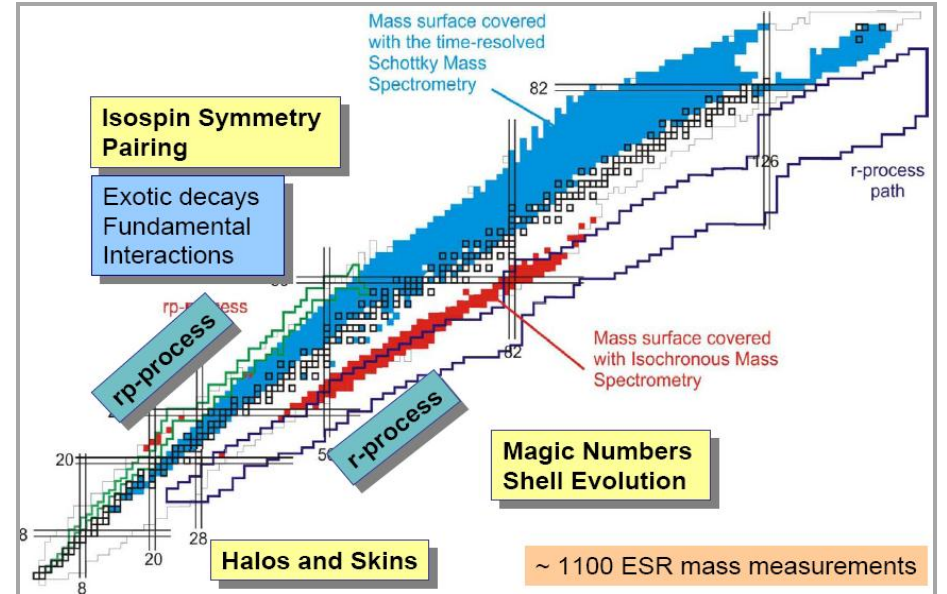
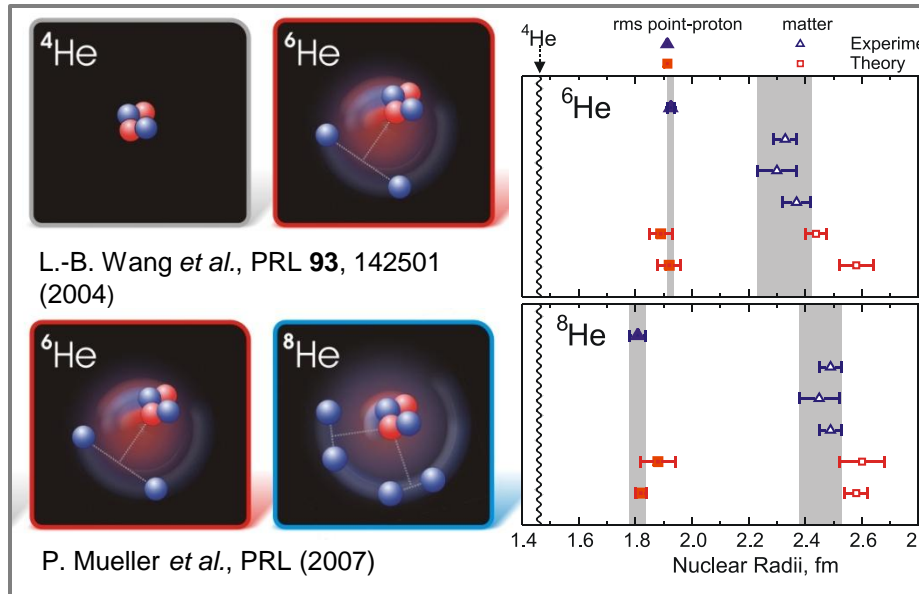
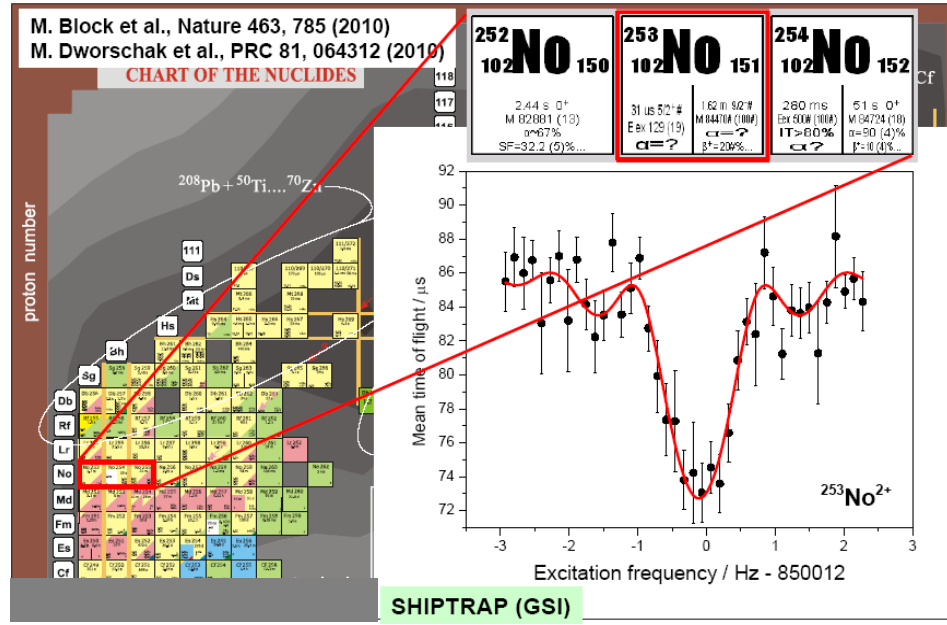
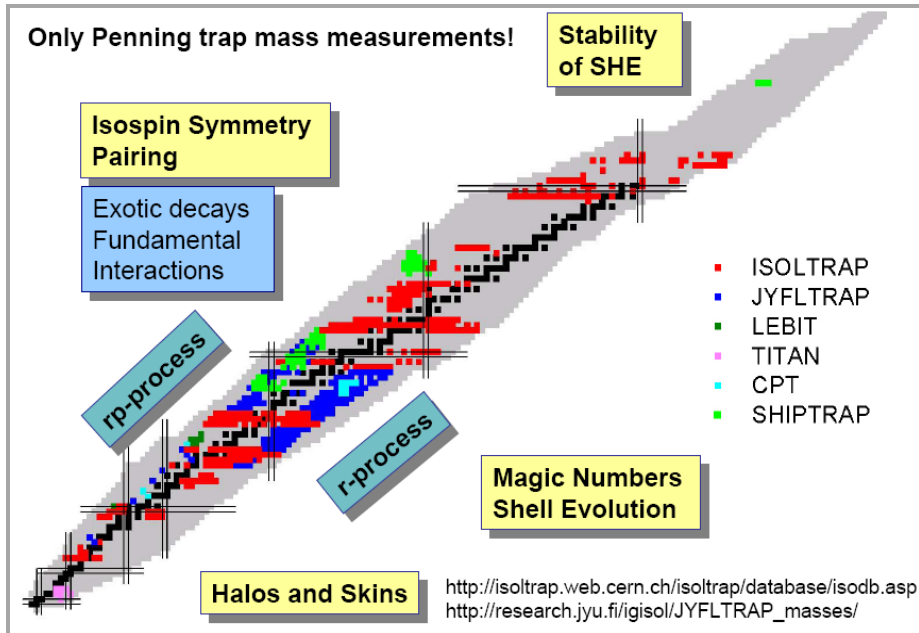
nuclear binding energy

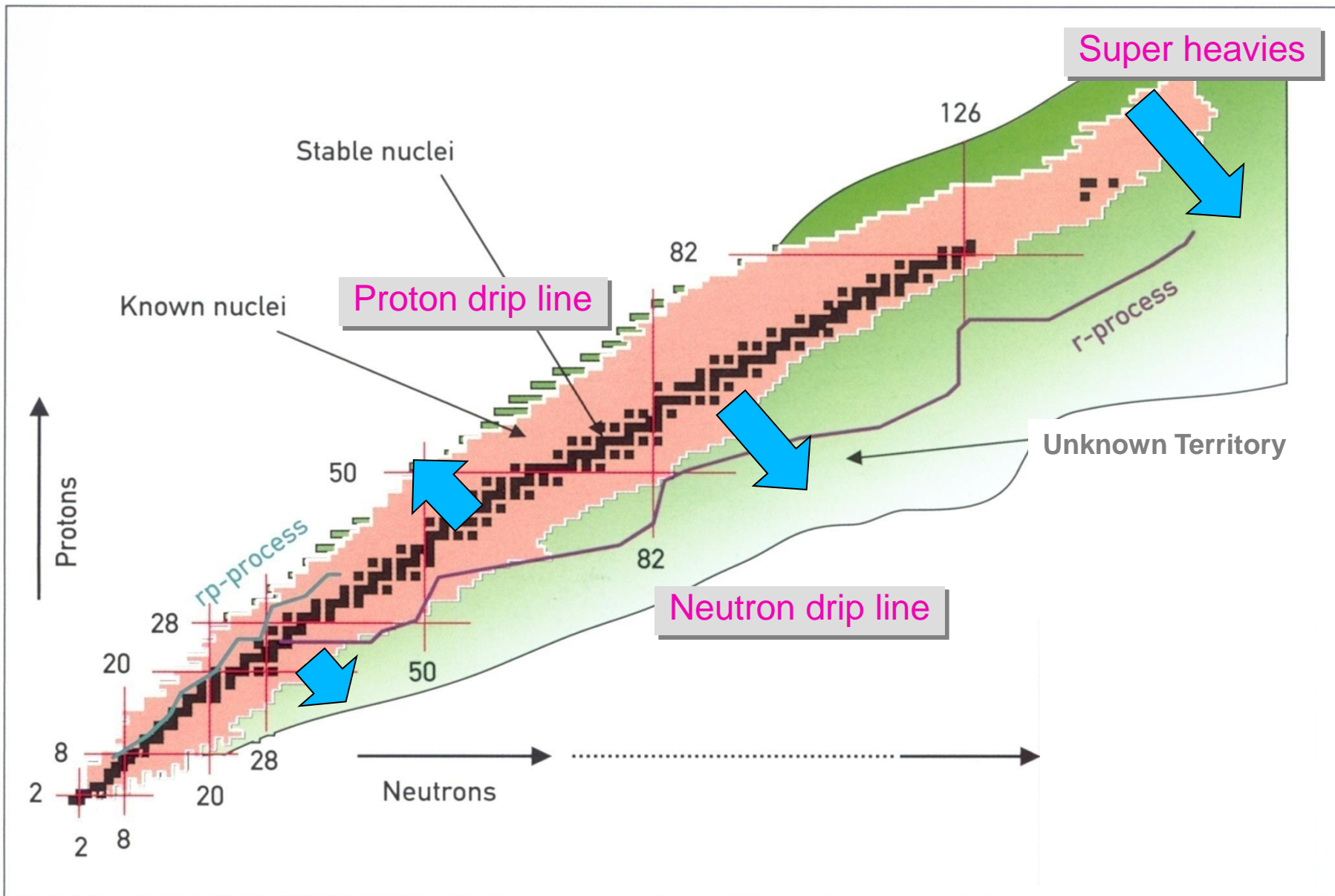
basic test of nuclear models

nuclear structure: shell closures, pairing, onset of deformation, drip lines, halos



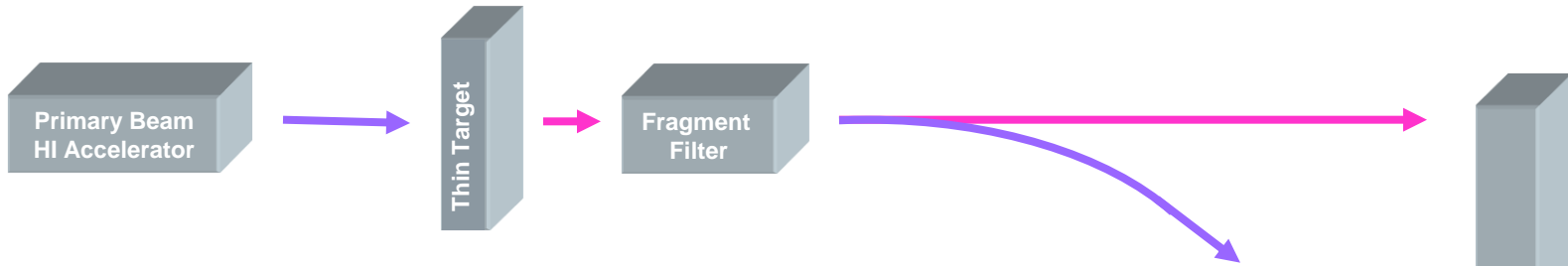
Mass measurements



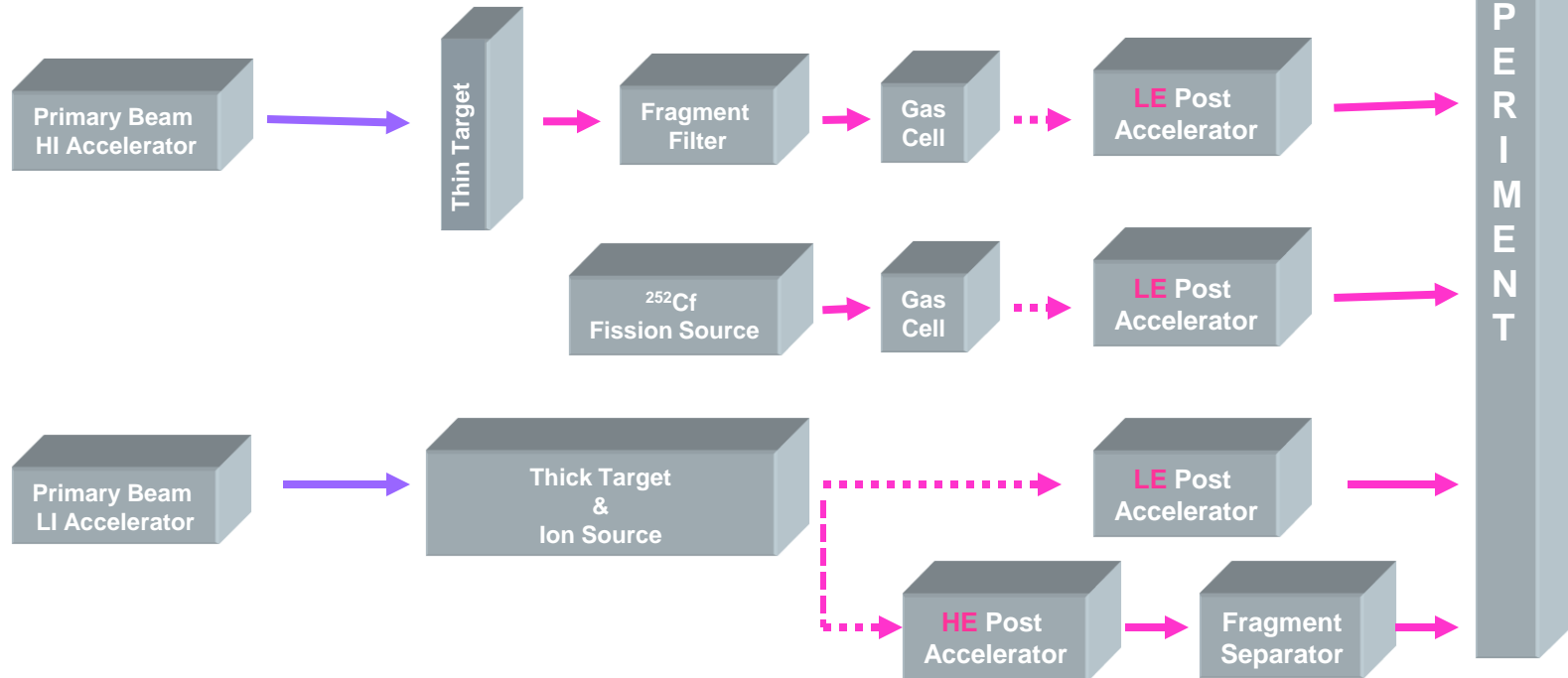


Production Methods for Radioactive Beams

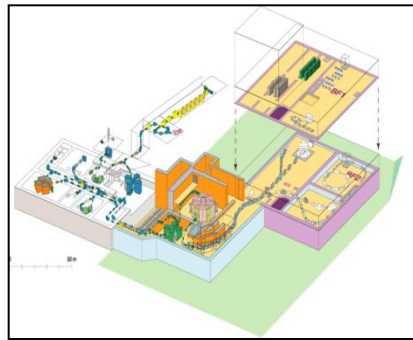
In-Flight Reactions



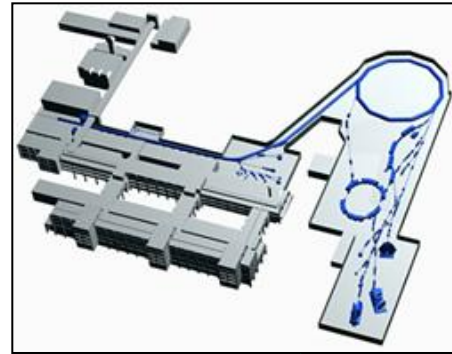
Stopping & Re-acceleration (ISOL)



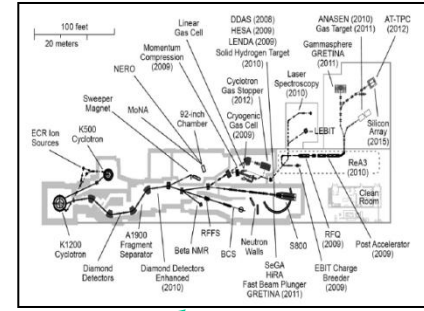
RIKEN-RIBF



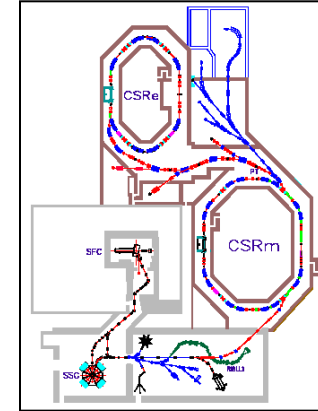
GSI-FRS/SIS18



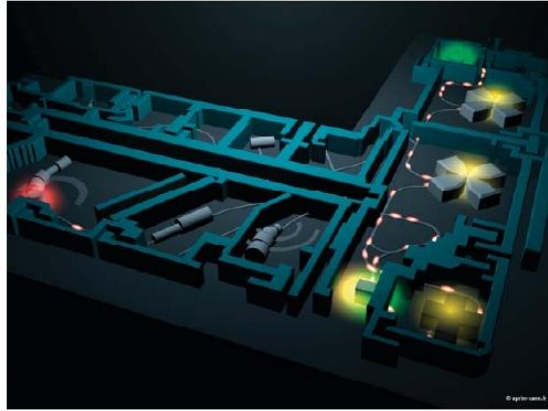
MSU-NSCL



CSR-Lanzhou



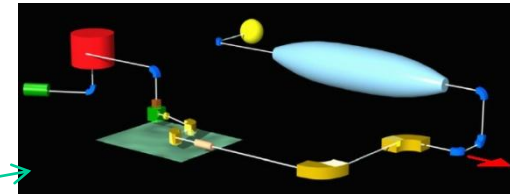
GANIL



In-Flight

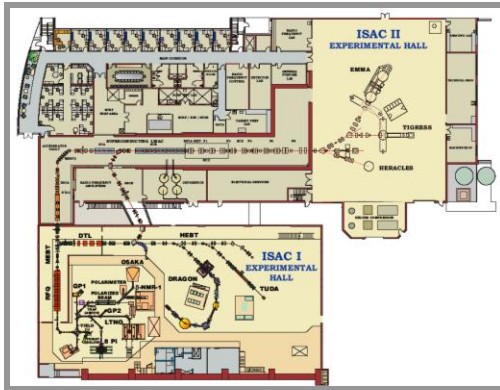
OPERATING

ISOL

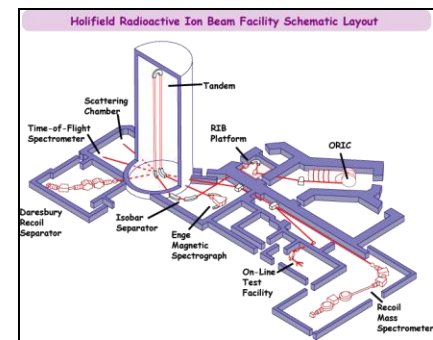
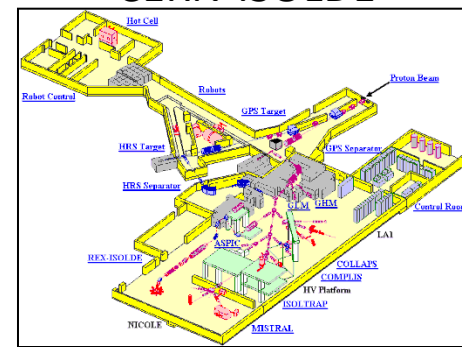


EXCYT-Catania

SPIRAL



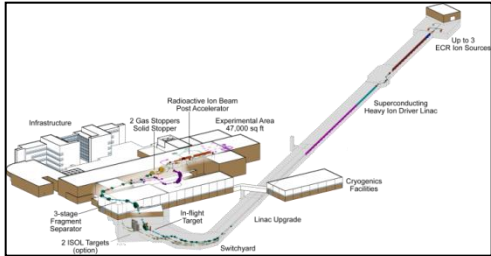
CERN-ISOLDE



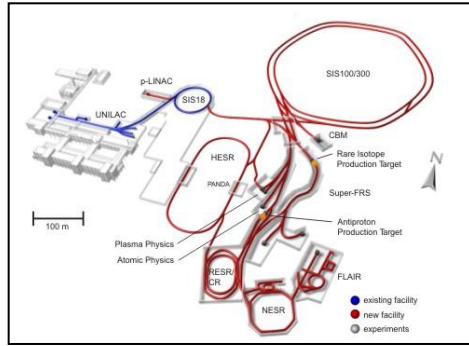
ORNL-HIRF

TRIUMF ISAC-2

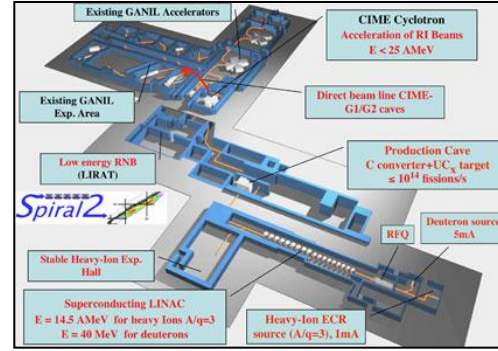
FRIB-MSU



FAIR-GSI



SPIRAL2-GANIL

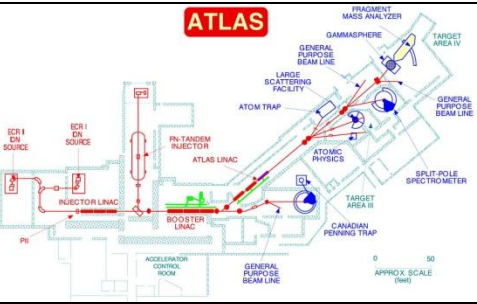


SPES-Legnaro

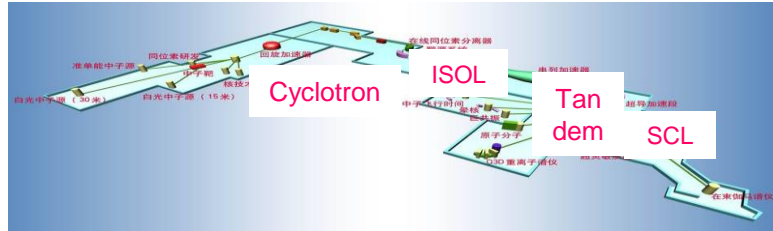
An ISOL RIB Facility is planned for construction: SPES



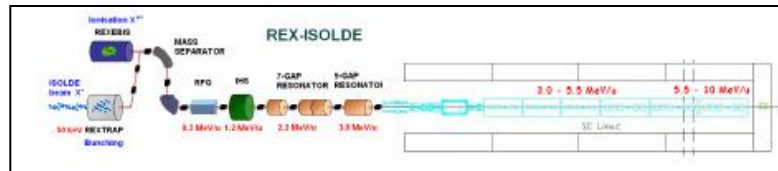
UNDER CONSTRUCTION
or
Funding Approved



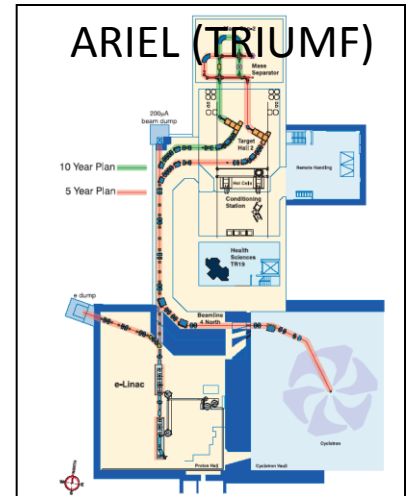
CARIBU (Argonne)



BRIF – Beijing)



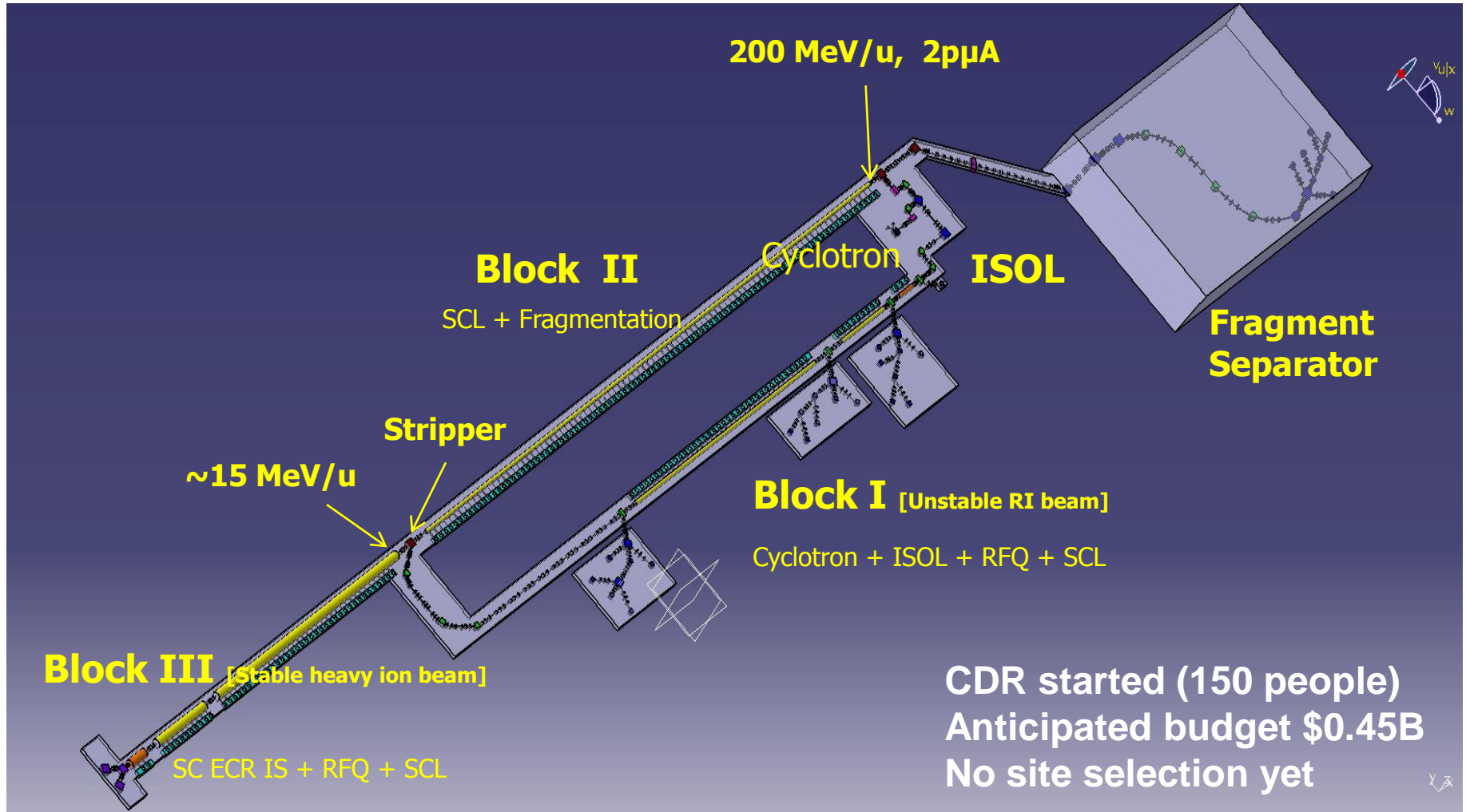
HIE-ISOLDE (CERN)

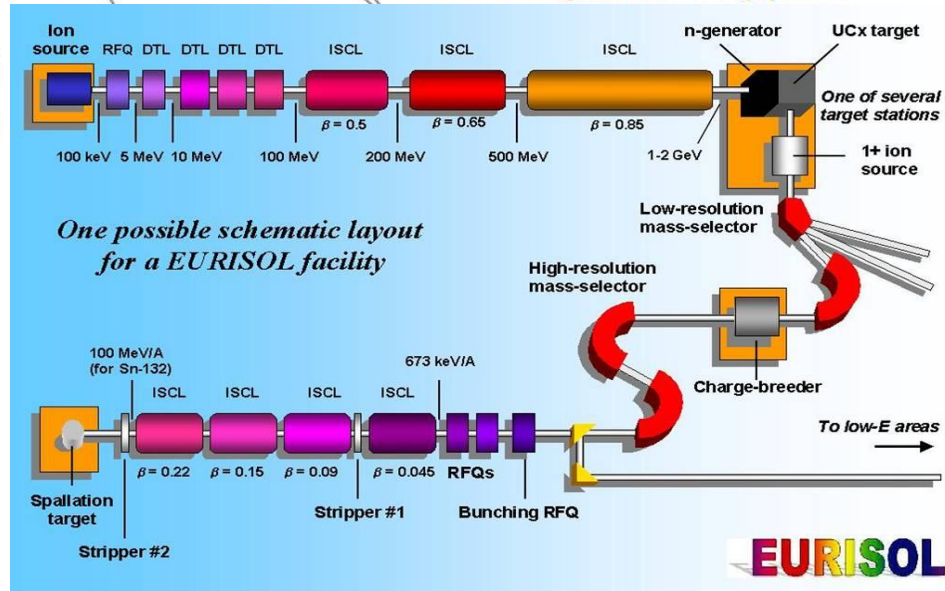
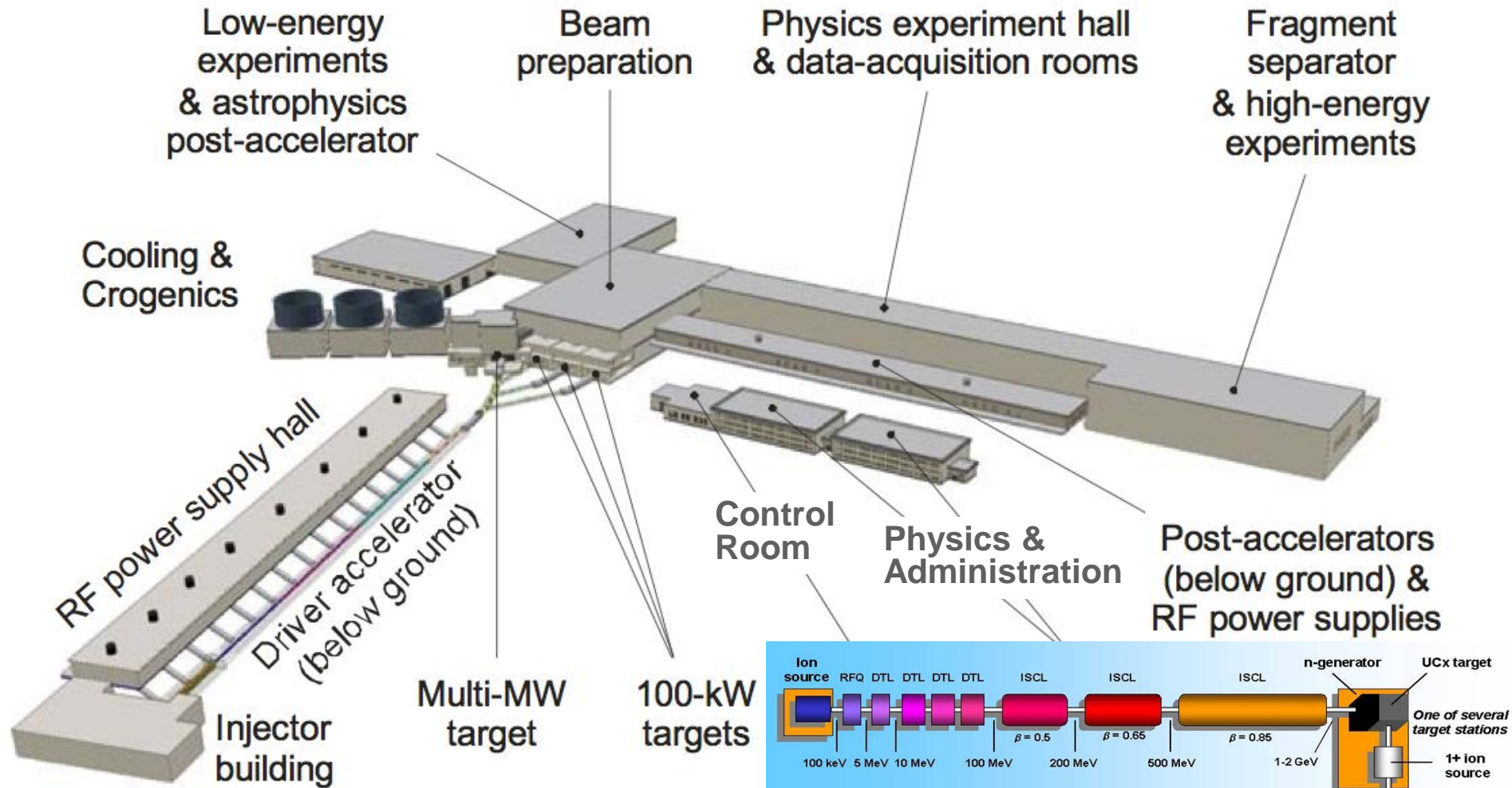


KoRIA (Korean Rare Isotope Accelerator)

Driver cyclotron (H 70MeV 1mA and D 35MeV 0.1mA)

Phase 1 (ISOL to 15MeV/u) / **Phase 2** (Linac to 200MeV/u for fragmentation beams)



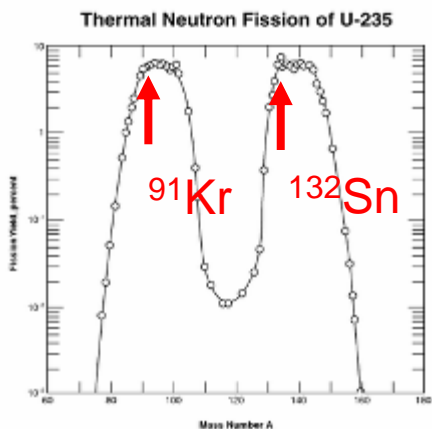
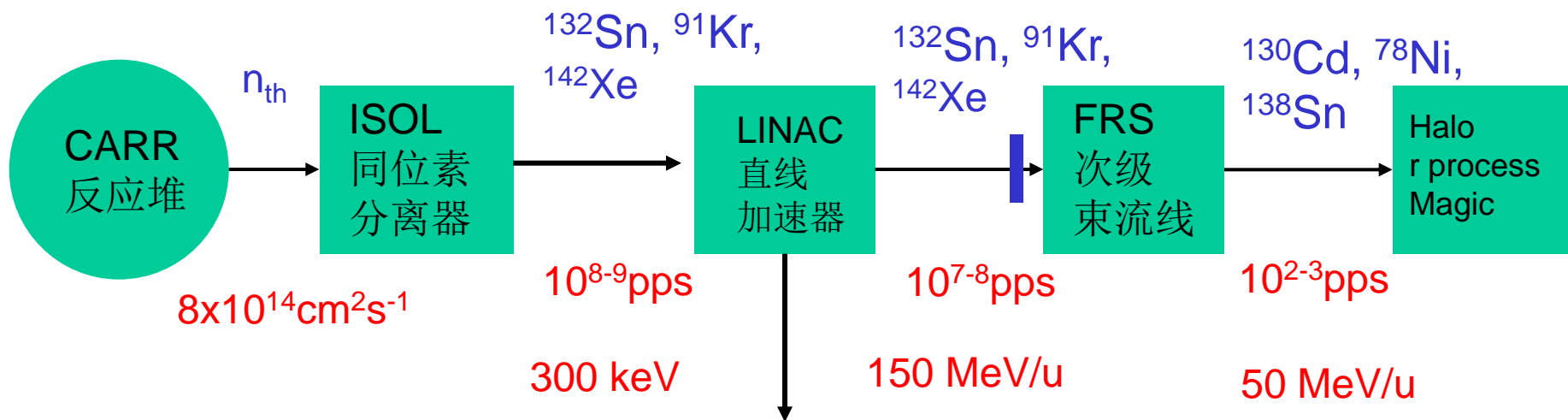


Enormous amount of work done in EU funded Design Study (Euro 30M)
<http://www.eurisol.org>
 Cost estimate €1.3B.
 Possible first beam 2015?

The CARIF Concept

China Advanced Rare Ion-beam Facility

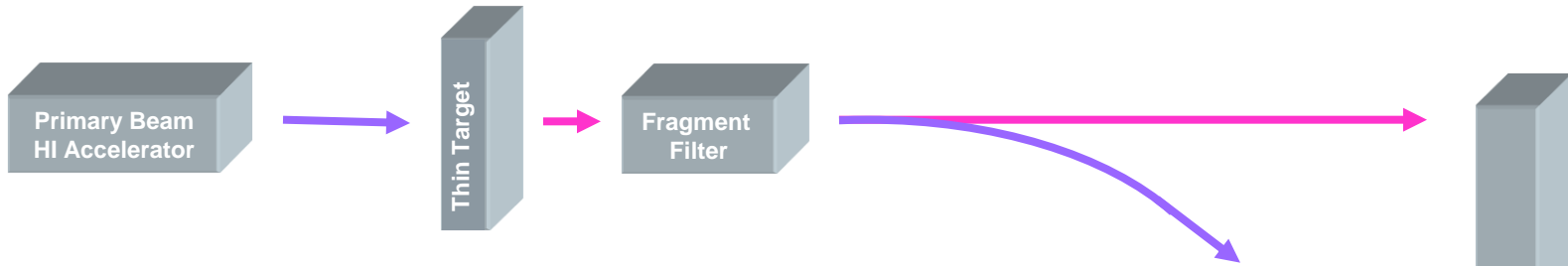
I. Tanihata, NIM B266 (2008) 4067



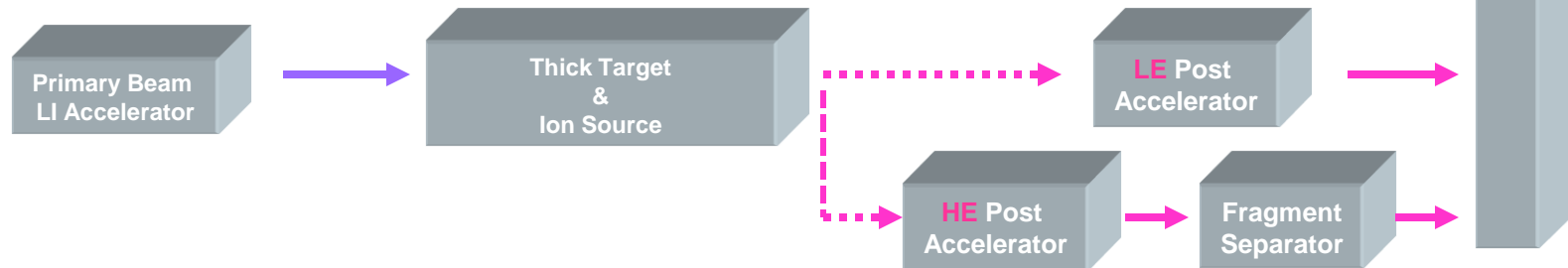
Expectation: Fragmentation of fission products can result in highly n-rich beams

Production Methods for Radioactive Beams

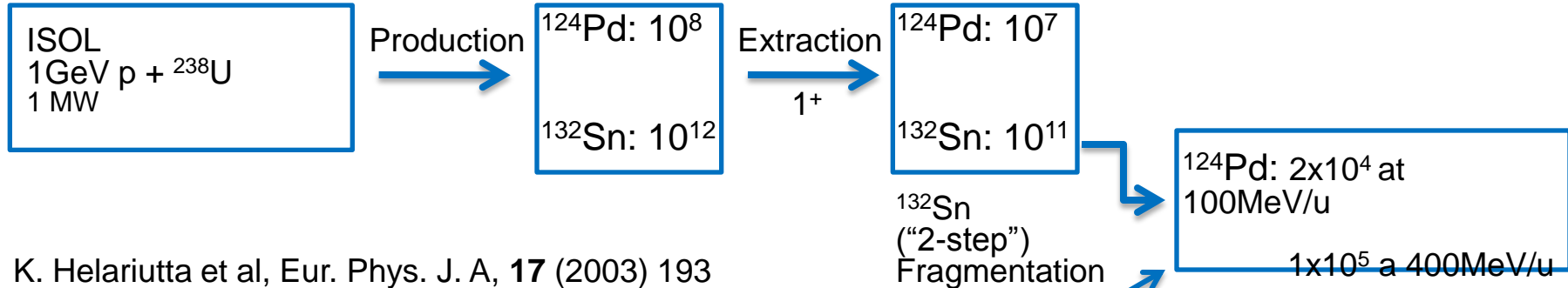
In-Flight Reactions



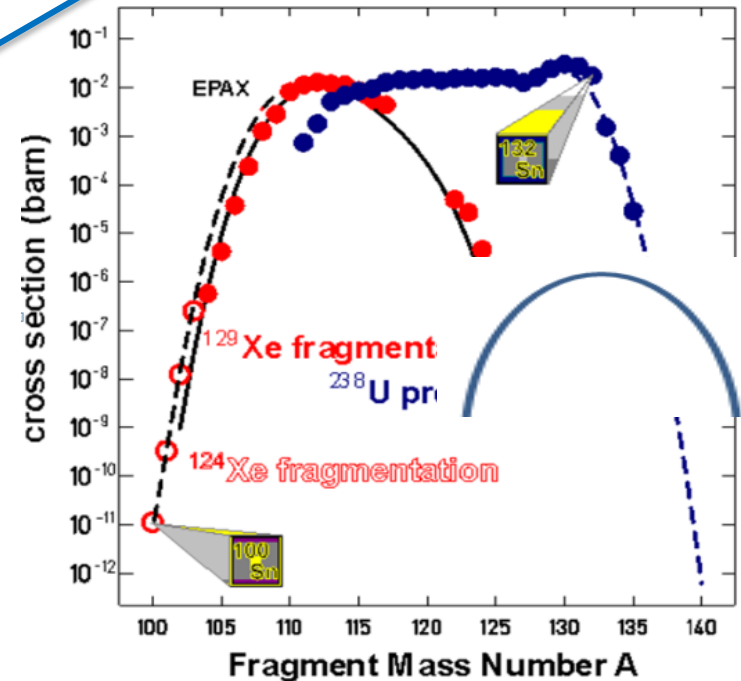
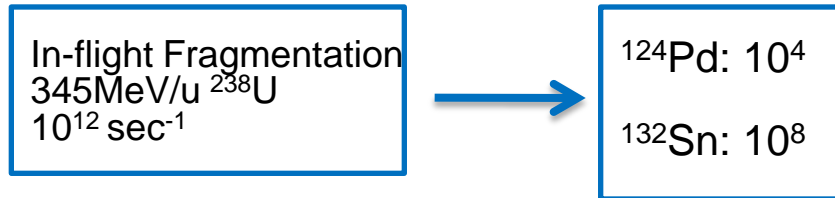
Stopping & Re-acceleration (ISOL)



Some comments on 2-step RIB Production

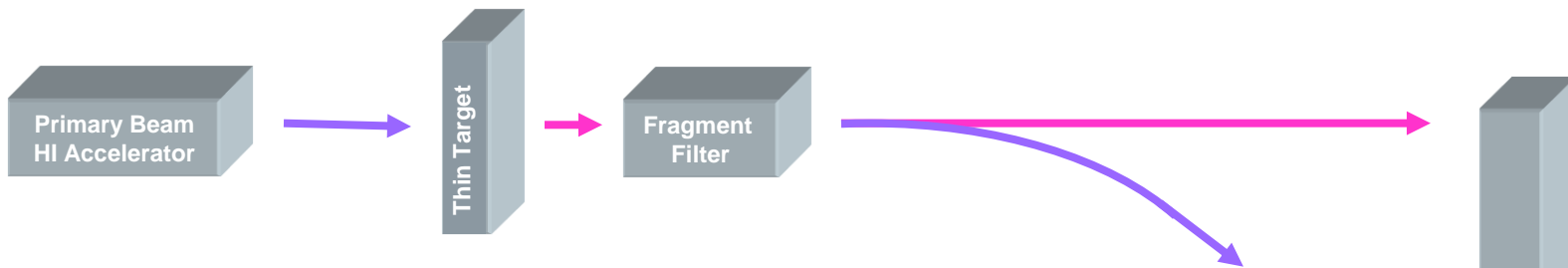


K. Helariutta et al, Eur. Phys. J. A, **17** (2003) 193

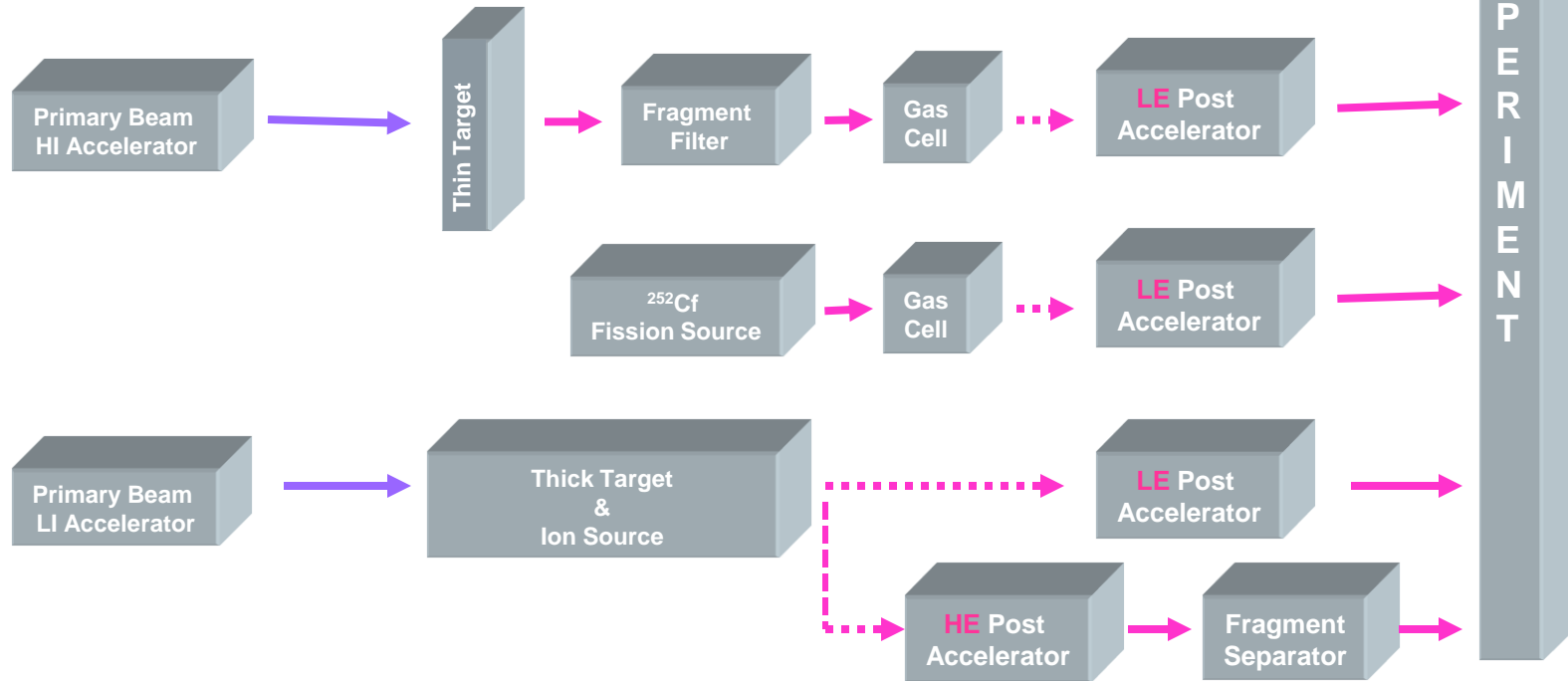


Production Methods for Radioactive Beams

In-Flight Reactions

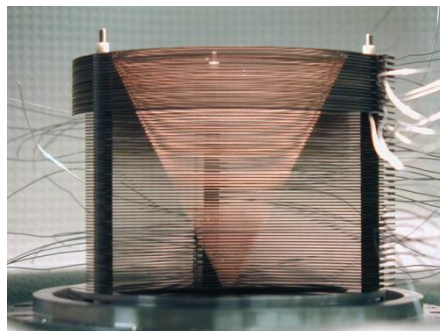


Stopping & Re-acceleration (ISOL)



Gas catcher at the ATLAS Canadian Penning Trap

- 20 cm long gas cell with second generation RF cone

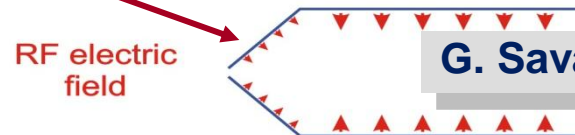
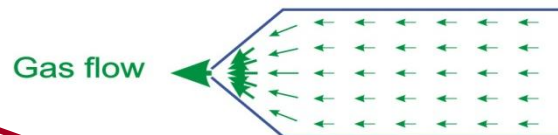
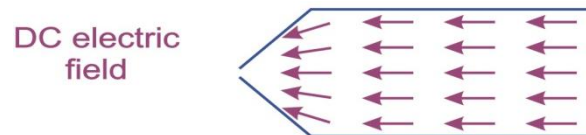


- $\epsilon \sim 45\%$

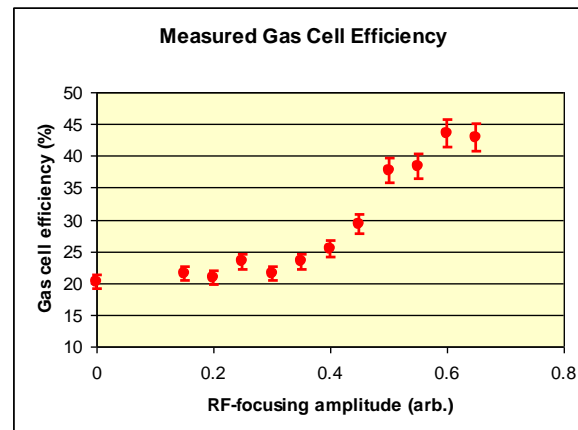
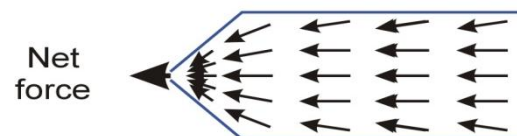
- mean delay time below 10 ms

- demonstrated off-line with fission products and on-line with fusion–evaporation reactions

- several years of operating experience

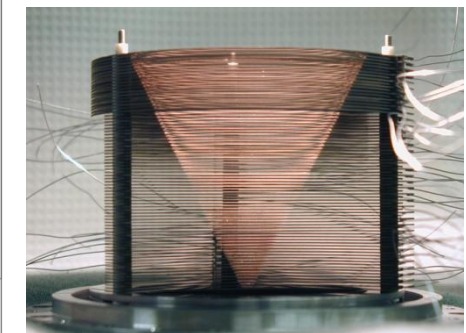
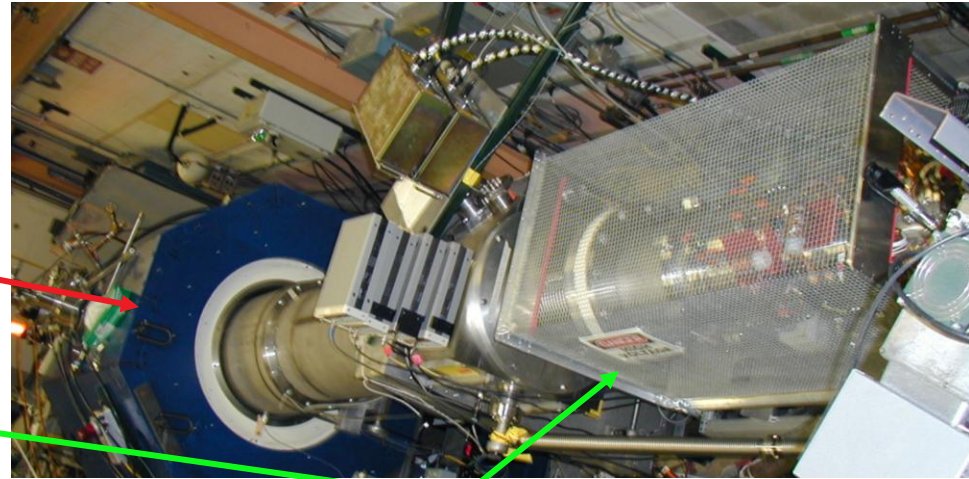


G. Savard 1997

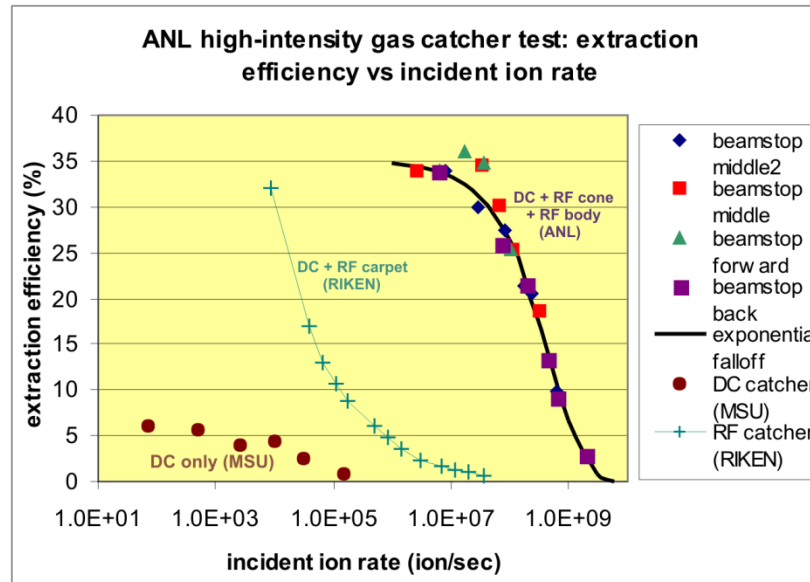


Gas catcher operation at FRIB intensities

- High intensity beamline built at ANL to reproduce ion distribution expected behind RIA/AEBL fragment separator
- Focusing by 68 cm bore superconducting solenoid
- Large gas catcher using large RF cone developed for high energy test at GSI, combined with full body RF focusing



- First tests at high intensity beamline in Sept and Oct 06
- High efficiency obtained at up to 10^9 incoming particles per second



What species have been extracted so far

Gas catcher should be a **“universal” approach** ... how diverse are the species that have been extracted so far:

- **Neutron-deficient isotopes**

- **Cs, Xe, Te, Sb, Sn, In, Rh, Ru, Tc, Mo, Se, As, Ge, Ga, As, Zn, Co, Fe, V, Ti, Al, Mg, Na, O, C**

- **Neutron-rich isotopes**

- **Nd, Ce, Pr, La, Ba, Rh, Ru, Te, Tc, Mo, Sr**

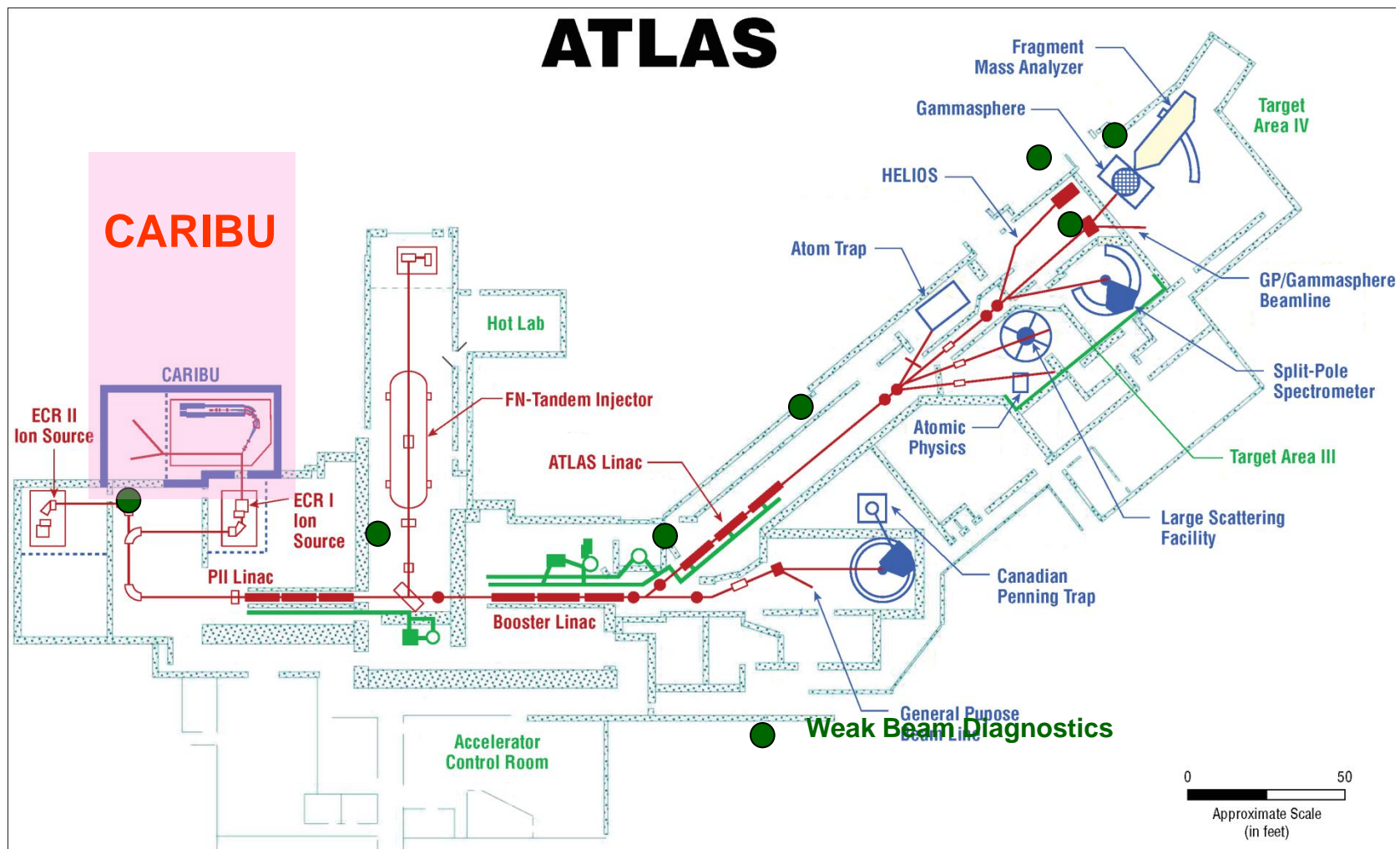
- **Stable isotopes**

- **Xe, Kr, Ar, Ne**

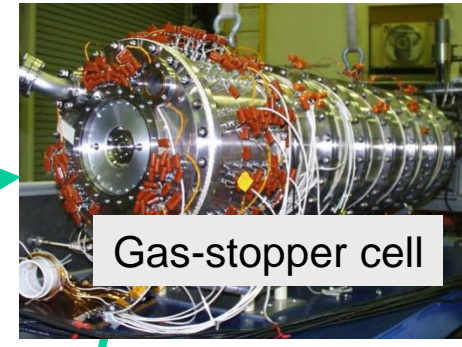
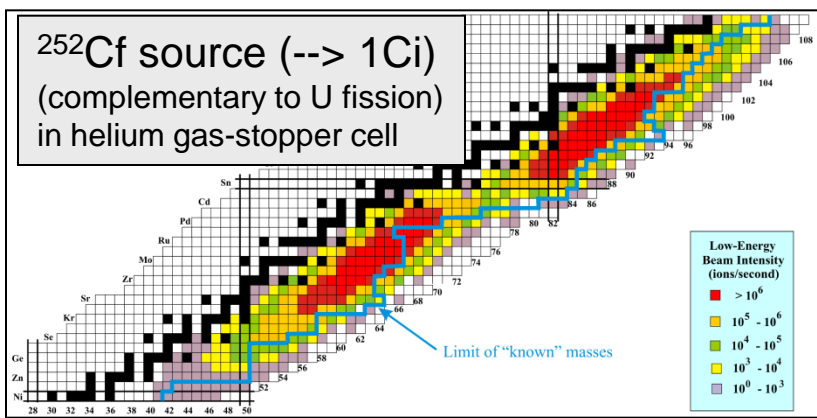
Essentially all attempted species, from the “easy” (alkali atoms and noble gases) up to the very refractory cases (Mo, Tc, Co, etc ...), have been extracted **as singly or doubly charged ions**.

All with **good efficiency**.

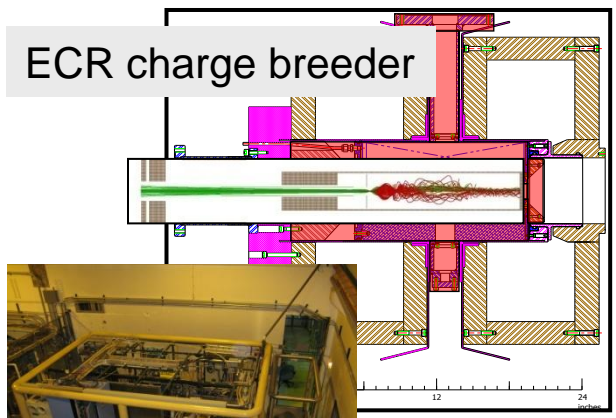
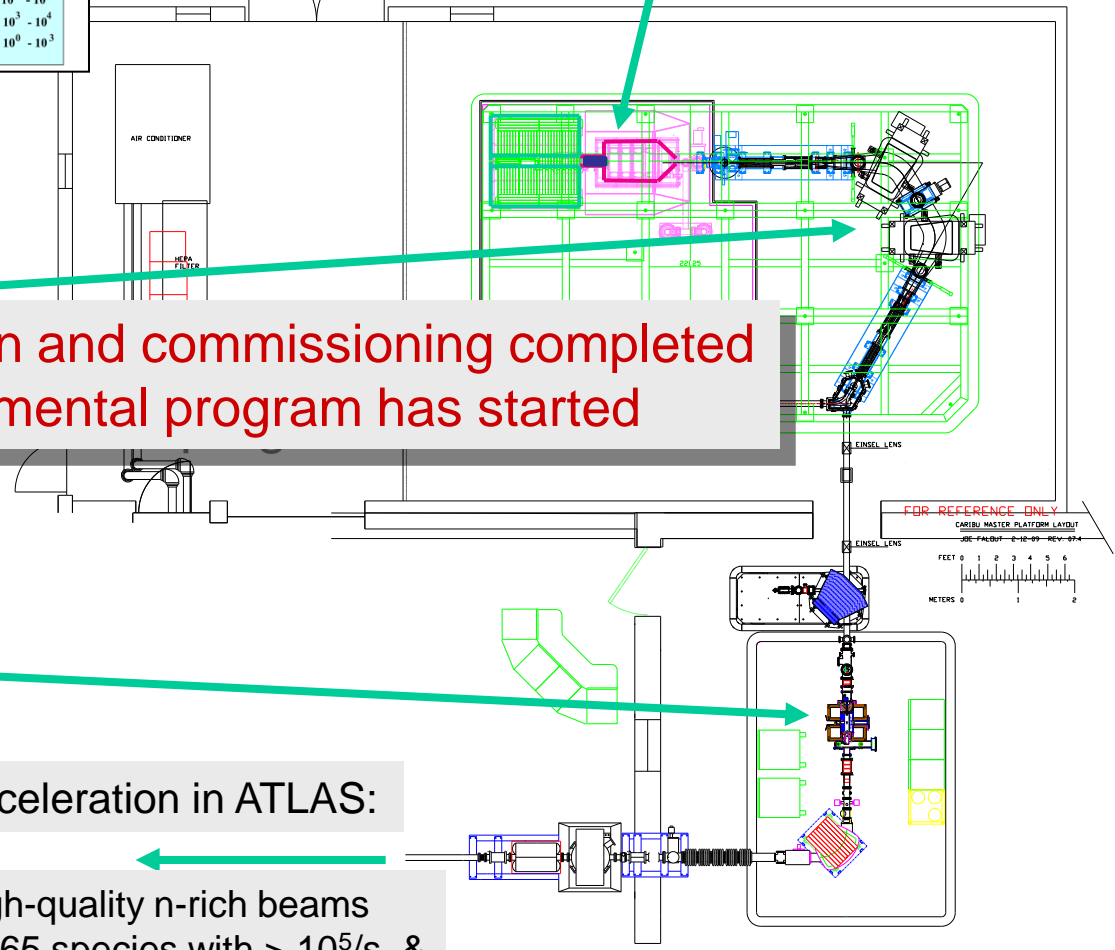
Layout for ^{252}Cf fission source system at ATLAS



^{252}Cf source (\rightarrow 1Ci)
(complementary to U fission)
in helium gas-stopper cell



Construction and commissioning completed
Experimental program has started



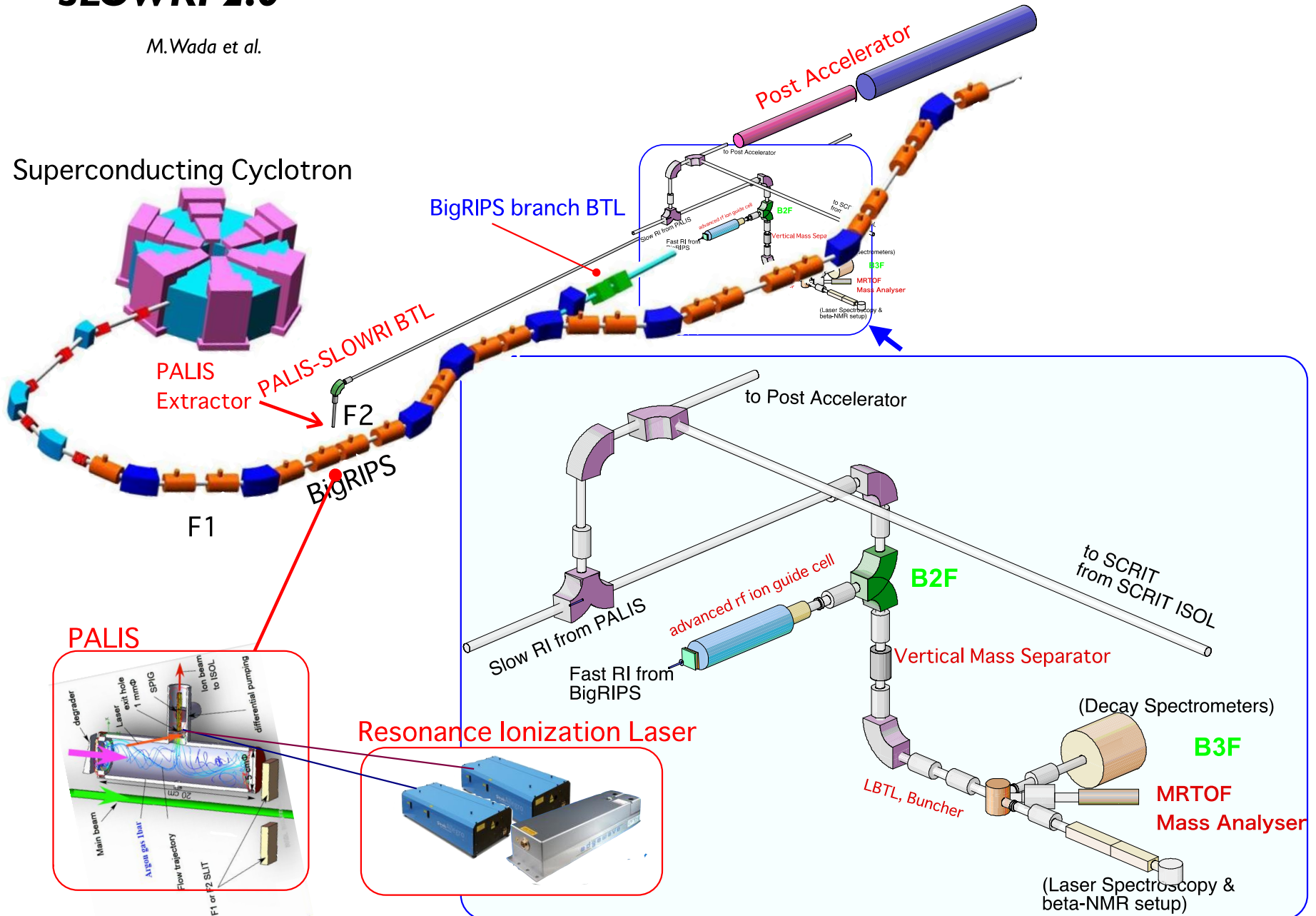
Acceleration in ATLAS:

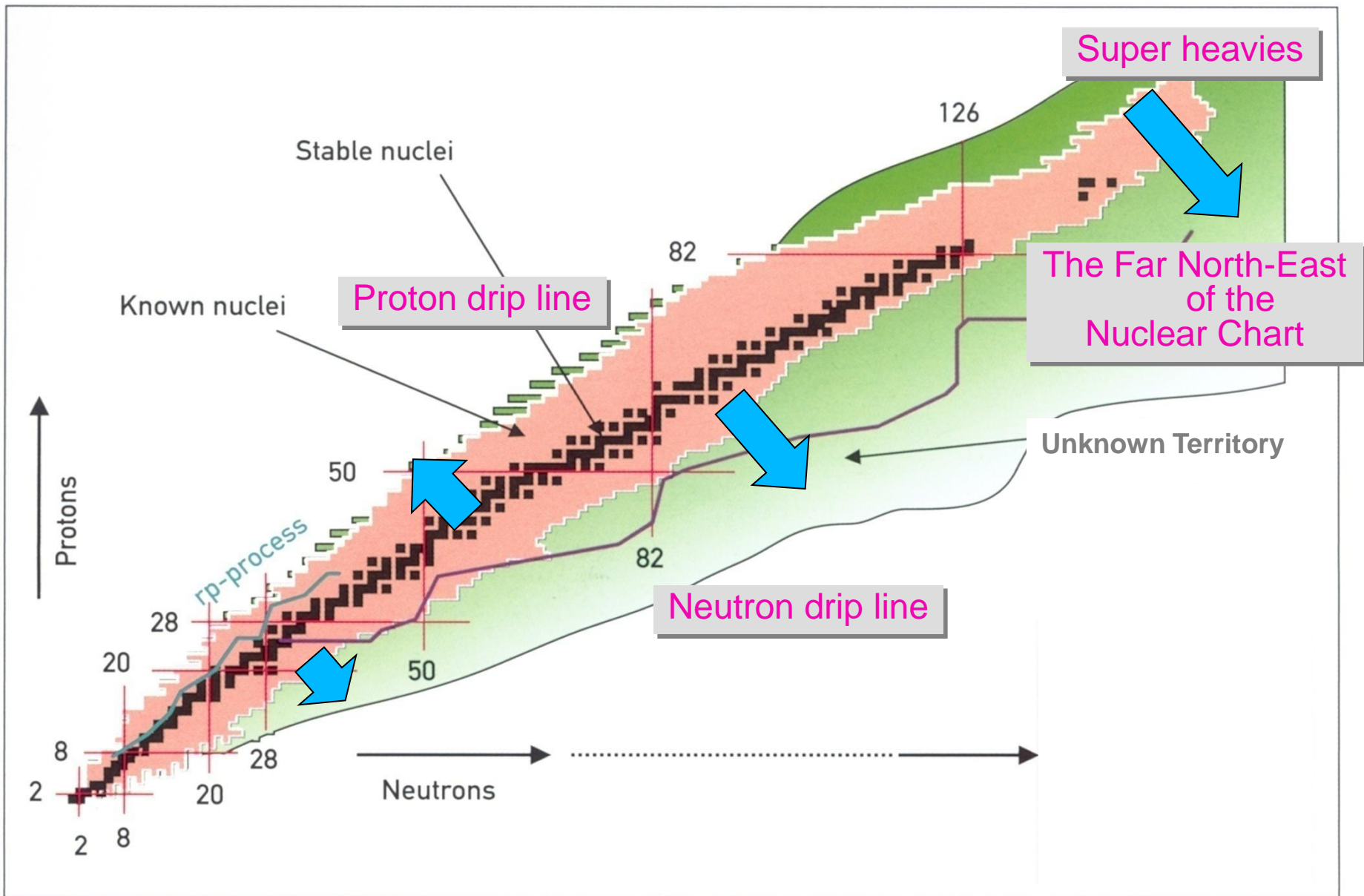
High-quality n-rich beams
~ 65 species with $> 10^5/s$ &
 > 150 species with $> 10^4/s$

SLOWRI 2.0

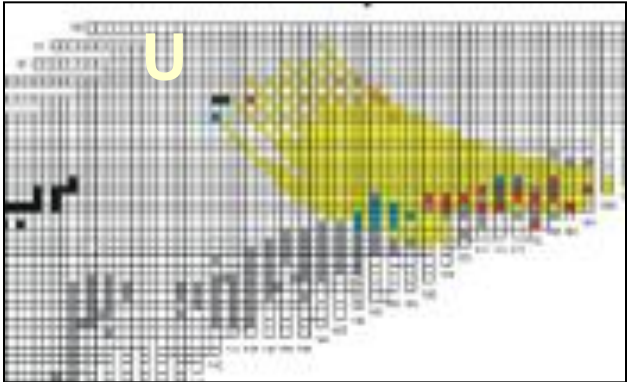
M.Wada et al.

Superconducting Cyclotron





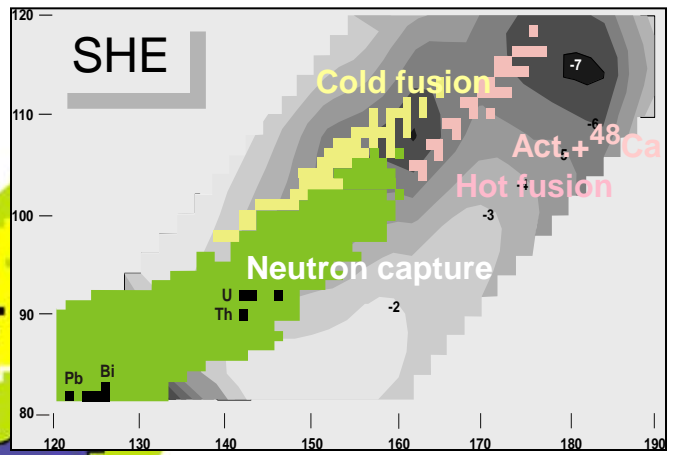
Proton number Z



r-Process production of ^{232}Th and ^{238}U -- and fission termination

^{208}Pb

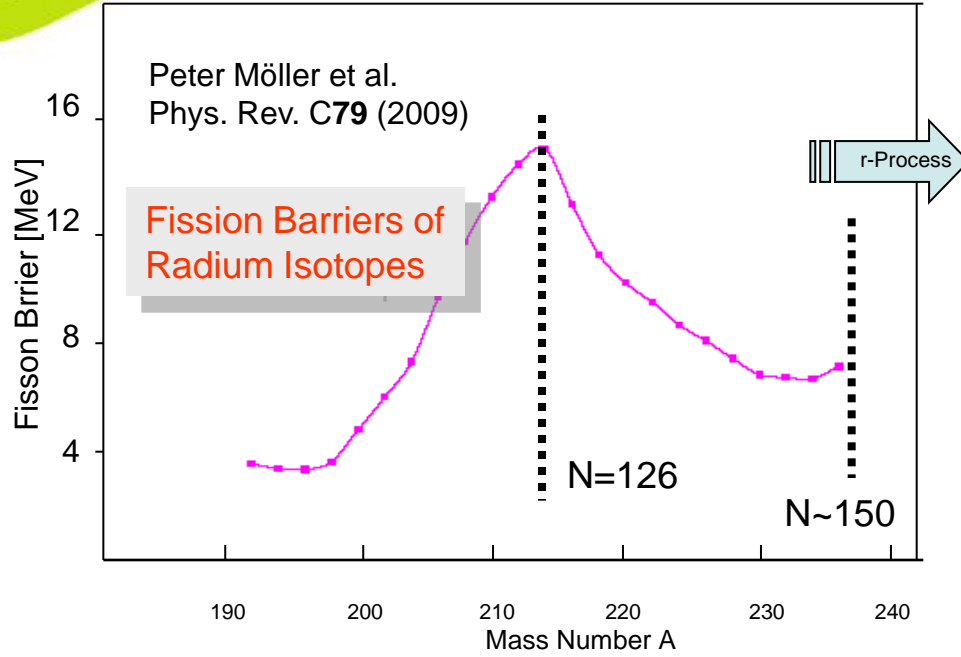
^{238}U



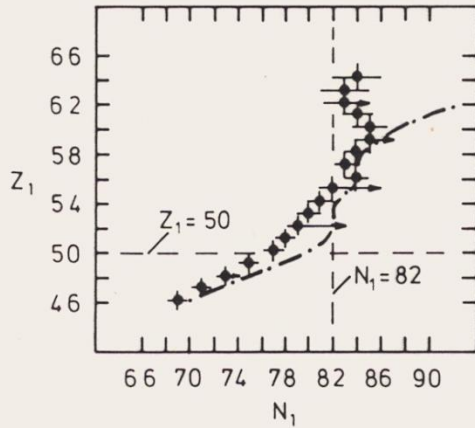
Fission of Neutron-Rich Nuclei

The Science:

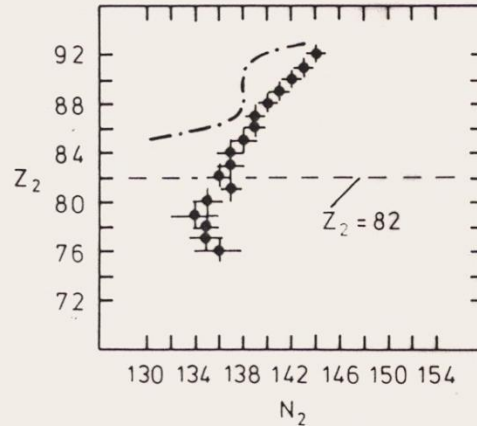
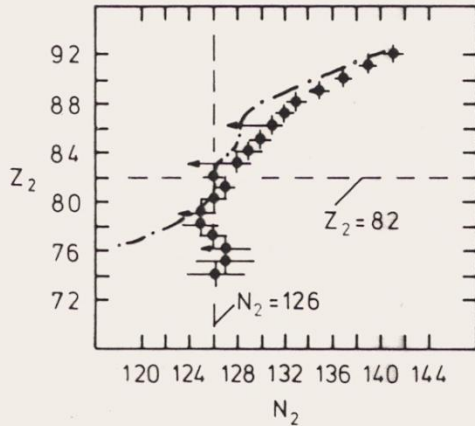
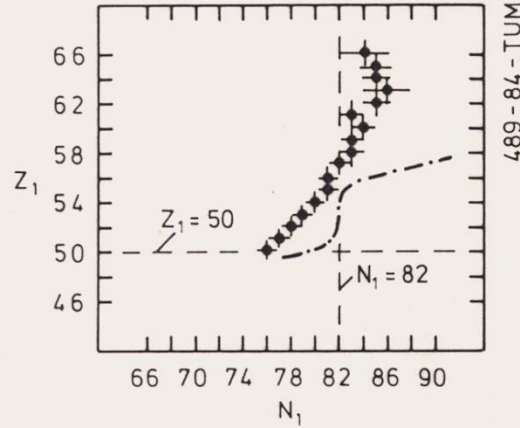
- nuclear shell structure at the extremes
- r-process: fission recycling and termination
- fission barriers of neutron-rich nuclei and symmetry energy
- connection of hot-fusion SHE island and mainland
- Unusual GT vs 1st forbidden decay ratios



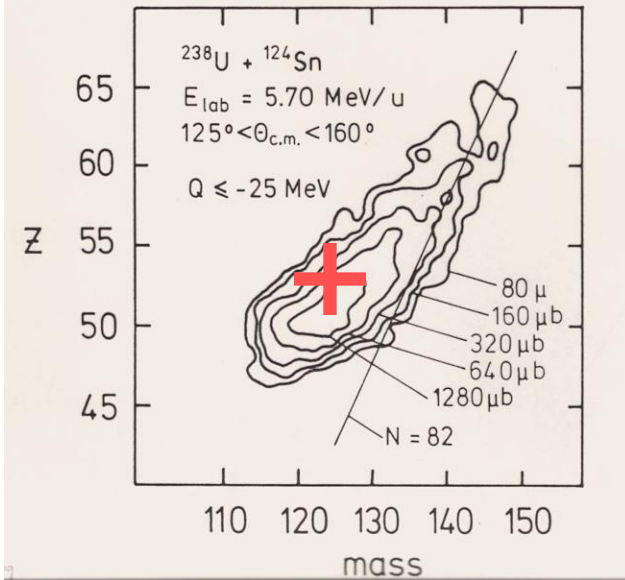
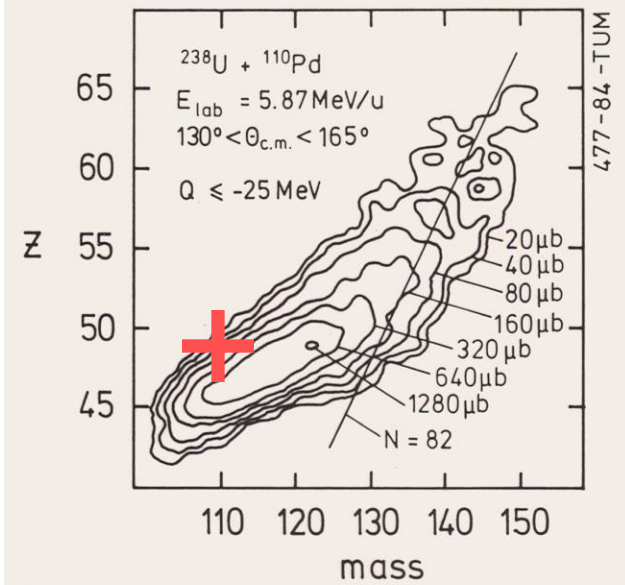
$^{238}\text{U} + ^{110}\text{Pd}$
 $E_{\text{lab}} = 5.87 \text{ MeV/u}$
 $130^\circ < \theta_{\text{c.m.}} < 165^\circ$
 $Q \leq -25 \text{ MeV}$



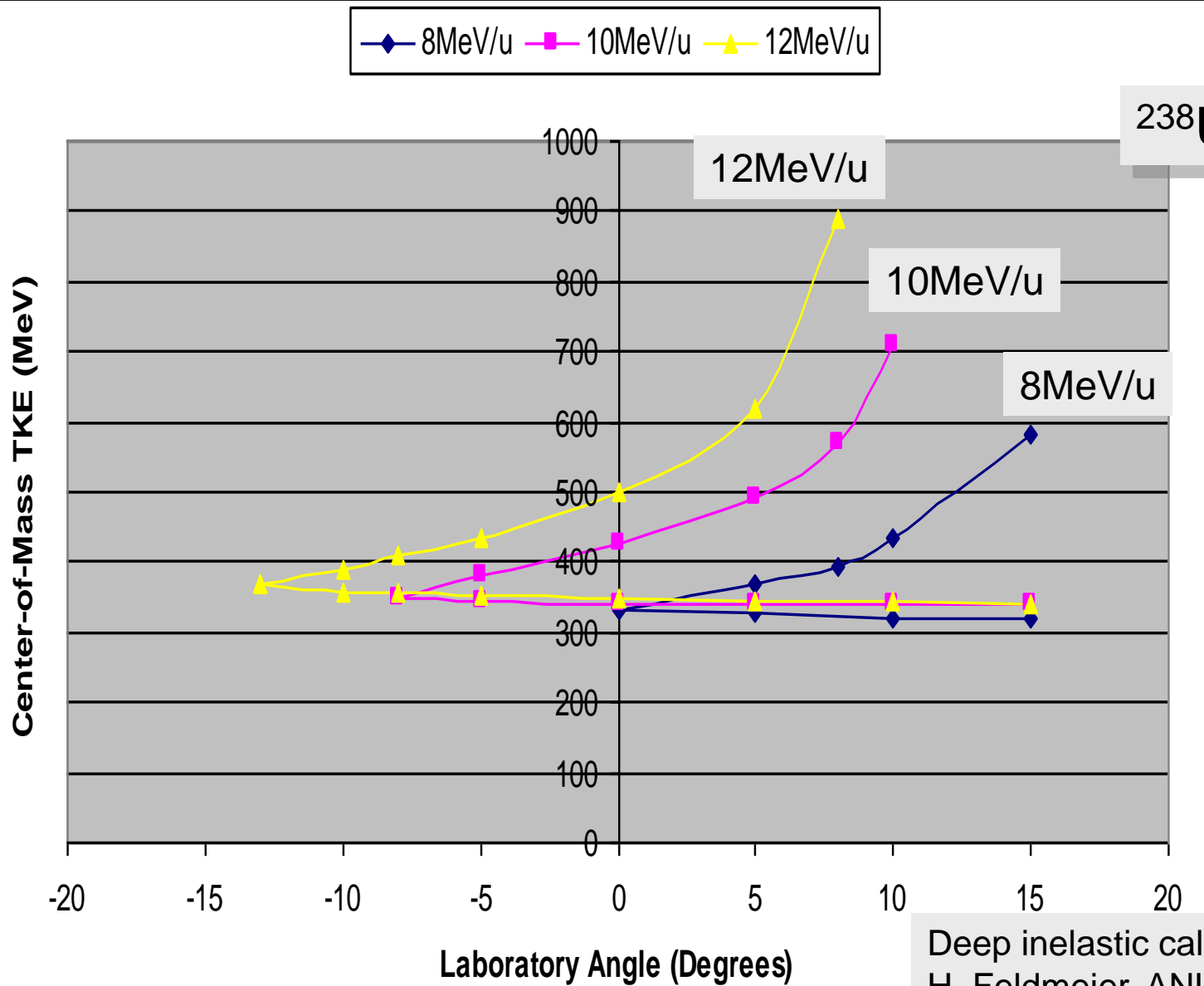
$^{238}\text{U} + ^{124}\text{Sn}$
 $E_{\text{lab}} = 5.70, 5.87, 6.02 \text{ MeV/u}$
 $125^\circ < \theta_{\text{c.m.}} < 160^\circ$
 $Q \leq -25 \text{ MeV}$



Waltraud Mayer et al., Phys. Lett. **B115** (1985) 162



$^{238}\text{U} + ^{110}\text{Pd}$



Deep inelastic calculations with code
H. Feldmeier, ANL-PHY-85-2 (1985)

A long, straight road stretches into the distance under a blue sky, with mountains in the background. The road is flanked by dry grass and gravel shoulders. The sky is a deep, clear blue, and the mountains in the distance are also blue, creating a monochromatic effect. The road has a yellow dashed line down the center and white lines on the sides. There are some small structures and trees on the right side of the road in the distance.

Extremes of the (nuclear) landscape...

Exploring the unknown...

Many facets of new physics...

RIBF is well positioned for the future...

Thank you for your attention!