## Out-of-equilibrium field theories coupled to strong external sources

Kyoto University, December 2013

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#### **Outline**



- Preamble : classical statistical method in Quantum Mechanics
- QFT with strong sources, Inclusive observables at LO and NLO
- 3 Instabilities and resummation
- 4 Example : Schwinger mechanism

# Quantum Mechanics

#### **Classical phase-space formulation of Quantum Mechanics**



Consider the von Neumann equation for the density operator :

$$\frac{\partial \widehat{\rho}_{\tau}}{\partial \tau} = i \frac{\hbar}{[\widehat{H}, \widehat{\rho}_{\tau}]} \qquad (**)$$

Introduce the Wigner transforms :

$$\begin{array}{lcl} {\cal W}_{\tau}(x,p) & \equiv & \int ds \; e^{i\,\mathbf{p}\cdot\mathbf{s}} \; \left\langle x+\frac{s}{2} \big| \widehat{\rho}_{\tau} \big| x-\frac{s}{2} \right\rangle \\ & {\cal H}(x,p) & \equiv & \int ds \; e^{i\,\mathbf{p}\cdot\mathbf{s}} \; \left\langle x+\frac{s}{2} \big| \widehat{H} \big| x-\frac{s}{2} \right\rangle \; \; \mbox{(classical Hamiltonian)} \end{array}$$

#### (\*\*) is equivalent to:

$$\begin{array}{lcl} \frac{\partial W_{\tau}}{\partial \tau} & = & \mathfrak{H}(x,p) \; \frac{2}{\mathfrak{i} \, \hbar} \; \sin \left( \frac{\mathfrak{i} \, \hbar}{2} \left( \stackrel{\leftarrow}{\partial}_{\, p} \stackrel{\rightarrow}{\partial}_{\, x} - \stackrel{\leftarrow}{\partial}_{\, x} \stackrel{\rightarrow}{\partial}_{\, p} \right) \right) \; W_{\tau}(x,p) \\ & = & \underbrace{\left\{ \mathfrak{H}, W_{\tau} \right\}}_{\text{Poisson bracket}} \; + \mathcal{O}(\hbar^2) \end{array}$$

#### **Classical statistical method in Quantum Mechanics**

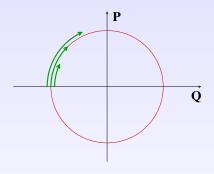


- Quantum effects in the time evolution are O(ħ²) corrections (i.e. they appear at NNLO and beyond)
- ①(ħ) (NLO) contributions can only come from the initial state
   Uncertainty principle : Δx · Δp ≥ ħ
   The initial Wigner distribution W<sub>τ=0</sub>(x, p) must have a support of area at least ħ (minimal area realized by coherent states)
- All the O(ħ) effects can be accounted for by a Gaussian initial distribution W<sub>τ=0</sub>(x, p)

#### Classical statistical method

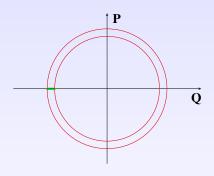
- Sample by a Monte-Carlo the Gaussian distribution that approximates the initial distribution  $W_{\tau=0}(\mathbf{x},\mathbf{p})$
- For each initial (x, p), solve the classical equation of motion up to the time of interest





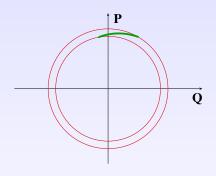
• For non-harmonic oscillators, the oscillation frequency depends on the initial condition





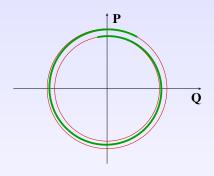
- For non-harmonic oscillators, the oscillation frequency depends on the initial condition
- $\bullet$  Because of QM, the initial ensemble is a set of width  $\gtrsim \! \hbar$





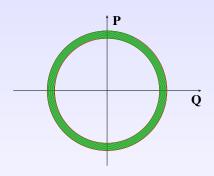
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- This ensemble of initial configurations spreads in time





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- For non-harmonic oscillators, the oscillation frequency depends on the initial condition
- Because of QM, the initial ensemble is a set of width  $\gtrsim \hbar$
- This ensemble of initial configurations spreads in time
- At large times, the ensemble fills densely all the region allowed by energy conservation ⇒ microcanonical equilibrium

## **Quantum Field Theory**

w/ Strong Sources

#### Typical situation

$$\frac{1}{2}(\vartheta_{\mu}\varphi)(\vartheta^{\mu}\varphi)-V(\varphi)+J\varphi$$

- In general, the source J is space and time dependent
- The system starts at  $t = -\infty$  from a known initial state (example: vacuum state)
- The source may be turned off at some point, and the system evolves by itself afterwards

#### Strong source : J ~ inverse coupling

- Non-perturbative
  - ⇒ can we expand in powers of the coupling?
- Spectrum of produced particles?
- After the sources are switched off: how does the system equilibrate?

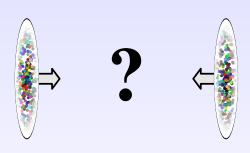
#### **Example I: Schwinger mechanism**



- Consider a constant and uniform electrical field  $\vec{E}_{\text{ext}}$
- Perturbatively, energy conservation prevents the production of  $e^+e^-$  pairs
- Pairs can be produced via a vacuum instability
- Rate :  $\exp(-\pi m^2/eE)$  (non analytic in the coupling *e*)
- In Quantum Field Theory, can be obtained at one loop (but one must use the propagator dressed by the external field)

#### **Example II: Nucleus-Nucleus collisions at high energy**



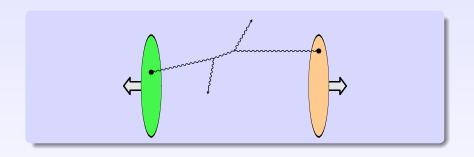


$$\mathcal{L} = -\frac{1}{4} \; F_{\mu\nu} F^{\mu\nu} + (\underbrace{J_1^\mu + J_2^\mu}_{J^\mu}) A_\mu \label{eq:lambda}$$

- Given the sources J<sub>1,2</sub> in each projectile, how do we calculate observables? Is there some kind of perturbative expansion?
- Loop corrections, factorization?
- Thermalization?

#### Strong source regime



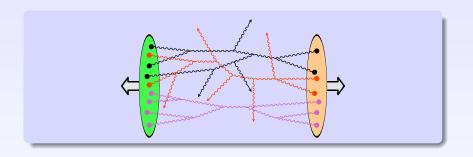


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· Weak sources : perturbative treatment

#### Strong source regime



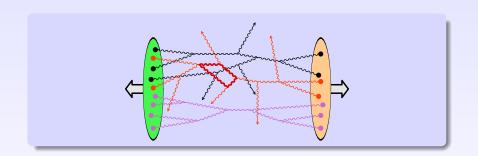


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- · Weak sources : perturbative treatment
- Strong sources: non-perturbative (when J ~ 1/g)

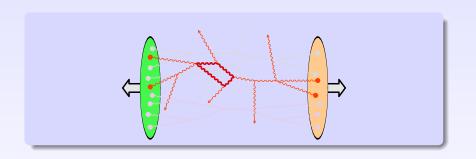
### **Power counting**





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### Order of connected subdiagram when $J \sim g^{-1}$ :

$$\frac{1}{g^2}$$
 g<sup>#</sup> produced gluons  $g^{2(\# loops)}$ 

#### **Power counting**



Example : single particle spectrum :

$$\frac{dN_1}{d^3\vec{p}} = \frac{1}{g^2} \left[ c_0 + c_1 g^2 + c_2 g^4 + \cdots \right]$$

 The coefficients c<sub>0</sub>, c<sub>1</sub>, · · · are themselves series that resum all orders in (gJ)<sup>n</sup>. For instance,

$$c_0 = \sum_{n=0}^{\infty} c_{0,n} \left( \mathbf{gJ} \right)^n$$

 We want to calculate at least the entire c<sub>0</sub>/g<sup>2</sup> contribution, and a subset of the higher order terms

#### Inclusive observables



 Inclusive observables do not veto any final state Example: moments of the transition probabilities:

$$\frac{dN_1}{d^3 \vec{\boldsymbol{p}}} \sim \sum_{n=0}^{\infty} (n+1) \int \frac{1}{(n+1)!} \bigg[\underbrace{d\Phi_1 \cdots d\Phi_n}_{\text{n part. phase-space}}\bigg] \ \left| \left\langle \boldsymbol{p} \boldsymbol{p}_1 \cdots \boldsymbol{p}_{n \, \text{out}} \middle| \boldsymbol{0}_{\text{in}} \right\rangle \right|^2$$

(single inclusive particle distribution)

#### **Equivalent definition:**

$$\frac{dN_1}{d^3\vec{\boldsymbol{p}}} \sim \left\langle \mathbf{0}_{in} \middle| \mathbf{\alpha}_{out}^\dagger(\boldsymbol{p}) \mathbf{\alpha}_{out}(\boldsymbol{p}) \middle| \mathbf{0}_{in} \right\rangle$$

(completeness of the out-states)





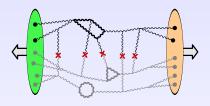
- Start with transition amplitudes : sources  $\rightarrow$  particles





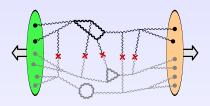
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#### Weight each cut by $z(p) \rightarrow generating functional$

$$F[\textbf{z}] \equiv \sum_{n} \frac{1}{n!} \int \left[ d\Phi_1 \cdots d\Phi_n \right] \frac{\textbf{z}(\textbf{p}_1) \cdots \textbf{z}(\textbf{p}_n)}{\left| \left\langle \textbf{p}_1 \cdots \textbf{p}_{n \, out} \middle| \textbf{0}_{in} \right\rangle \right|^2}$$

#### **Generating functional**



Observables are given by derivatives of F[z], e.g.

$$\frac{\mathrm{dN}_1}{\mathrm{d}^3\vec{\mathbf{p}}} = \left. \frac{\delta F[z]}{\delta z(\mathbf{p})} \right|_{z=1}$$

(inclusive observables are derivatives at the point z = 1)

unitarity implies F[1] = 1

#### Exact formula for the first derivative :

$$\frac{\delta \log F[z]}{\delta z(\textbf{p})} = \int d^4x d^4y \ e^{i\textbf{p}\cdot(\textbf{x}-\textbf{y})} \ \Box_\textbf{x} \Box_\textbf{y} \Big[ \textbf{A}_+(\textbf{x}) \textbf{A}_-(\textbf{y}) + \textbf{G}_{+-}(\textbf{x},\textbf{y}) \Big]$$

where  $A_{\pm}$  and  $g_{+-}$  are connected 1- and 2-point functions in the Schwinger-Keldysh formalism, with cut propagators weighted by z(p)

#### **Schwinger-Keldysh formalism**



- Set of Feynman rules to compute directly transition probabilities (i.e.  $\mathcal{A}\mathcal{A}^*$ )
- This can be achieved as follows:
  - ullet A vertex is -ig on one side of the cut, and +ig on the other side
  - There are four propagators, depending on the location w.r.t. the cut of the vertices they connect:

$$\begin{array}{ll} G_{++}^0(p)=i/(p^2-m^2+i\varepsilon) & (\text{standard Feynman propagator}) \\ G_{--}^0(p)=-i/(p^2-m^2-i\varepsilon) & (\text{complex conjugate of } G_{++}^0(p)) \\ G_{+-}^0(p)=2\pi\,\text{z}(\textbf{p})\,\theta(-p^0)\delta(p^2-m^2) \end{array}$$

At each vertex of a given diagram, sum over the types + and (2<sup>n</sup> terms for a diagram with n vertices)

# Inclusive Observables at Leading Order

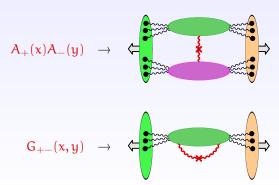
#### Single inclusive spectrum



The single inclusive spectrum is given by :

$$\frac{dN_1}{d^3p} = \left. \frac{\delta F[z]}{\delta(p)} \right|_{z=1} = \int d^4x d^4y \,\, e^{\mathfrak{i} p \cdot (x-y)} \,\, \Box_x \Box_y \Big[ \textcolor{red}{A_+(x)A_-(y)} + \textcolor{red}{G_{+-}(x,y)} \Big]_{z=1}$$

Two types of terms :



#### **Leading Order**



- LO ≡ tree diagrams
  - > the second terms can be ignored
- In each blob, we must sum over all the tree diagrams, and over all the possible cuts:

$$\frac{dN_1}{d^3p}\bigg|_{LO} = \sum_{\text{trees}} \sum_{\text{cuts}} \iff \text{tree}$$

• Note : at this point, we set  $z(\mathbf{p}) = 1$ 

#### Expression in terms of classical fields



 When summing over the cuts in a tree diagram, we only get the following combinations of propagators:

$$G_{++}^{0}(p) - G_{+-}^{0}(p)$$

$$G_{-+}^{0}(p) - G_{--}^{0}(p)$$

#### Retarded propagator

$$G_{++}^0(p) - G_{+-}^0(p) = G_{-+}^0(p) - G_{--}^0(p) = \frac{G_R^0(p)}{G_R^0(p)} \quad \text{(retarded propagator)}$$

• For any tree diagram contributing to the 1-point functions  $A_{\pm}$ , the sum over the  $\pm$  indices at the vertices simply transforms all the propagators into retarded propagators



#### Sum of trees with retarded propagators :

$$\square \mathcal{A} + U'(\mathcal{A}) = \mathfrak{j} \qquad , \quad \lim_{x_0 \to -\infty} \mathcal{A}(x) = 0$$

Expansion in powers of J (for cubic interactions) :



Built with retarded propagators



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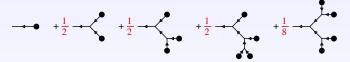
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- · Built with retarded propagators
- · Classical solutions resum the full series of tree diagrams

#### Inclusive particle spectrum at LO



The particle spectrum at LO is given by :

$$\left.\frac{dN_1}{d^3\vec{\boldsymbol{p}}}\right|_{LO} = \frac{1}{16\pi^3}\int_{x,y}\;e^{i\mathbf{p}\cdot(x-y)}\;\Box_x\Box_y\;\boldsymbol{\mathcal{A}}(x)\boldsymbol{\mathcal{A}}(y)$$

where  $\mathcal{A}(x)$  is the classical solution such that  $\lim_{x^0\to -\infty}\mathcal{A}(x)=0$ 

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where  $\mathcal{A}(x)$  is the classical solution such that  $\lim_{x^0\to -\infty}\mathcal{A}(x)=0$ 

• NOTE : if the source J is time independent, no particles can be produced at LO (because  $\mathcal{A}(x)$  has no time-like Fourier modes)

## Leading Order

Next-to-

#### Why the LO may be insufficient?



Naive perturbative expansion :

$$\frac{dN}{d^{3}\vec{p}} = \frac{1}{g^{2}} \left[ c_{0} + c_{1} g^{2} + c_{2} g^{4} + \cdots \right]$$

Note : so far, we have seen how to compute  $c_0$ 

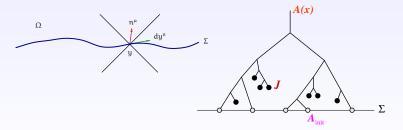
- The source is time dependent, and particle production is impossible at LO
- The description of the projectiles as external sources is valid for modes with large longitudinal momenta  $k^{\pm} > \Lambda$ . Loop corrections produce logs of this unphysical cutoff. The logs must be computed and resummed
- In QCD and other theories, there are instabilities that cause the coefficients  $c_n$  to grow indefinitely with time. These secular terms must be resummed

#### Cauchy problem for classical fields



#### Green's formula for classical solutions

$$\frac{\mathcal{A}(x)}{\mathcal{A}(x)} = i \int\limits_{y \in \Omega} G_{\text{R}}^{0}\left(x,y\right) \left[J(y) - V'(\underline{\mathcal{A}(y)})\right] + i \int\limits_{y \in \Sigma} G_{\text{R}}^{0}\left(x,y\right) \left(n \cdot \stackrel{\leftrightarrow}{\vartheta}_{y}\right) \underline{\mathcal{A}}_{\text{init}}(y)$$



#### Small perturbations of a classical field

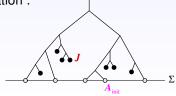


#### Linearized equation of motion around a classical background

$$\left[\Box_x + V''(\mathcal{A}(x))\right] \frac{\alpha(x)}{\alpha(x)} = 0 \qquad , \quad \frac{\alpha(x)}{\alpha(x)} = \frac{\alpha(x)}{\alpha(x)} \text{ on } \Sigma$$

#### **Formal solution**

· Diagrammatic interpretation :



A(x)

#### Small perturbations of a classical field

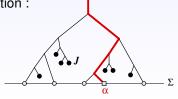


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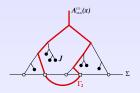


#### Translation operator for the initial field

$$\mathfrak{F}[\mathcal{A}_{\text{initial}} + \alpha] \equiv \exp\left[\int_{\vec{u} \in \Sigma} \left[\alpha \cdot \mathbb{T}\right]_{\vec{u}}\right] \, \mathfrak{F}[\mathcal{A}_{\text{initial}}]$$

• This formula means that  $\mathbb{T}_u$  is the generator of shifts of the initial value of the classical field  $\mathcal A$ 

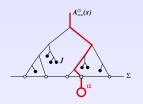




 A loop can be obtained by shifting the initial condition of A at two points

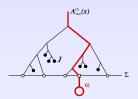
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- A loop can be obtained by shifting the initial condition of A at two points
- A term linear in  $\ensuremath{\mathbb{T}}$  is necessary if the loop is entirely below the initial surface



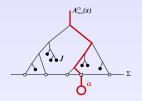


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#### Single field at NLO

$$\boldsymbol{A}_{\text{NLO}} = \left[\frac{1}{2}\int\limits_{\mathbf{u},\mathbf{v}} \boldsymbol{\Gamma}_{2}(\mathbf{u},\boldsymbol{v}) \, \mathbb{T}_{\mathbf{u}} \mathbb{T}_{\mathbf{v}} + \int\limits_{\mathbf{u}} \boldsymbol{\alpha}(\mathbf{u}) \, \mathbb{T}_{\mathbf{u}} \right] \, \boldsymbol{A}_{\text{LO}}$$





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#### Works also for any observable expressible in terms of the field at LO

#### But there is no free lunch...



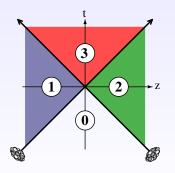
- For this formula to be true, the functions  $\Gamma_2$  and  $\alpha$  must be determined consistently :
  - $\Gamma_2$  = dressed propagator with endpoints on  $\Sigma$
  - $\alpha$  = 1-point function at 1-loop with endpoint on  $\Sigma$
- This is almost as hard as doing the NLO calculation!
   (law of conservation of difficulty...)
- In some cases, there is an advantage :
  - If  $\Sigma$  is at  $t = -\infty$ , then  $\Gamma_2$  is trivial and  $\alpha = 0$
  - When the classical field is simple below  $\Sigma$  and complicated above, then  $\Gamma_2$  and  $\alpha$  are simpler to calculate than the NLO observable above  $\Sigma$

#### **Example: nucleus-nucleus collisions**



Sources located on the light-cone:

$$J^{\mu} = \delta^{\mu+} \underbrace{\rho_1(x^-, x_{\perp})}_{\sim \delta(x^-)} + \delta^{\mu-} \underbrace{\rho_2(x^+, x_{\perp})}_{\sim \delta(x^+)}$$



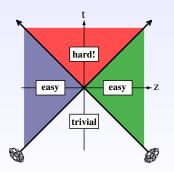
- Region 0 :  $A^{\mu} = 0$
- Regions 1,2 :  $\mathcal{A}^{\mu}$  depends only on  $\rho_1$  or  $\rho_2$  (known analytically)
- Region 3 : A<sup>μ</sup> = radiated field after the collision, only known numerically

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 $\Longrightarrow$  choose  $\Sigma$  just above the forward light-cone

## and Resummation

Instabilities

#### Mode decomposition of $\Gamma_2$



$$\Gamma_2(x,y) = \int_{\text{modes } k} \alpha_k(x) \alpha_k^*(y)$$

with

$$\left[\Box_x + V''(\mathcal{A}(x))\right] \frac{\alpha_k(x)}{\alpha_k(x)} = 0 \qquad , \quad \lim_{t \to -\infty} \frac{\alpha_k(x)}{\alpha_k(x)} = e^{\mathrm{i} k \cdot x}$$

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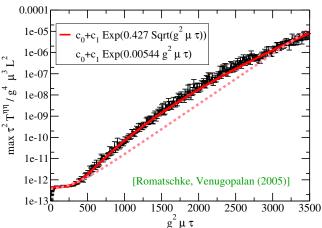
with

$$\left[\Box_x + V''({\color{black}\mathcal{A}}(x))\right] \frac{\alpha_k(x)}{\alpha_k(x)} = 0 \qquad , \quad \lim_{t \to -\infty} \frac{\alpha_k(x)}{\alpha_k(x)} = e^{i\,k\cdot x}$$

- The equation of motion of the of the mode functions  $\alpha_k$  is linear
- Some of the modes can be unstable
- What happens to the NLO observables?

#### Yang-Mills theory: Weibel instabilities for small perturbations

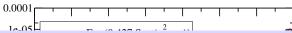
[Mrowczynski (1988), Romatschke, Strickland (2003), Arnold, Lenaghan, Moore (2003), Rebhan, Romatschke, Strickland (2005), Arnold, Lenaghan, Moore, Yaffe (2005), Romatschke, Rebhan (2006), Bodeker, Rummukainen (2007),...



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- For some k's, the field fluctuations α<sub>k</sub> diverge like  $\exp \sqrt{\mu \tau}$  when  $\tau \to +\infty$
- Some components of  $T^{\mu\nu}$  have secular divergences when evaluated at fixed loop order
- When  $\alpha_k \sim A \sim q^{-1}$ , the power counting breaks down and additional contributions must be resummed:

$$g e^{\sqrt{\mu \tau}} \sim 1$$
 at  $\tau_{max} \sim \mu^{-1} \log^2(g^{-1})$ 

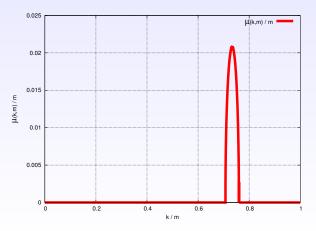
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#### $\phi^4$ scalar field theory



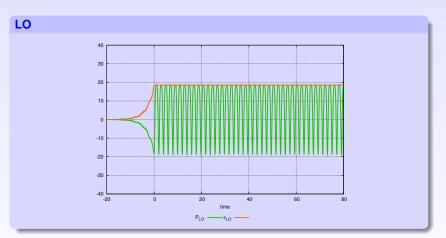
· Lyapunov exponent for the mode k:

$$\mu_{\mathbf{k}} \equiv \frac{1}{T} \ln \left( \frac{\alpha_{\mathbf{k}}(t+T)}{\alpha_{\mathbf{k}}(t)} \right)$$



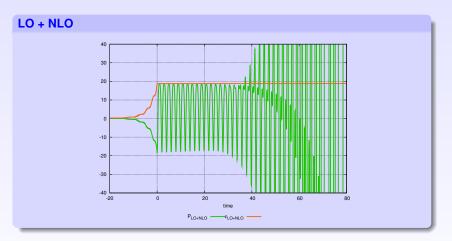
#### $\varphi^4$ scalar field theory : pathologies in fixed order calculations





#### $\varphi^4$ scalar field theory : pathologies in fixed order calculations





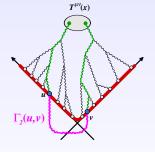
- Small correction to the energy density (protected by energy conservation)
- Secular divergence in the pressure

#### Improved power counting and resummation





$$\mathbb{T} \sim e^{\sqrt{\mu\tau}}$$



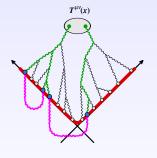
• 1 loop :  $(ge^{\sqrt{\mu\tau}})^2$ 

#### Improved power counting and resummation



$$\mathsf{Loop} \sim \mathsf{g}^2 \qquad , \qquad \mathbb{T} \sim \mathsf{e}^{\sqrt{\mu\tau}}$$

$$\mathbb{T} \sim e^{\sqrt{\mu\tau}}$$



 1 loop :  $(ge^{\sqrt{\mu\tau}})^2$ 

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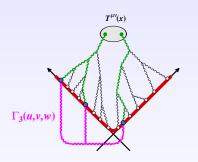
• 2 disconnected loops :  $(ge^{\sqrt{\mu\tau}})^4$ 

#### Improved power counting and resummation



Loop 
$$\sim g^2$$
 ,  $\mathbb{T} \sim e^{\sqrt{\mu \tau}}$ 

$$\mathbb{T} \sim e^{\sqrt{\mu \tau}}$$



- 1 loop :  $(qe^{\sqrt{\mu\tau}})^2$
- 2 disconnected loops  $(ge^{\sqrt{\mu\tau}})^4$
- 2 entangled loops :  $q(qe^{\sqrt{\mu\tau}})^3 > \text{subleading}$

#### **Leading terms**

- All disconnected loops to all orders
  - > exponentiation of the 1-loop result

#### Interlude...



$$e^{\frac{\alpha}{2}\delta_x^2} f(x) = \int_{-\infty}^{+\infty} dz \, \frac{e^{-z^2/2\alpha}}{\sqrt{2\pi\alpha}} f(x+z)$$

#### Resummation of the leading secular terms



$$\begin{array}{lcl} T_{\text{resummed}}^{\mu\nu} & = & \exp\left[\frac{1}{2}\int\limits_{\mathbf{u},\nu} \Gamma_{2}(\mathbf{u},\nu) \mathbb{T}_{\mathbf{u}} \mathbb{T}_{\nu}\right] T_{\text{LO}}^{\mu\nu} [\mathcal{A}_{\text{init}}] \\ \\ & = & \underbrace{T_{\text{LO}}^{\mu\nu} + T_{\text{NLO}}^{\mu\nu}}_{\text{in full}} + \underbrace{T_{\text{NNLO}}^{\mu\nu} + \cdots}_{\text{partially}} \end{array}$$

 The exponentiation of the 1-loop result collects all the terms with the worst time behavior

#### Resummation of the leading secular terms



- The exponentiation of the 1-loop result collects all the terms with the worst time behavior
- Equivalent to Gaussian fluctuations of the initial field + classical time evolution

#### Note: Classical field + Fluctuations = Coherent state



• This Gaussian distribution of initial fields is the Wigner distribution of a **coherent state**  $|\mathcal{A}\rangle$ 

Coherent states are the "most classical quantum states"

Their Wigner distribution has the minimal support permitted by the uncertainty principle ( $\mathfrak{O}(\hbar)$  for each mode)

•  $|\mathcal{A}\rangle$  is not an eigenstate of the full Hamiltonian  $\rhd$  decoherence via interactions

# **Example:**Schwinger mechanism

[FG, N. Tanji (2013)]

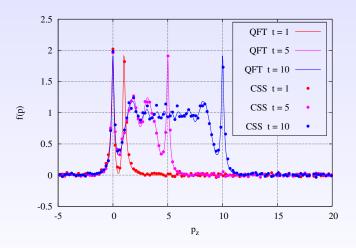
#### Scalar QED model



$$\begin{split} \mathcal{L} \equiv &\underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu}}_{\text{photons}} + \underbrace{(D_{\mu} \varphi) (D^{\mu} \varphi)^* - m^2 \varphi^* \varphi - \frac{\lambda}{4} (\varphi \varphi^*)^2}_{\text{charged scalars}} + J^{\mu}_{\text{ext}} A_{\mu} \\ F^{\mu\nu} = &\partial^{\mu} A^{\nu} - \partial^{\nu} A^{\mu} \quad , \quad D^{\mu} \equiv \partial^{\mu} - i e A^{\mu} \quad , \end{split}$$

- Two coupling constants :
   e (electrical charge) and λ (self-coupling)
- When the external field is static, no perturbative production
- Non perturbative production  $\sim \exp(-\pi m^2/eE_{ext})$

• Comparison with the 1-loop QFT result (for  $\lambda = 0$ ) :



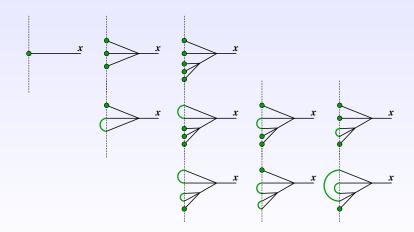
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- QFT = 1-loop quantum field theory
- CSS = classical statistical simulation

#### **Mass renormalization**

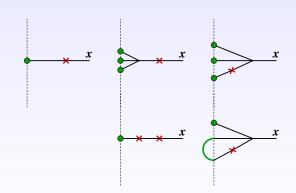


• When  $\lambda \neq 0$ , tadpoles give a quadratic cutoff dependence  $\sim \Lambda^2$ 

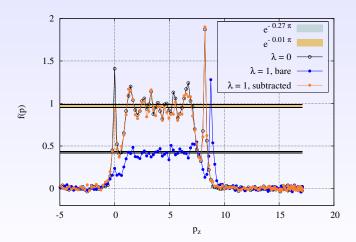


 This can be compensated by a counterterm in the equation of motion:

$$\left(D_0^2 - \sum_i D_i D_i + m_0^2 + \delta m^2\right) \phi + \frac{\lambda}{2} (\phi^* \phi) \phi = 0,$$



• Comparison of bare and mass-renormalized results, for  $\lambda=1$  (at very short time, so that we know that the scalar self-interactions should not have affected the system yet):



### Summary

#### Summary



- In Quantum Field Theories coupled to strong sources :
  - The LO is expressible in terms of classical fields
  - The NLO can be related to the LO by an operator acting on the initial fields
- When the classical fields are unstable (i.e. in theories where the classical Hamiltonian has chaotic dynamics), the loop expansion is ill behaved
- The terms that have the fastest growth can be summed to all loop orders
  - The result of the resummation can be obtained by averaging the LO result over a Gaussian ensemble of initial conditions
- Accuracy : LO + NLO + leading secular terms