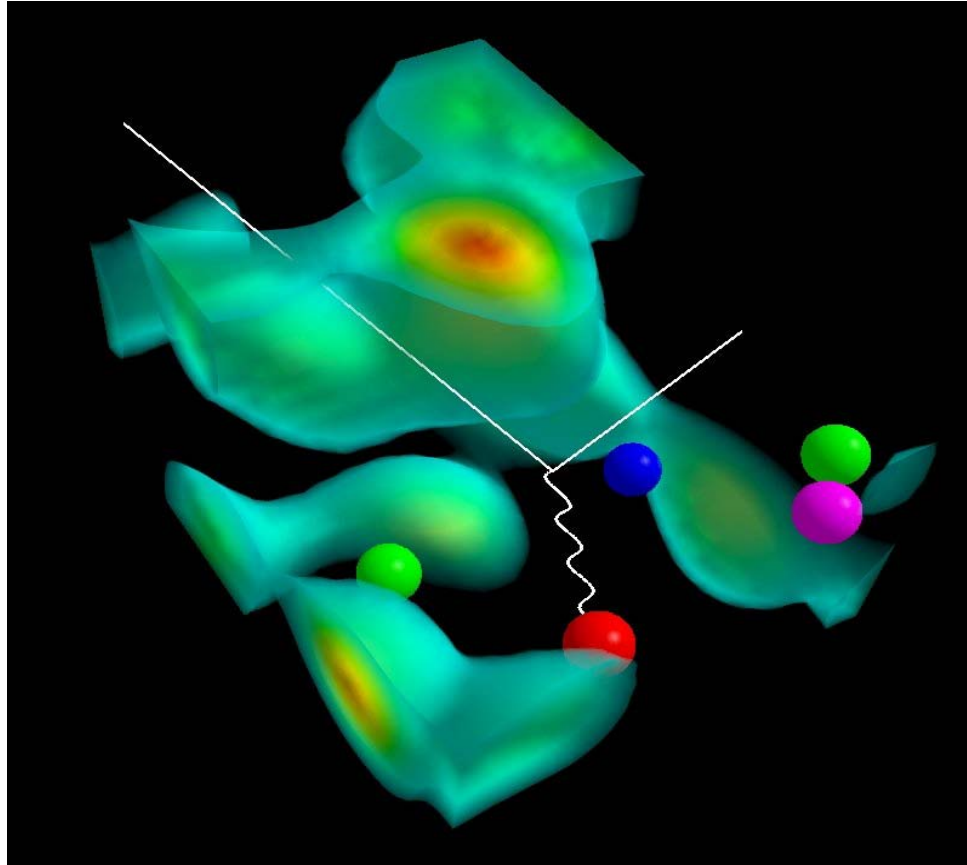


Welcome to Adelaide



Australian Government
Australian Research Council



Anthony W. Thomas

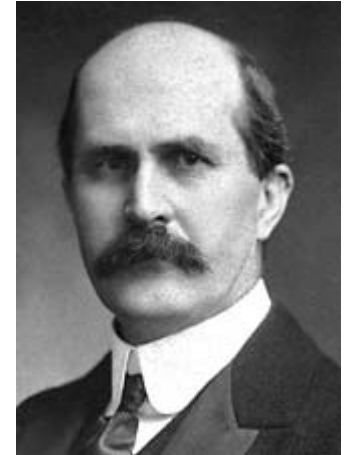
**ANPhA Symposium and Board Meeting
University of Adelaide : August 3rd 2012**



The Very Beginning

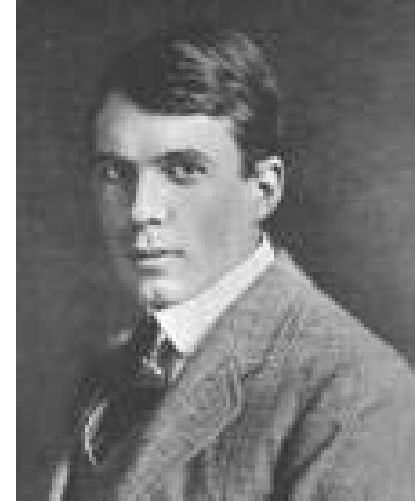
In **Adelaide** William Henry Bragg was appointed as Elder Professor of Physics in 1886.

Research aimed at understanding the nature of X-rays – he concluded they were particles



Data eventually forced a re-evaluation

**Working with William Lawrence Bragg,
WH was soon forced to re-examine
this and concluded that sometimes X-rays
act like waves and sometimes particles**



Only later understood in terms of Quantum Mechanics

A Very Brief History of Subatomic Physics

- **1995: Won national competition to host a National Institute of Theoretical Physics (NITP)**
- **1997: Awarded ARC Special Research Centre for the Subatomic Structure of Matter (CSSM)**
- **2009: CSSM re-invented as C²SSM (Adelaide University's Research Centre for Complex Systems and the Structure of Matter)**
- **2011: Node of the ARC Centre of Excellence for Particle Physics at the Terascale (CoEPP)**
- **Currently more than 40 staff and post-graduate students including experimental participation in ATLAS**

CSSM

- Centre for the Subatomic Structure of Matter formed as an ARC Special Research Centre in 1997
- To understand the quark structure of matter from protons and neutrons to atomic nuclei and neutron stars (pulsars)
- This has made Adelaide a major international player in subatomic physics – leading scientists from around the world come here to work with our staff and students



Australian Government

Australian Research Council

A few achievements since 1997

- **Students:**
 - 78 Graduate students 1997+
 - 54 Honours students
 - 46 undergraduates involved in research
- **Workshops:**
 - 35 Workshops and Conferences 1997+
 - Including QNP 2000; Symmetries in Subatomic Physics; Light-cone 2005; Pacific Spin 2011; Lattice 2012 and ICHEP (in Melbourne) 2012
- **Approx. 370 journal articles with more than 11,000 citations**

Just a few scientific highlights

Quark and Gluon Propagators

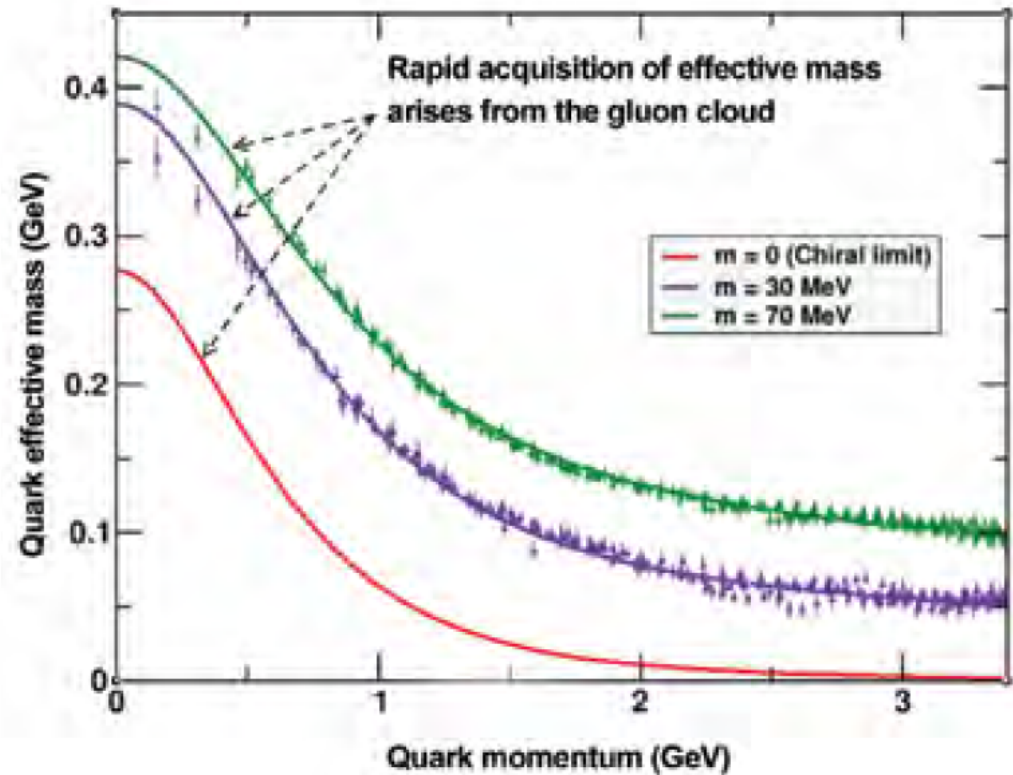
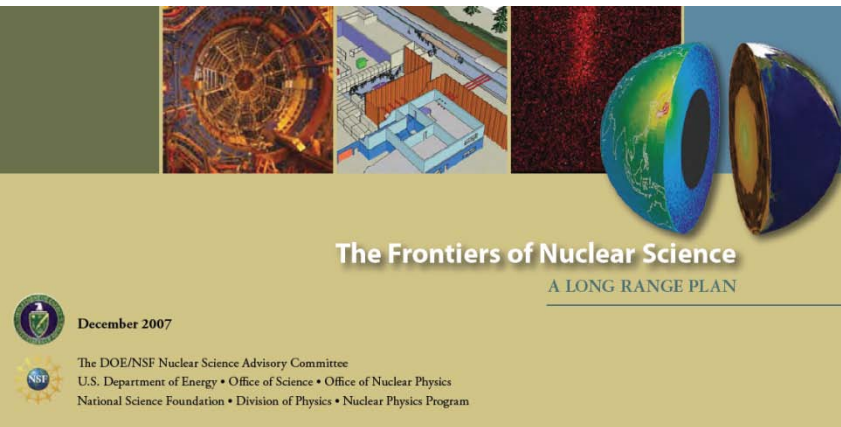
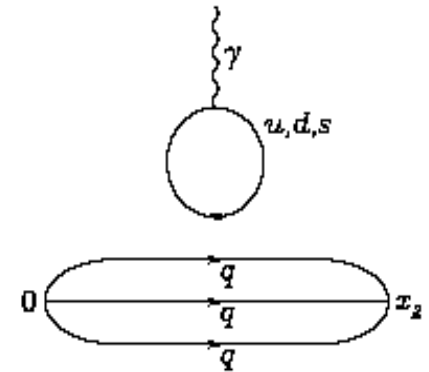
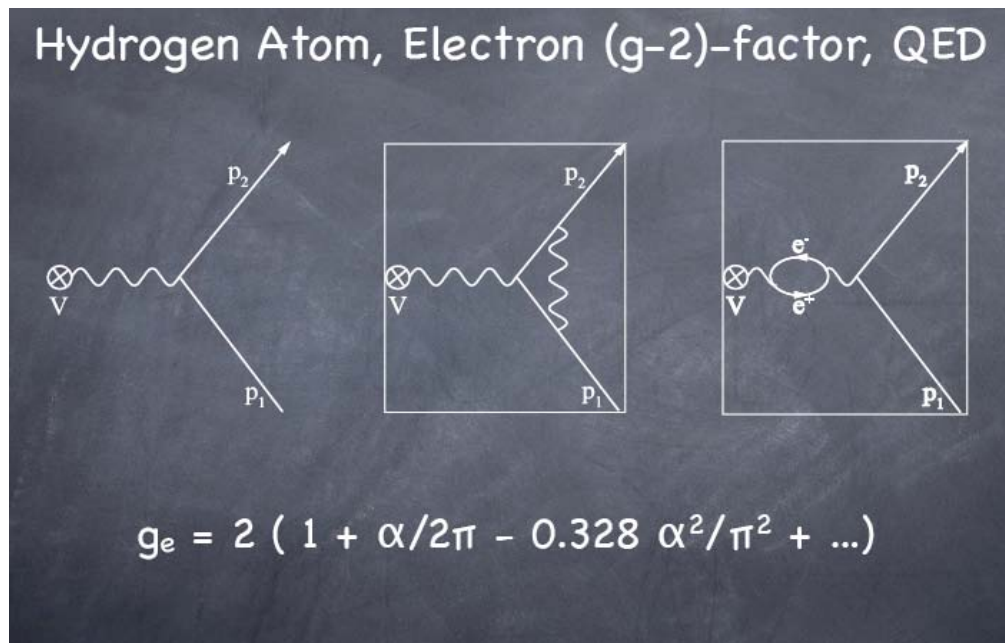


Figure 2.1: Mass from nothing. In QCD a quark's effective mass depends on its momentum. The function describing this can be calculated and is depicted here. Numerical simulations of lattice QCD (data, at two different bare masses) have confirmed model predictions (solid curves) that the vast bulk of the constituent-mass of a light quark comes from a cloud of gluons, that are dragged along by the quark as it propagates. In this way, a quark that appears to be absolutely massless at high energies ($m = 0$, red curve) acquires a large constituent mass at low energies.

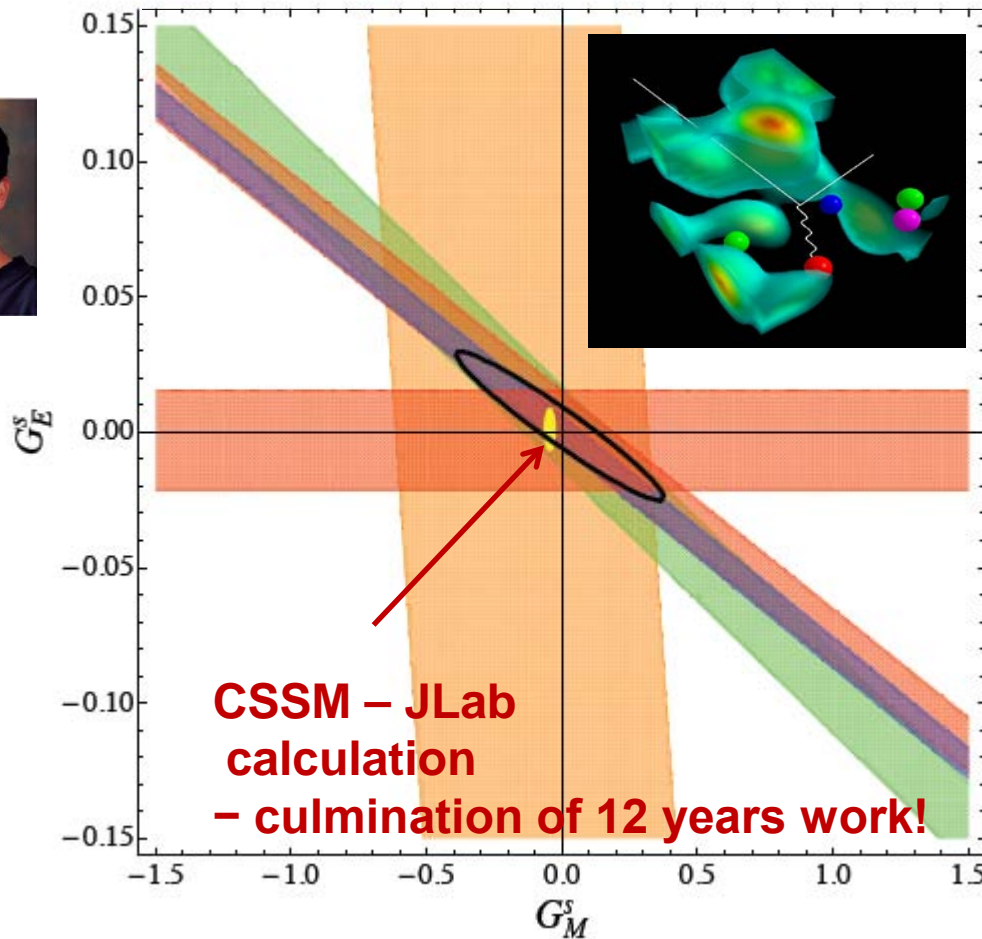
Fundamental Test of Non-Perturbative QCD

- Strangeness contribution is a vacuum polarization effect, *analogous to Lamb shift in QED*



- It is a fundamental test of QCD at long distances where the force is really *strong*

A unique case : theory 10 times more accurate than data



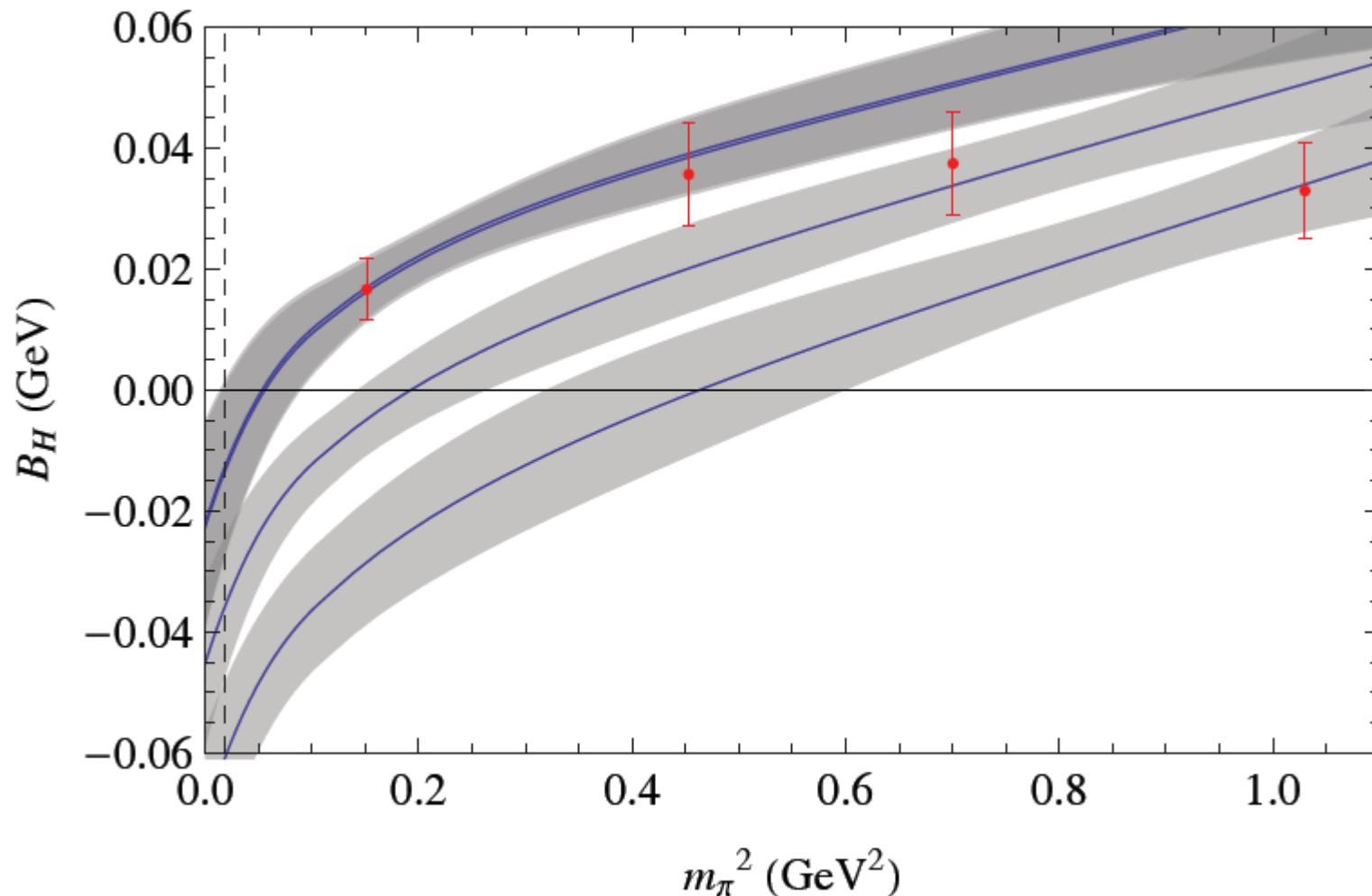
Experimental program
took three major
laboratories 20 years!



Thomas Jefferson National
Accelerator Facility (JLab)

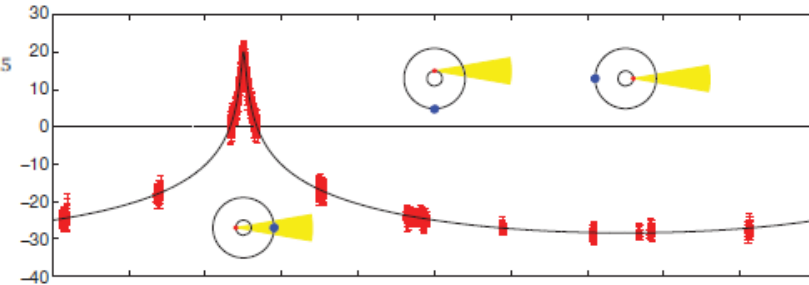
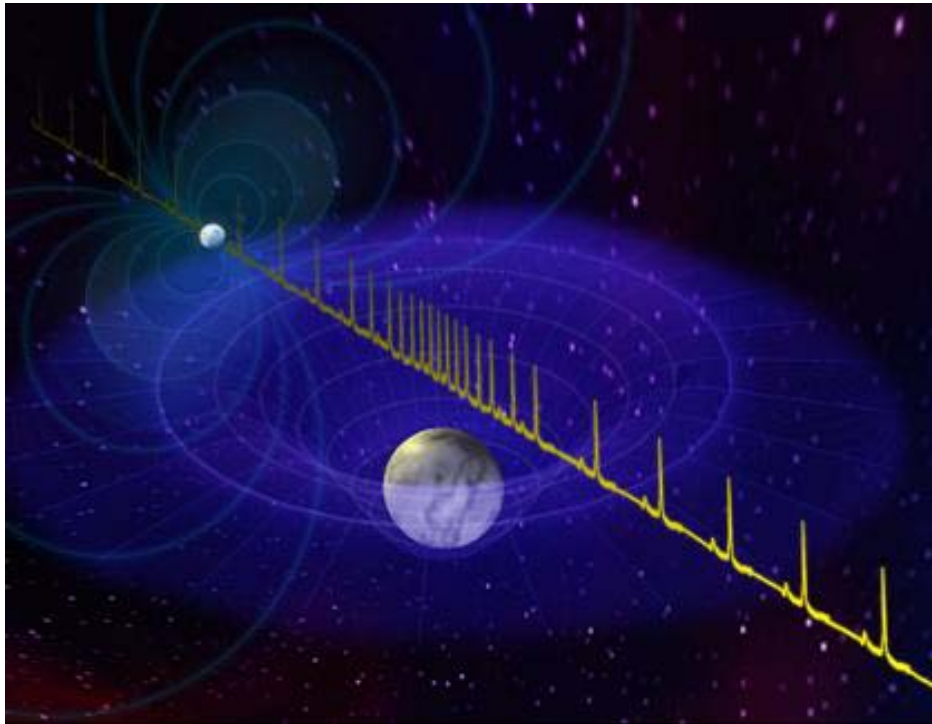
Last Strangeness Twist

Using data from NPLQCD & HAL:
find H-dibaryon almost bound – search at J-PARC



A two-solar-mass neutron star measured using Shapiro delay

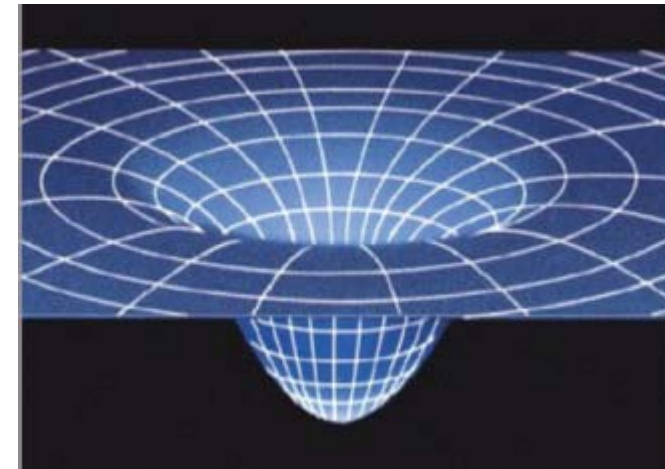
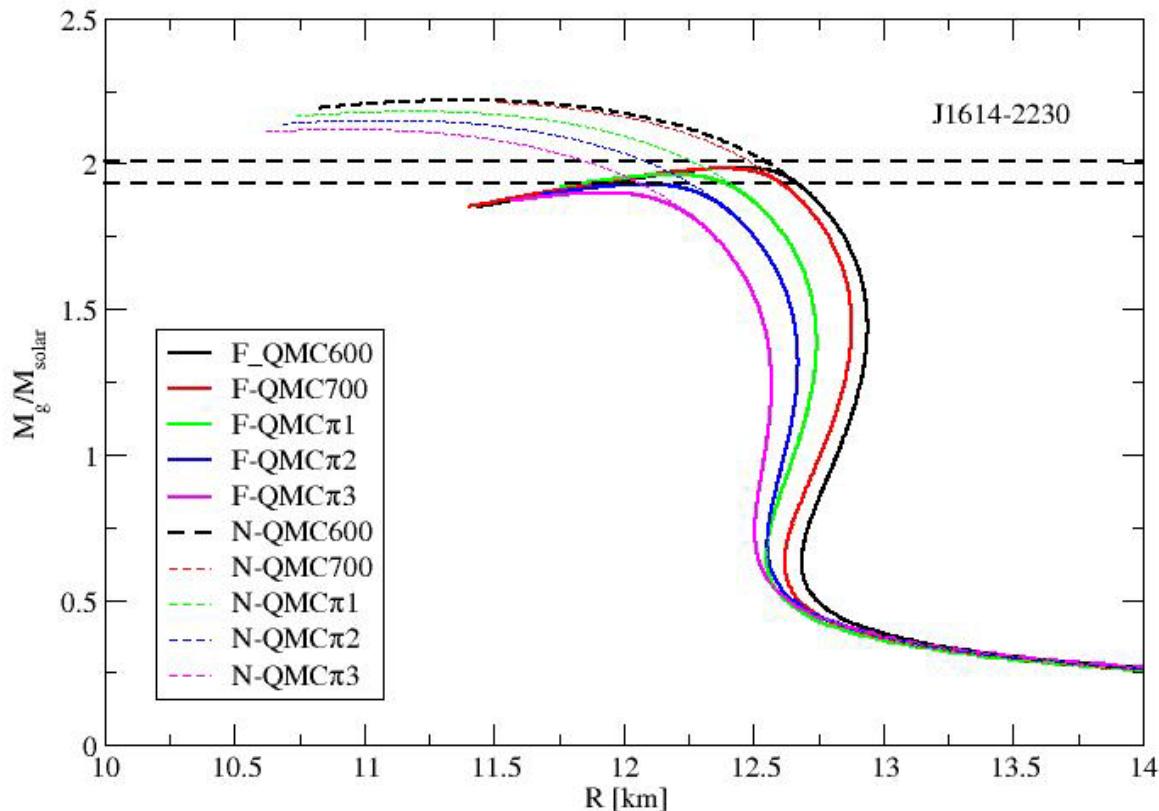
P. B. Demorest¹, T. Pennucci², S. M. Ransom¹, M. S. E. Roberts³ & J. W. T. Hessels^{4,5}



Report a very accurate pulsar mass much larger than seen before : 1.97 ± 0.04 solar mass

Claim it rules out hyperons (particles with strange quarks)

Just 3 years before*



We conclude that the Demorest et al. result, if confirmed, is very significant for neutron star physics and does indeed rule out all EoS which predict a mass-radius curve that does not intersect the J1614-2230 mass line. However, it does not provide any constraint on the possible 'exotic' composition of the high-density neutron star matter.

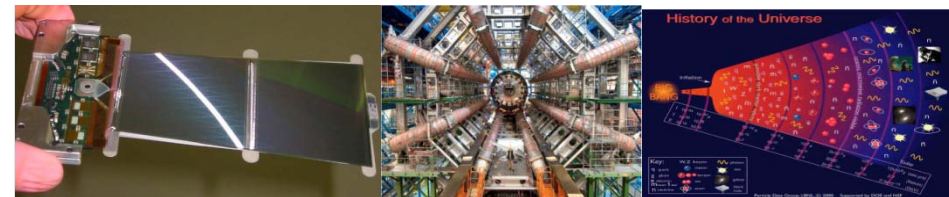
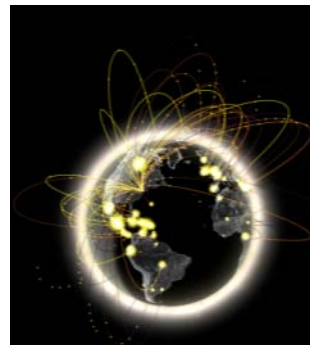
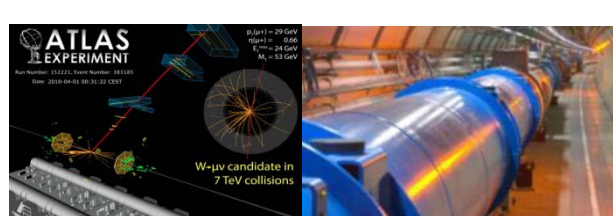
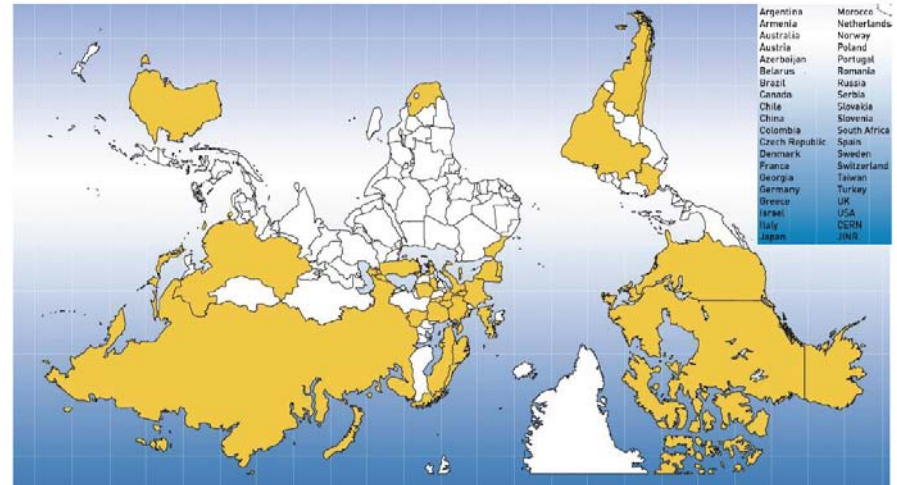
- Guichon et al., Nucl. Phys. A814 (2008) 66
- result of an on-going collaboration between CSSM & CEA France with Jirina Stone (Oxford)

ARC Centre of Excellence for Particle Physics at the Tera-Scale

The University of Adelaide
The University of Melbourne
The University of Sydney
Monash University



The ATLAS Collaboration



Physics at the LHC corresponds to conditions around here

Inflation

Particle Physics Regime

gim

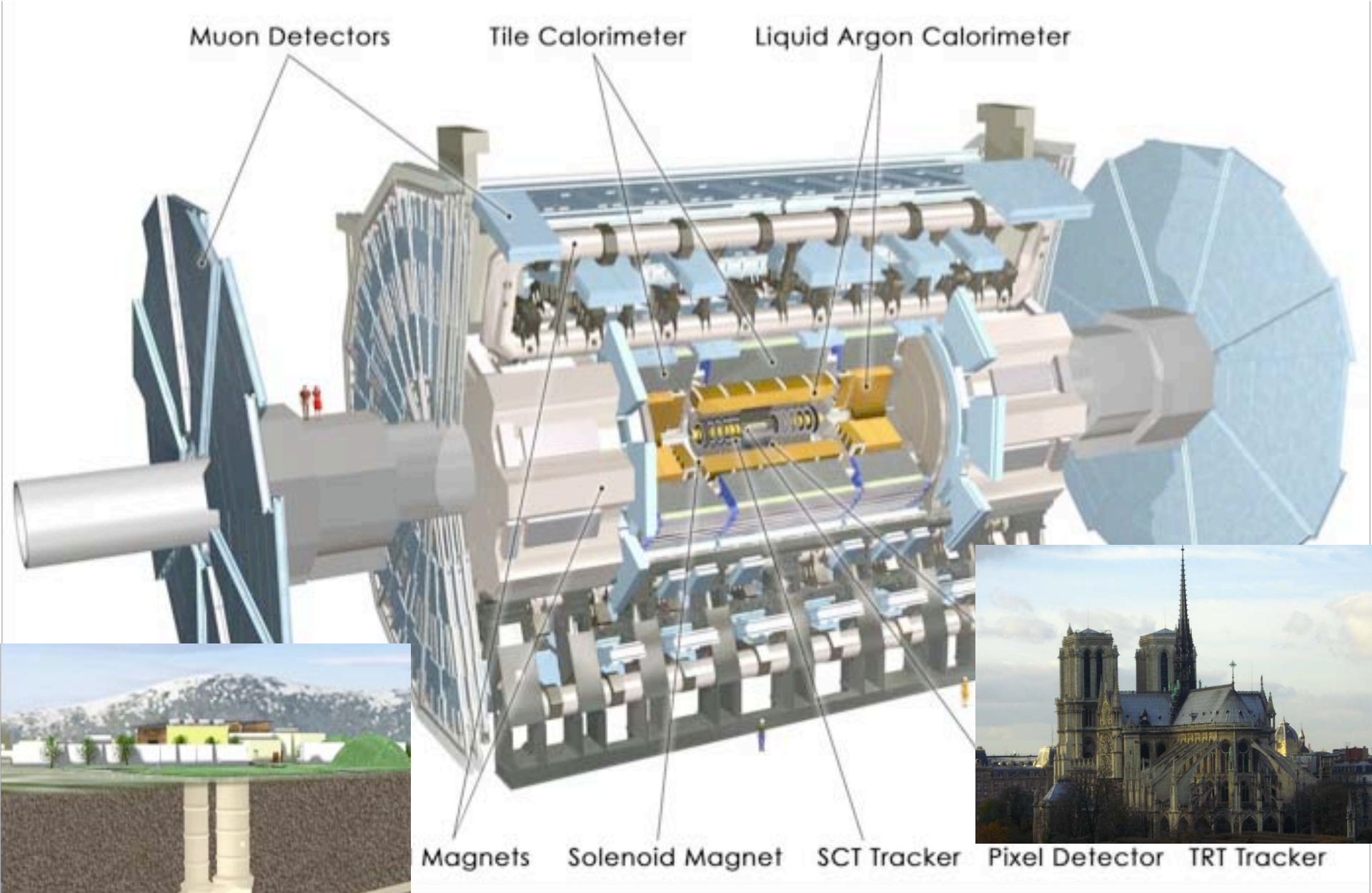
reflects

gim
u
on Visible

Today

12×10^9 y	(sec, yrs)
2.7	(Kelvin)
3×10^{-13}	(GeV)

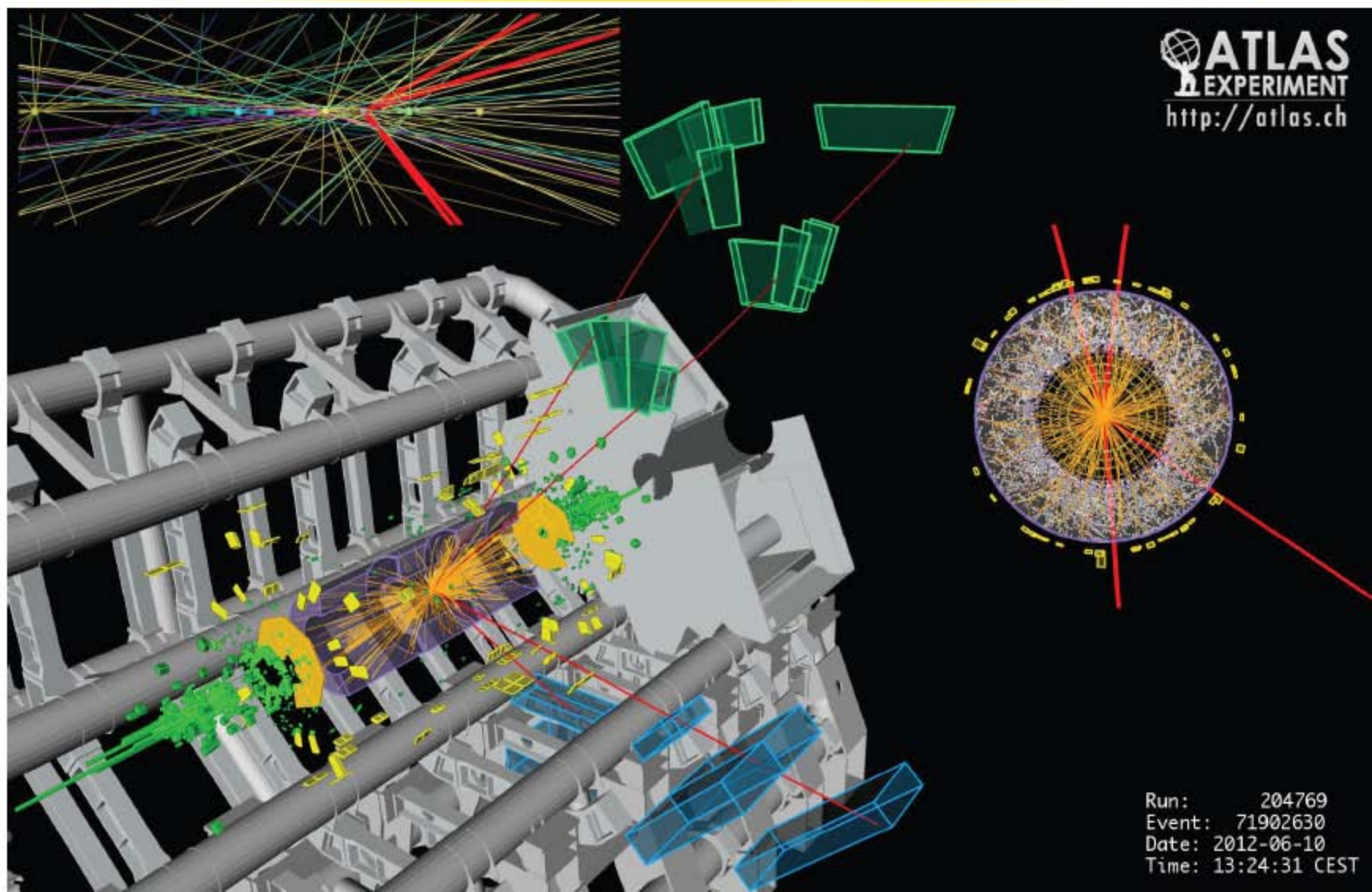
Key:	W, Z bosons		photon
q quark	 meson		star
g gluon	 baryon		galaxy
e electron	 ion		black hole
m muon	 atom		
t tau			
n neutrino			

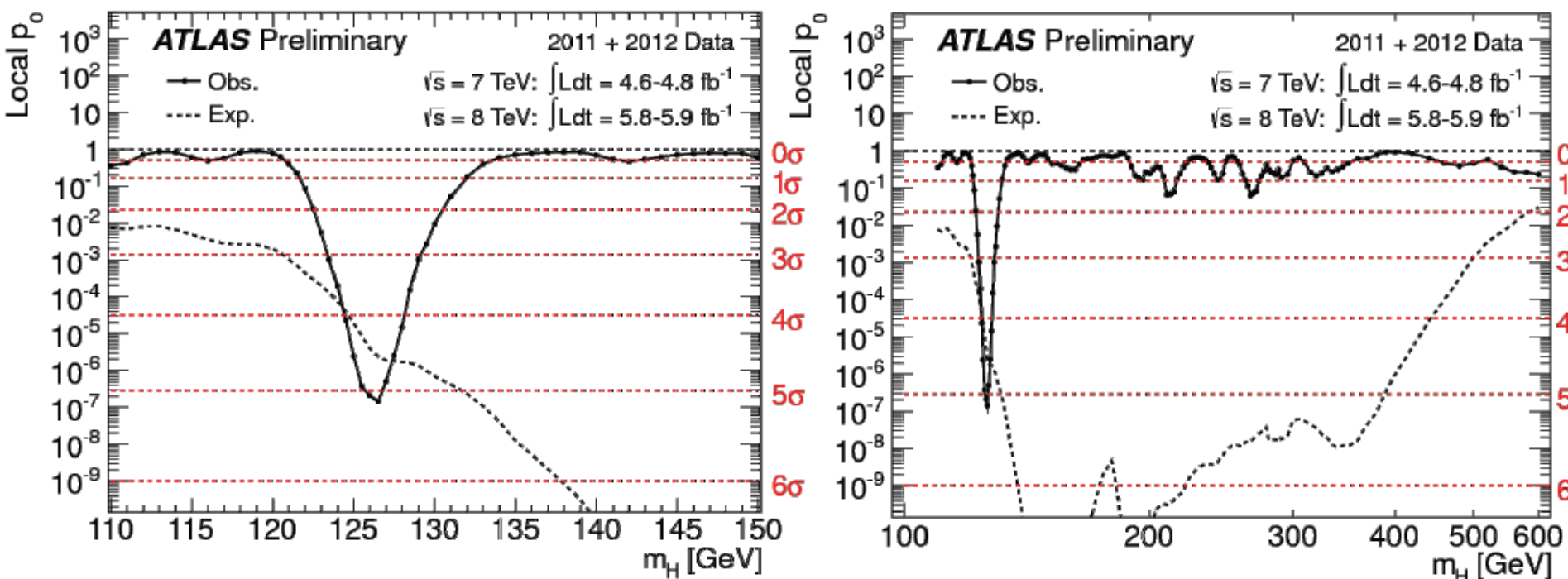


ARC CoE for Particle Physics at the Tera-Scale



Event Display: 4μ





- Smallest p_0 value corresponds to local significance of 5.1σ at $m_H = 126.5 \text{ GeV}$
 - Expected value if SM Higgs of mass 126.5 GeV exists is 4.6σ
- Significance mildly sensitive to energy scale and resolution systematics (ESS)
 - Important for photons and electrons, negligible for muons
 - ESS estimated to reduce local significance to 5.0σ

Summary

This is a wonderfully challenging and exciting time

The likely discovery of the Higgs boson constitutes the most exciting development in particle physics in almost 30 years

We know enough to realize just how much there must still be to learn

Through the CoEPP and CSSM Adelaide is ideally placed to play a major role in the coming decade and beyond



Event display: $2e2\mu$

